

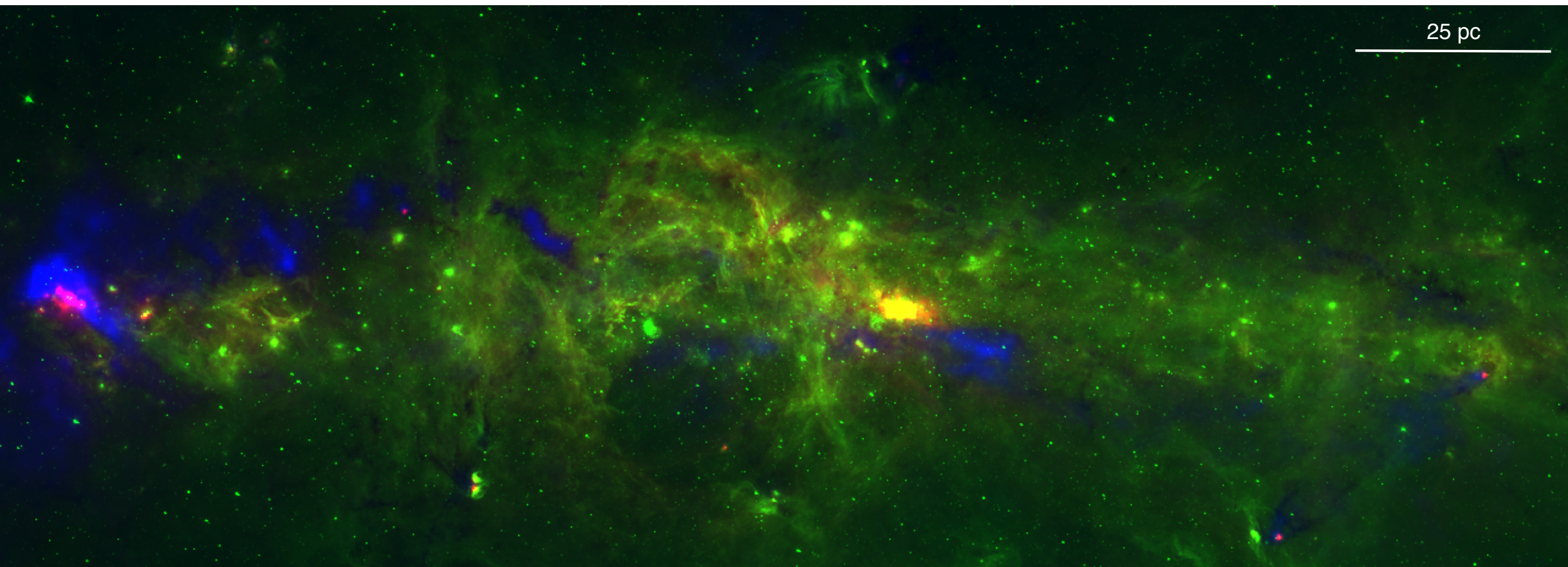
High-mass star formation in the Central Molecular Zone

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NAOJ / ALMA Fellow

Steve Longmore, Qizhou Zhang, Cara Battersby, Adam Ginsburg, Jonny Henshaw, Ashley Barnes, Diederik Kruijssen, Jens Kauffmann, Xing Lu, Betsy Mills, Thushara Pillai & many more!

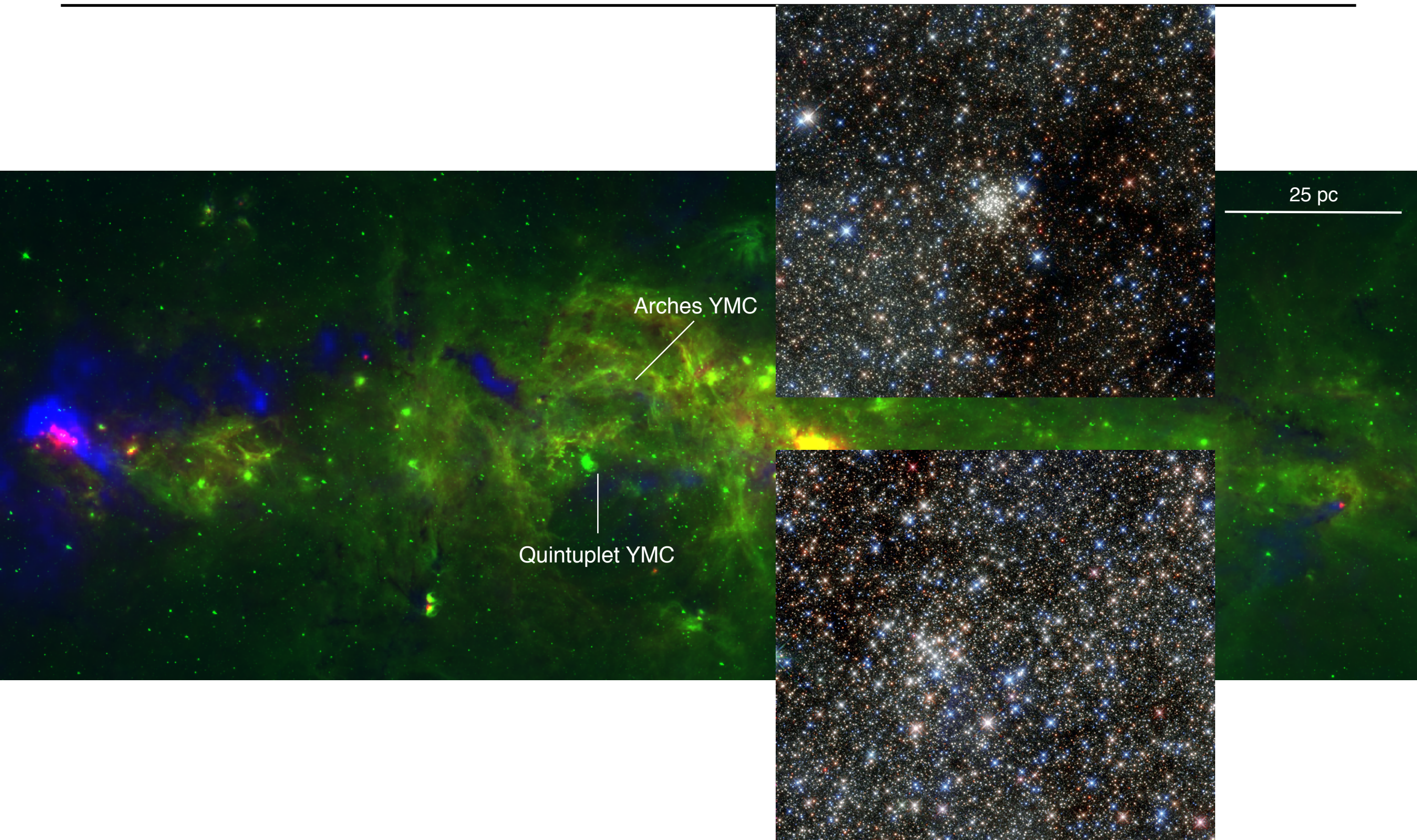
Central Molecular Zone (CMZ)

Green — 8 μm (GLIMPSE)
Red — 70 μm (Herschel)
Blue — dust column (HiGAL)



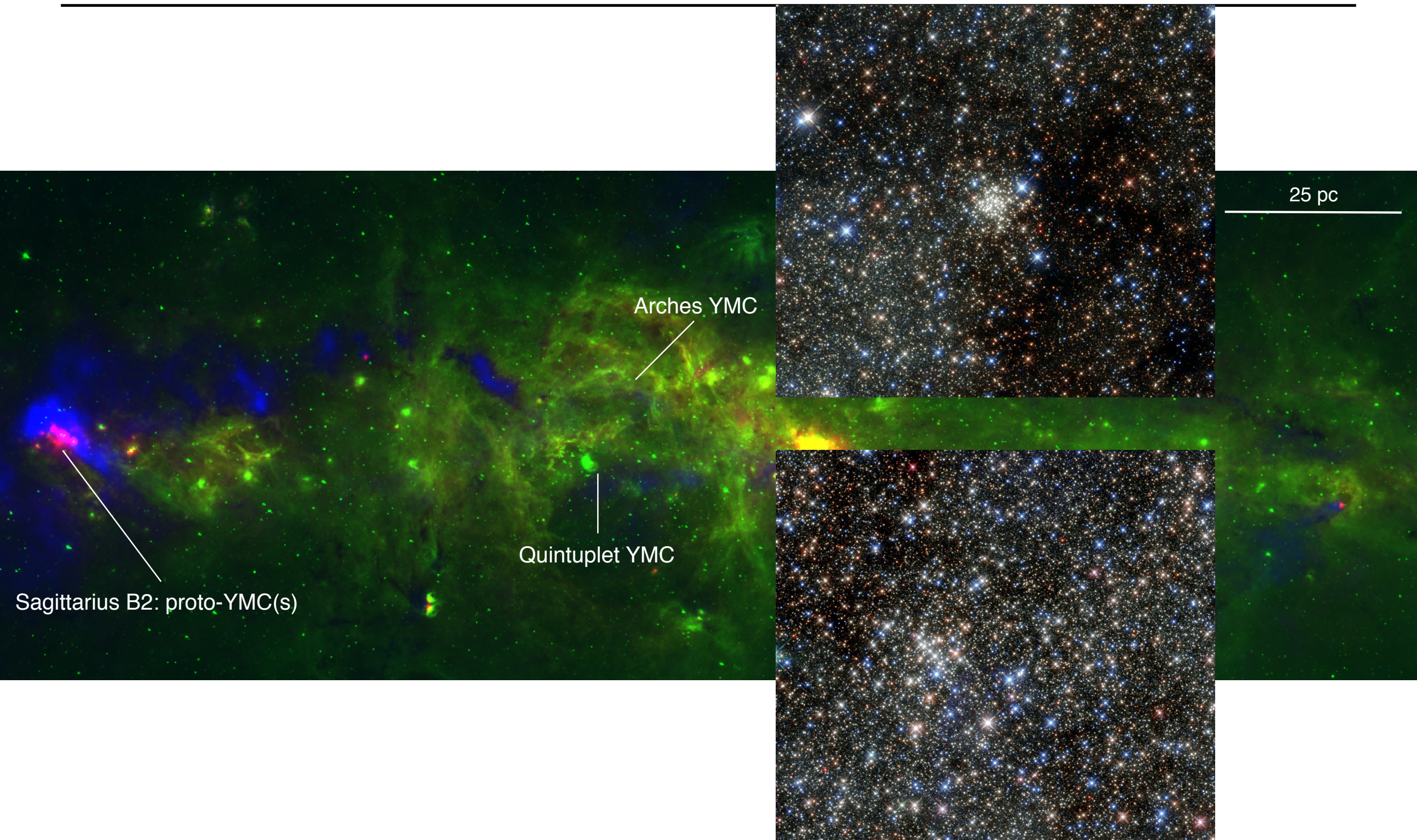
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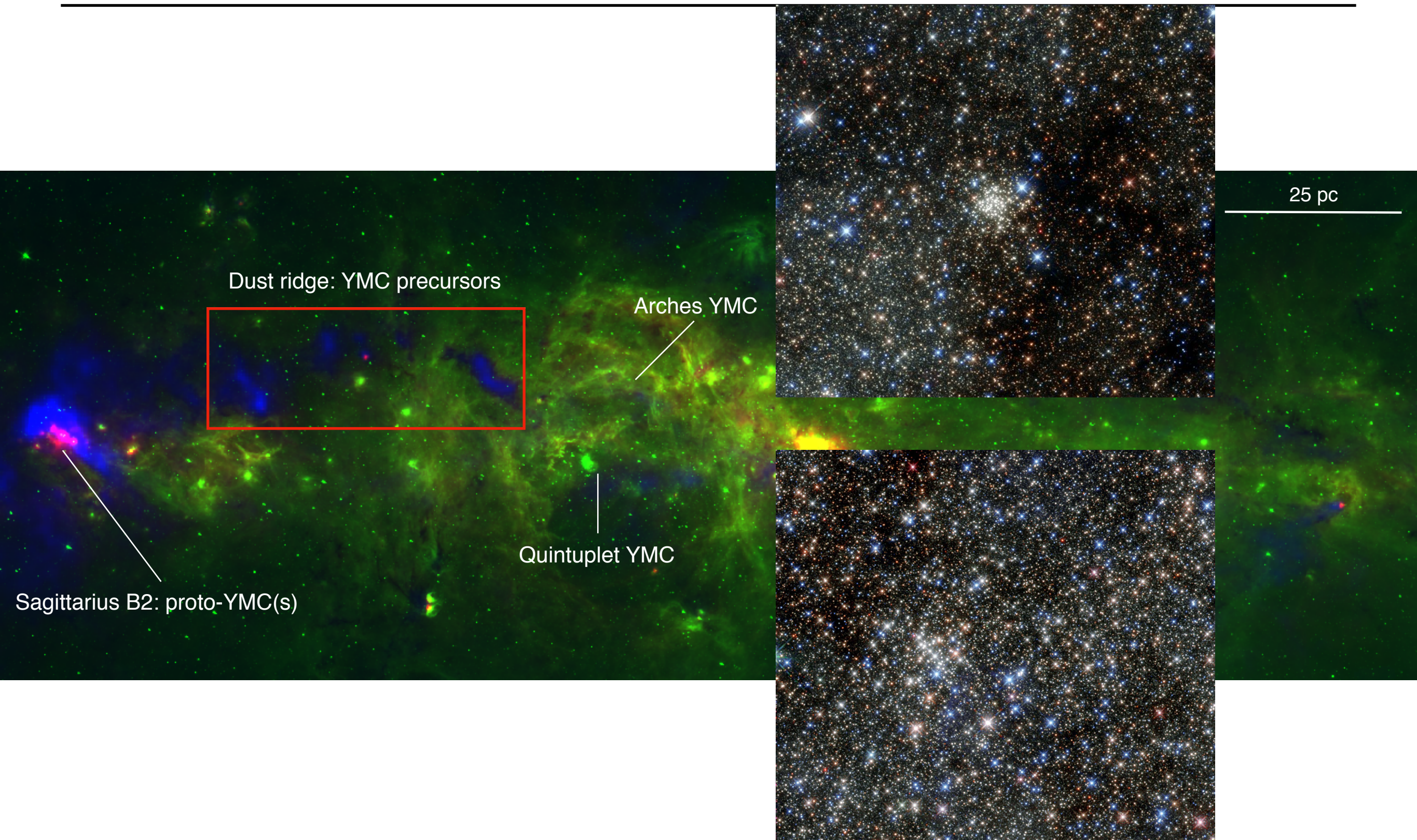
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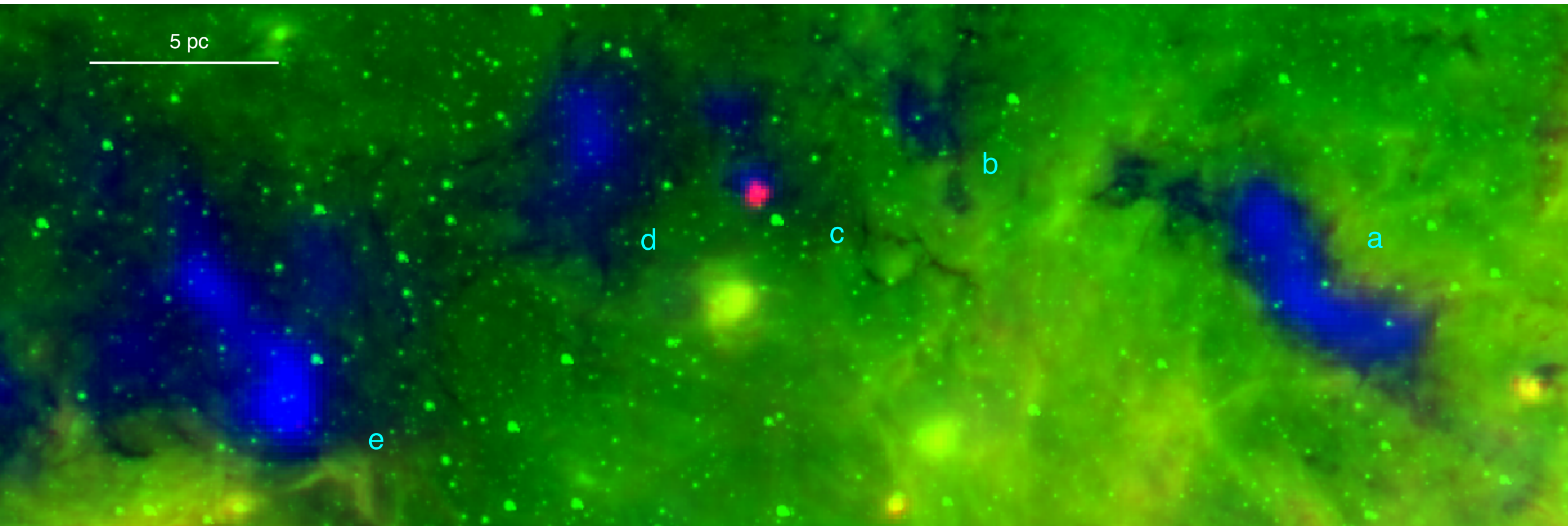
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The ‘Dust Ridge’

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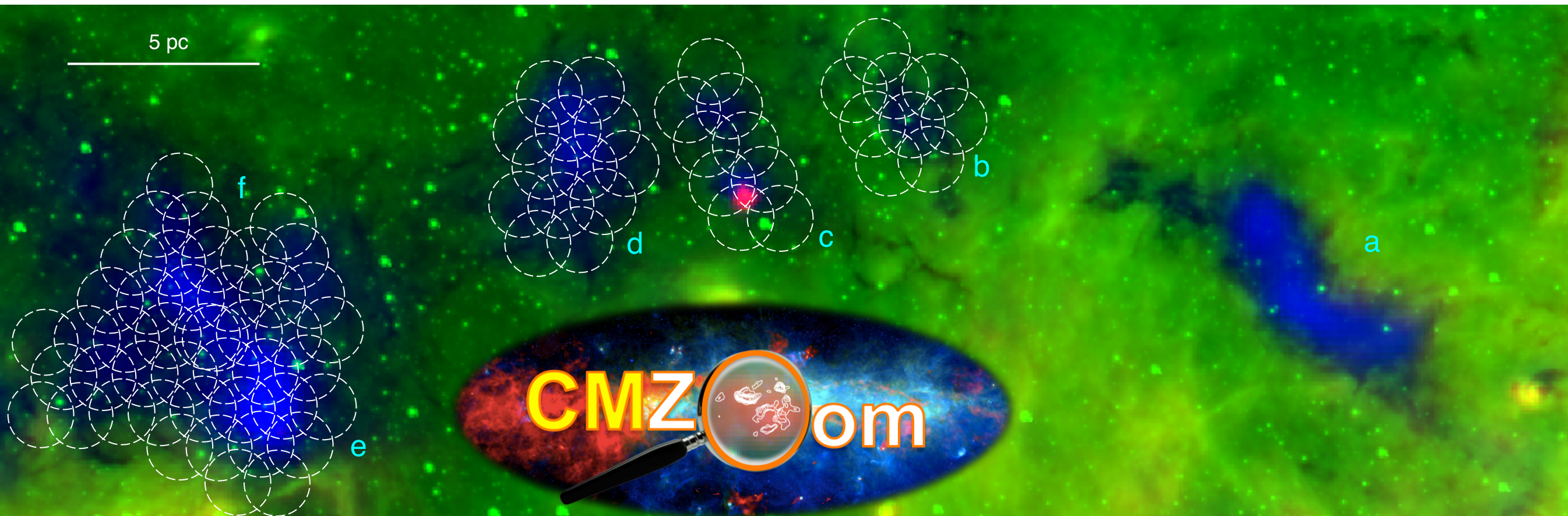
- 6 molecular clouds spanning ~ 0.25 deg (~ 40 pc)
 - $M \sim 10^{4-5} M_{\odot}$
 - $R \sim 2 - 3$ pc
 - No widespread SF
- ➔ Amongst the most massive and compact **quiescent** clouds known to exist in the Galaxy!

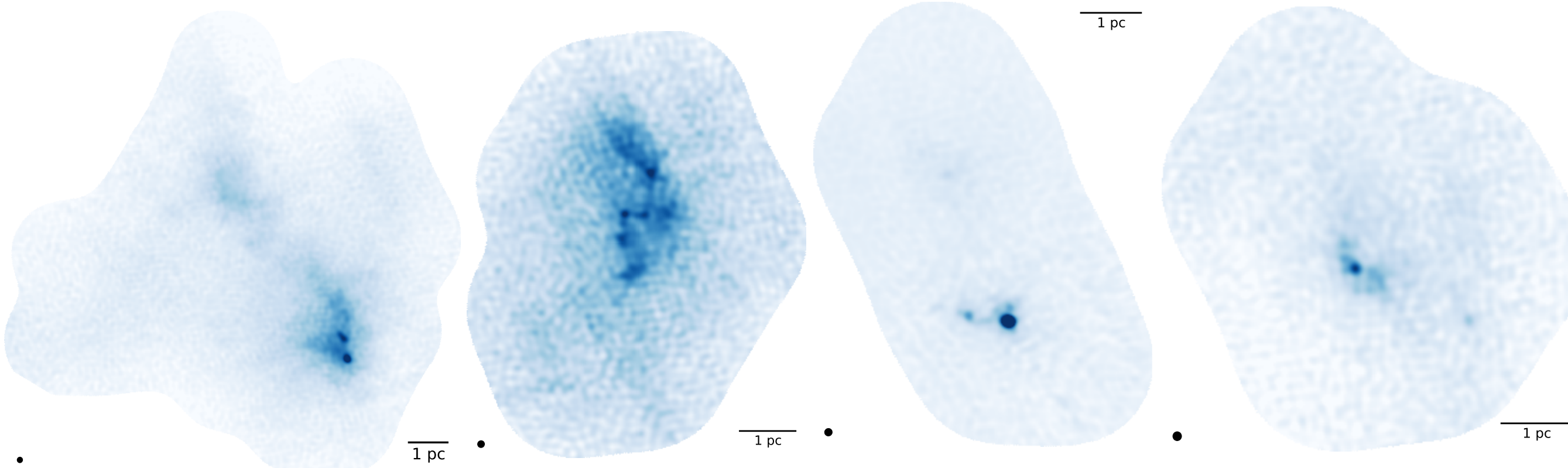


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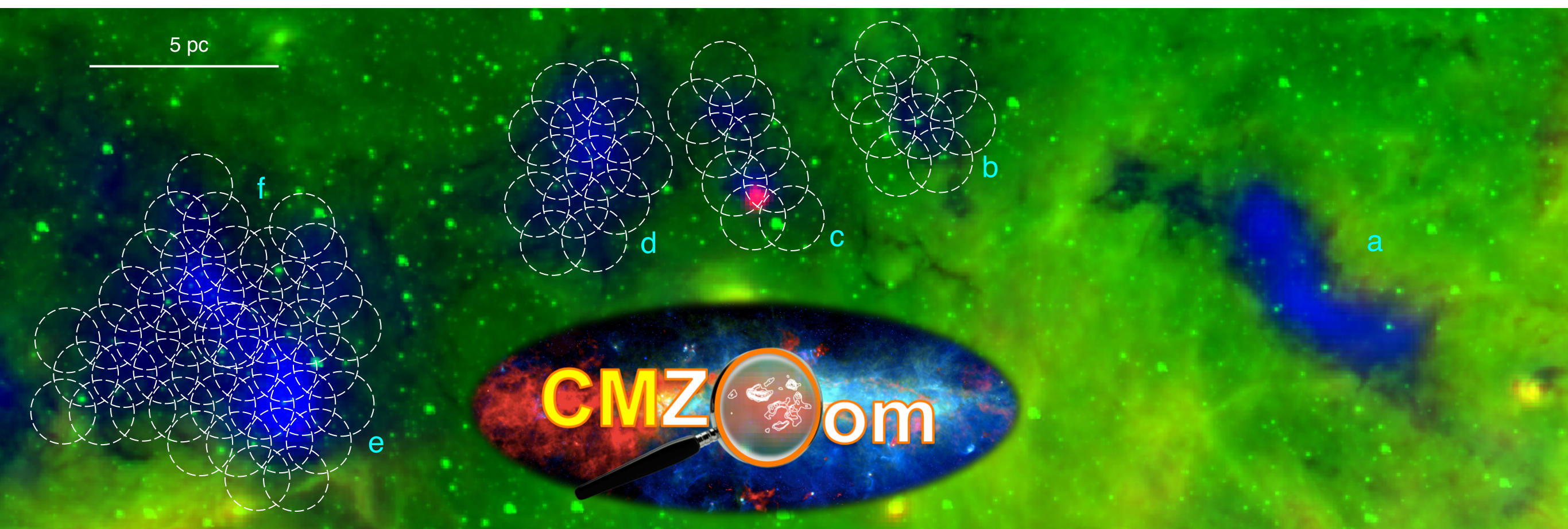
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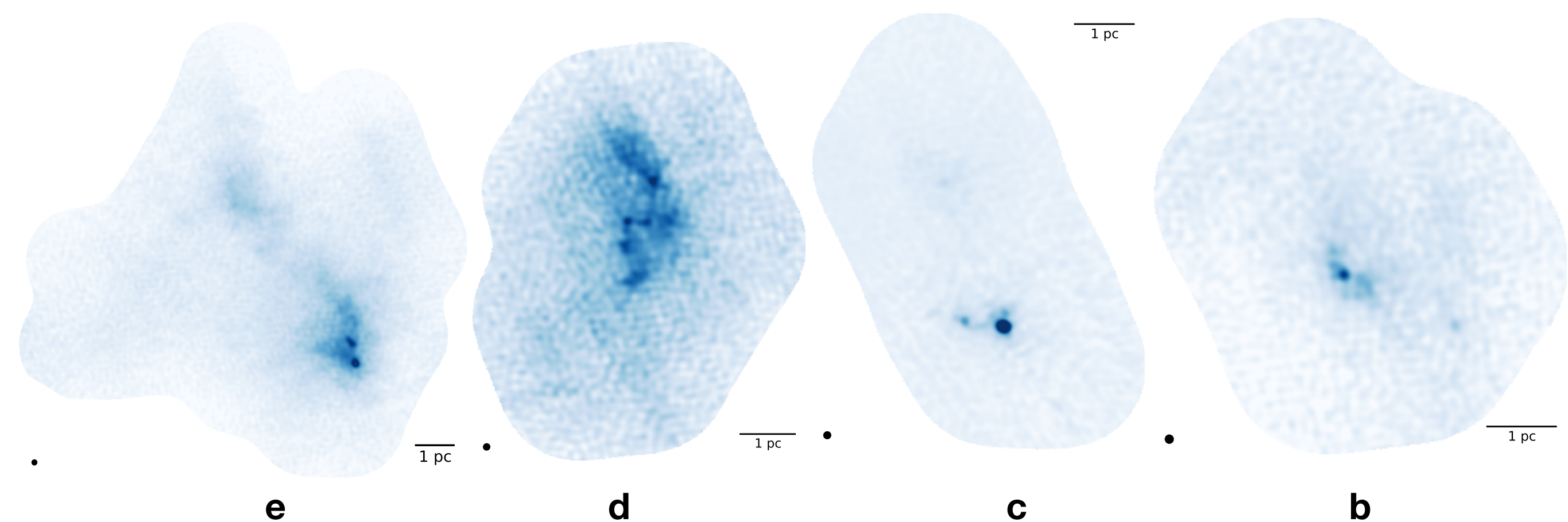
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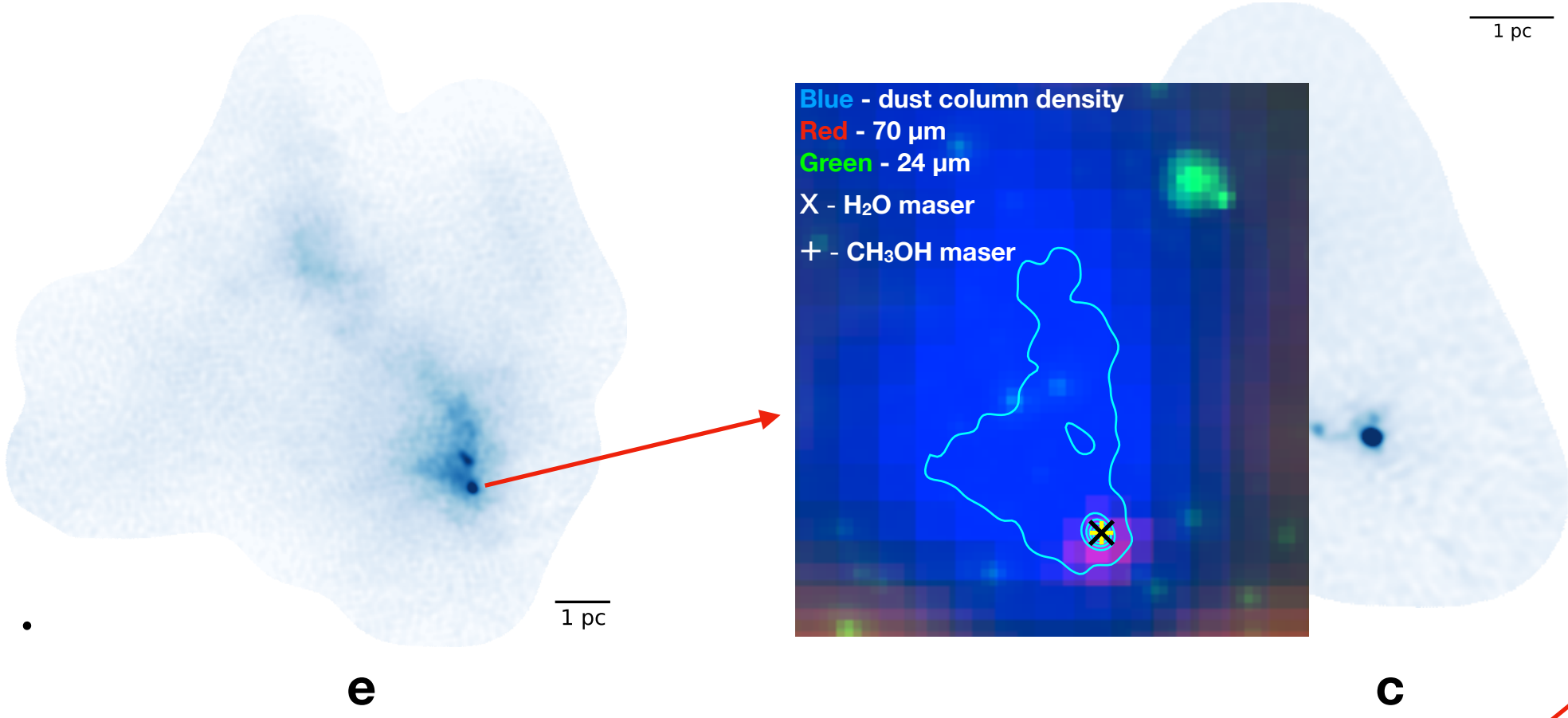


→ Amongst the most massive and compact **quiescent** clouds known to exist in the Galaxy!

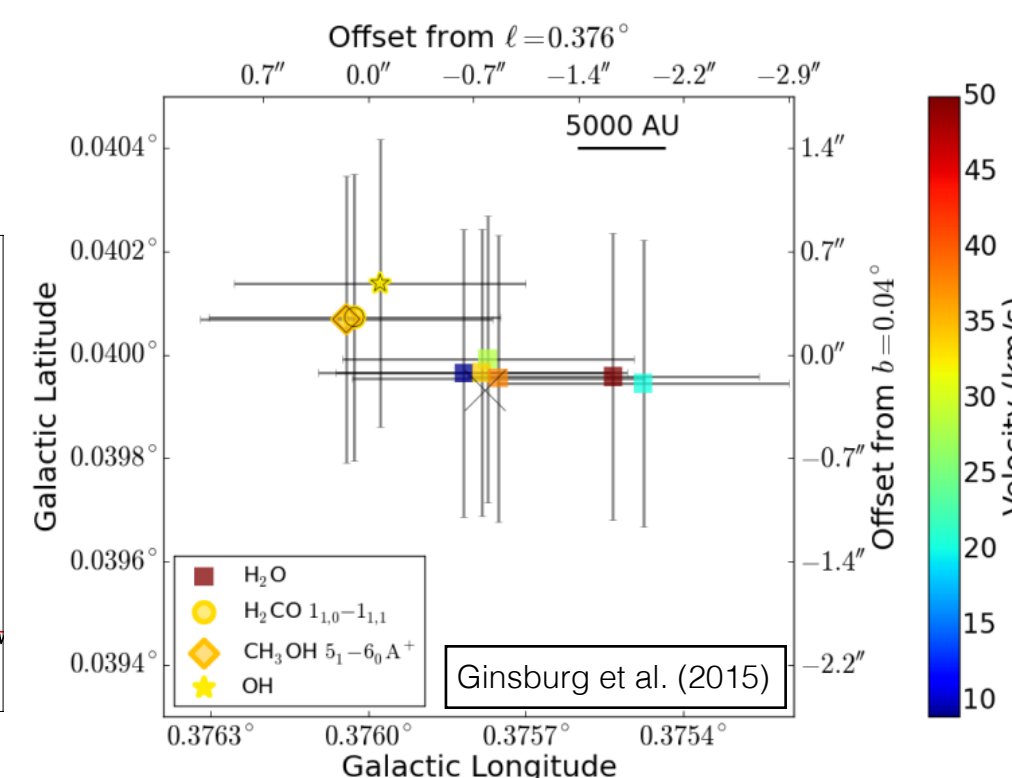
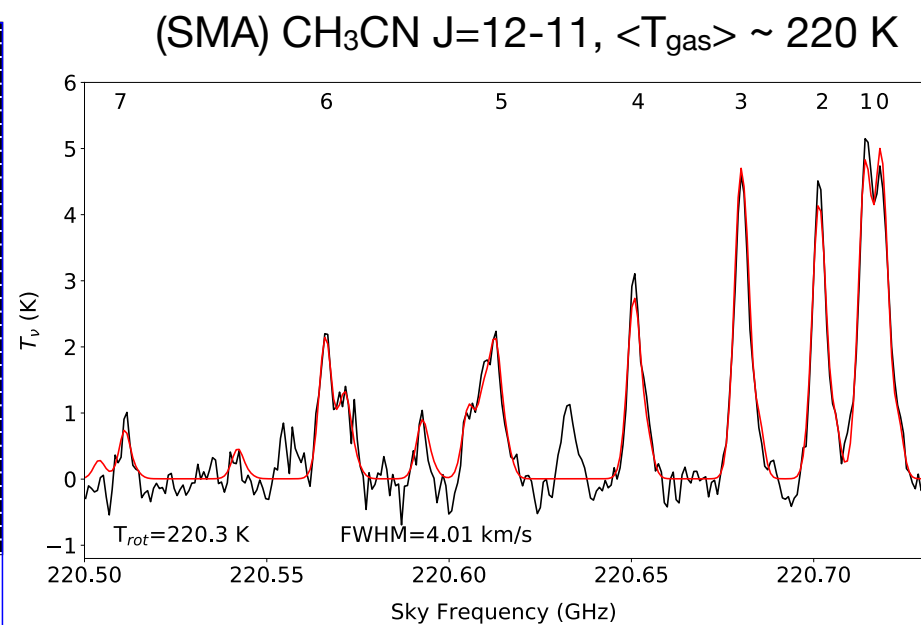
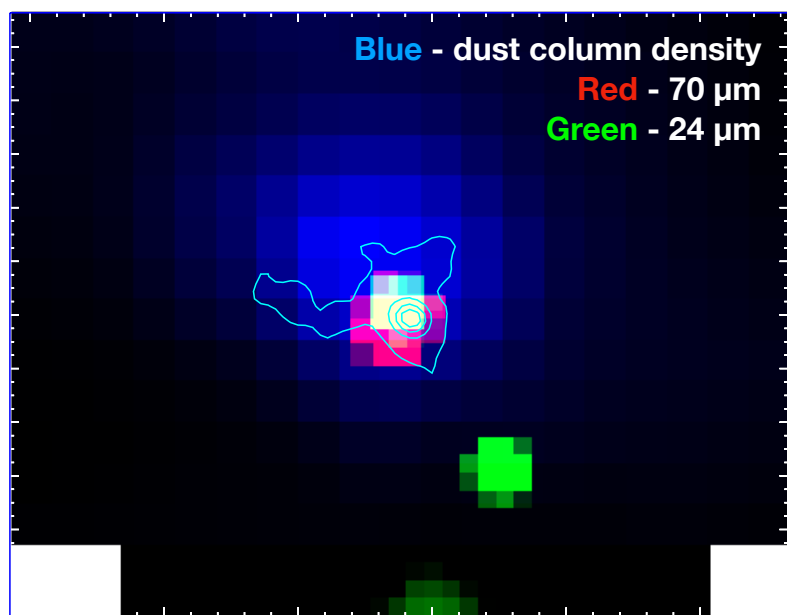




- Detected a robust sample of 15 cores with **$R \sim 0.1 - 0.25 \text{ pc}$** and **$M \sim 50 - > 1000 M_{\odot}$**
- $\Delta V \sim 5 \text{ km/s}$
 - Majority of cores appear to be super-virial
 - Pressure-bounded equilibrium requires $P/k \sim 10^8 \text{ K cm}^{-3}$, consistent with measurements in the CMZ
- Only **2** of the **15** cores show signatures of embedded star formation, despite all being **$> 10^6 \text{ cm}^{-3}$**
 - *Further evidence for a heightened critical density threshold for SF in the CMZ*

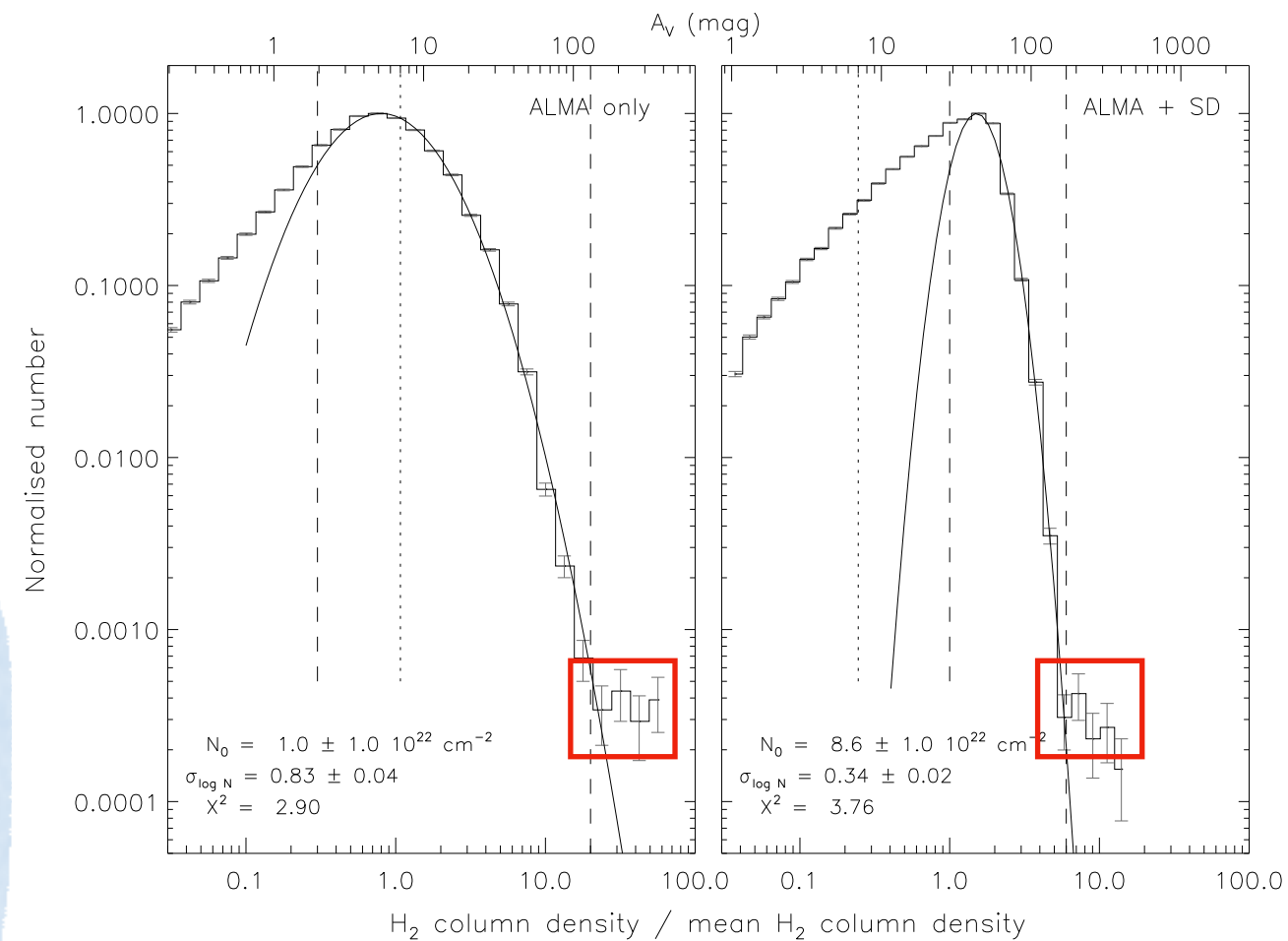
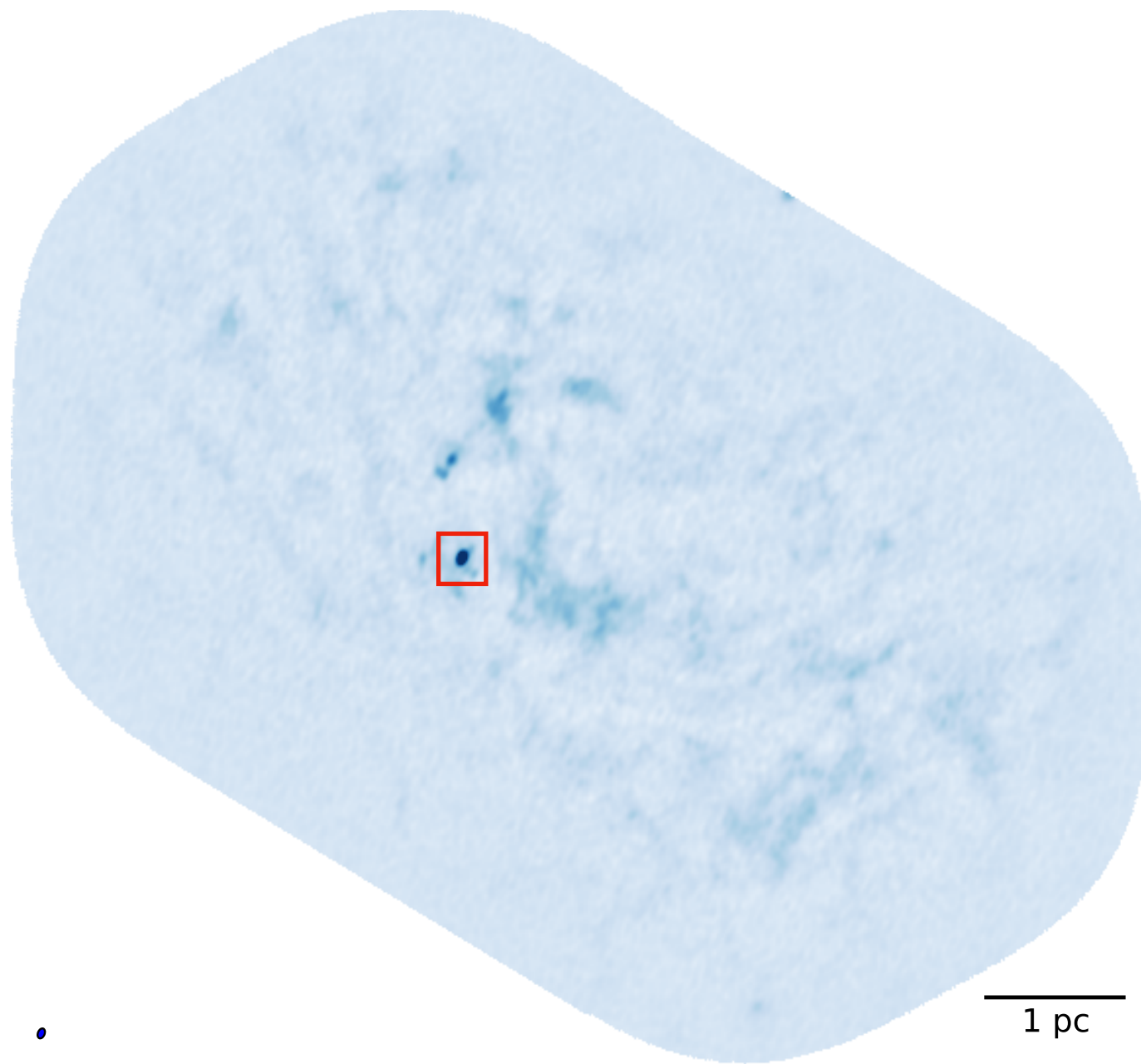


Two newly discovered high-mass star-forming regions in the dust ridge!



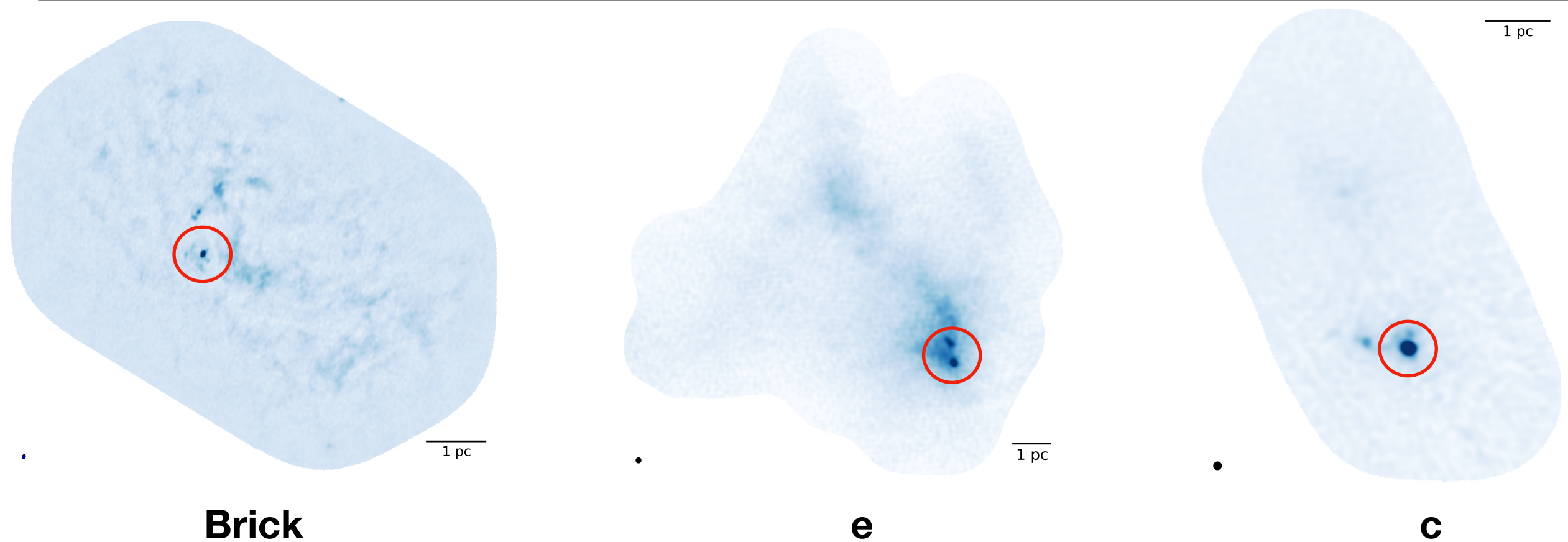
ALMA follow-up observations ... but first ...

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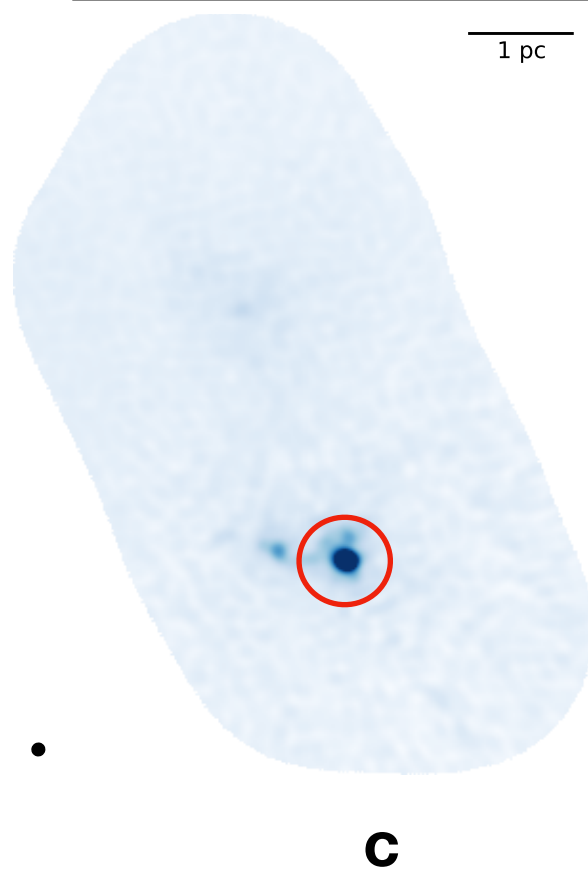
Mass $\sim 100 M_\odot$ and coincides with water maser emission $\rightarrow$ potential site of (massive) star formation in the Brick

ALMA follow-up observations

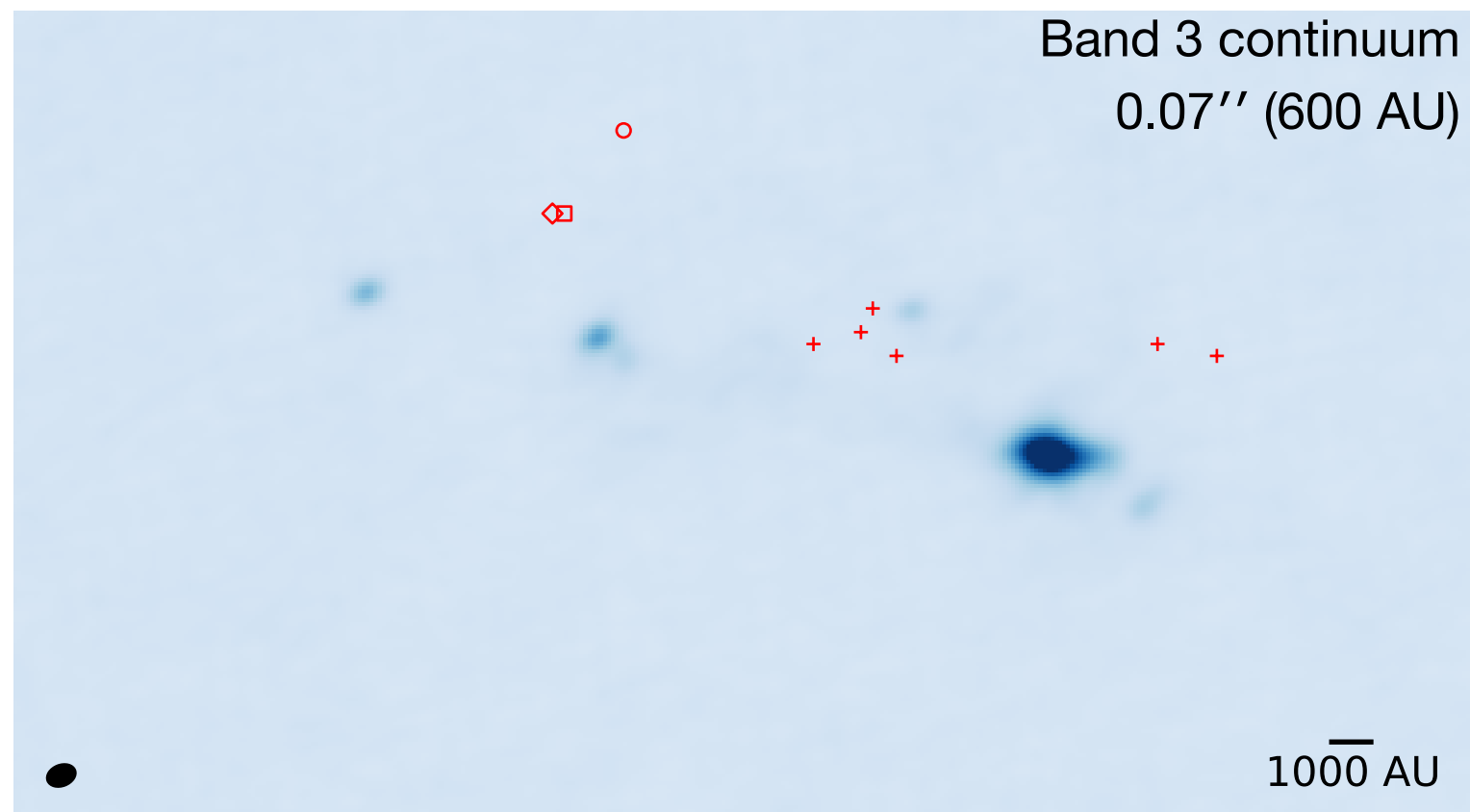
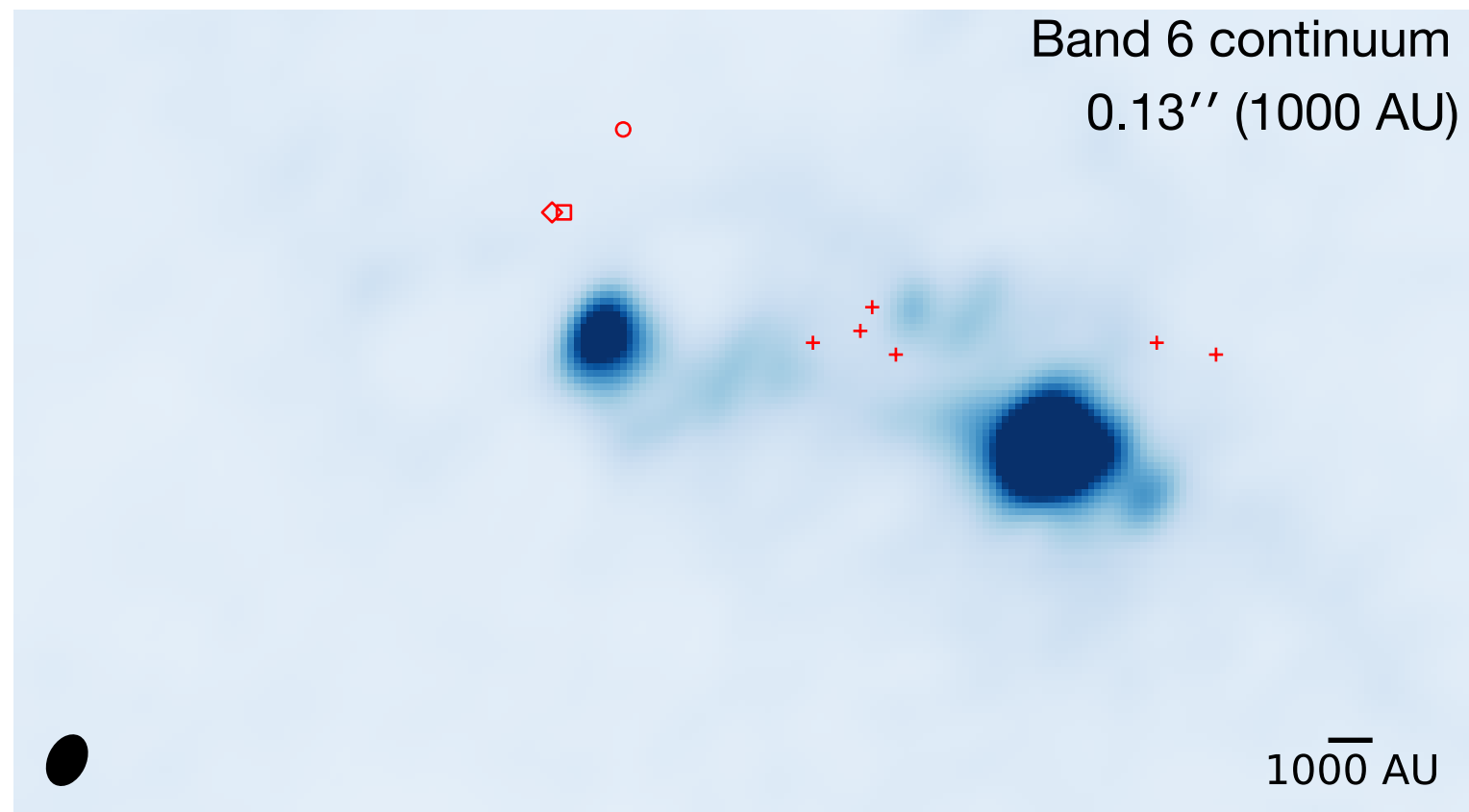
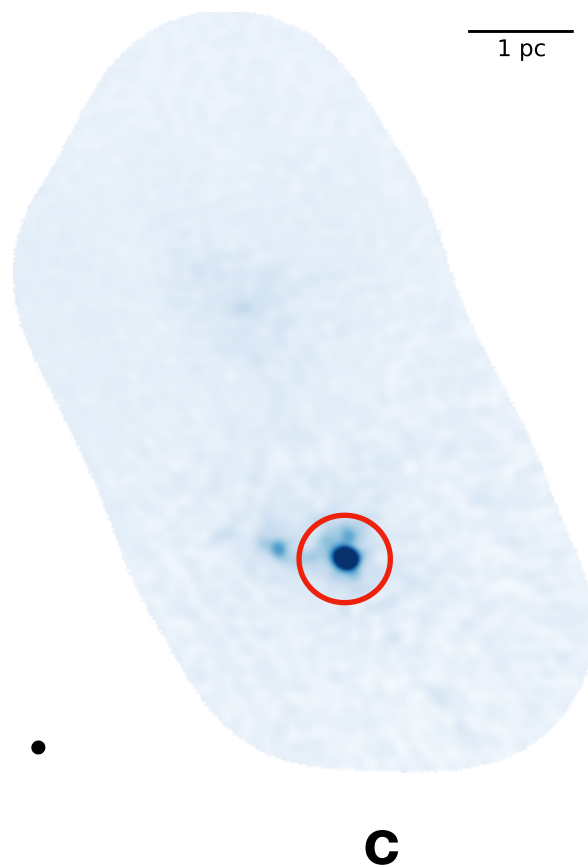


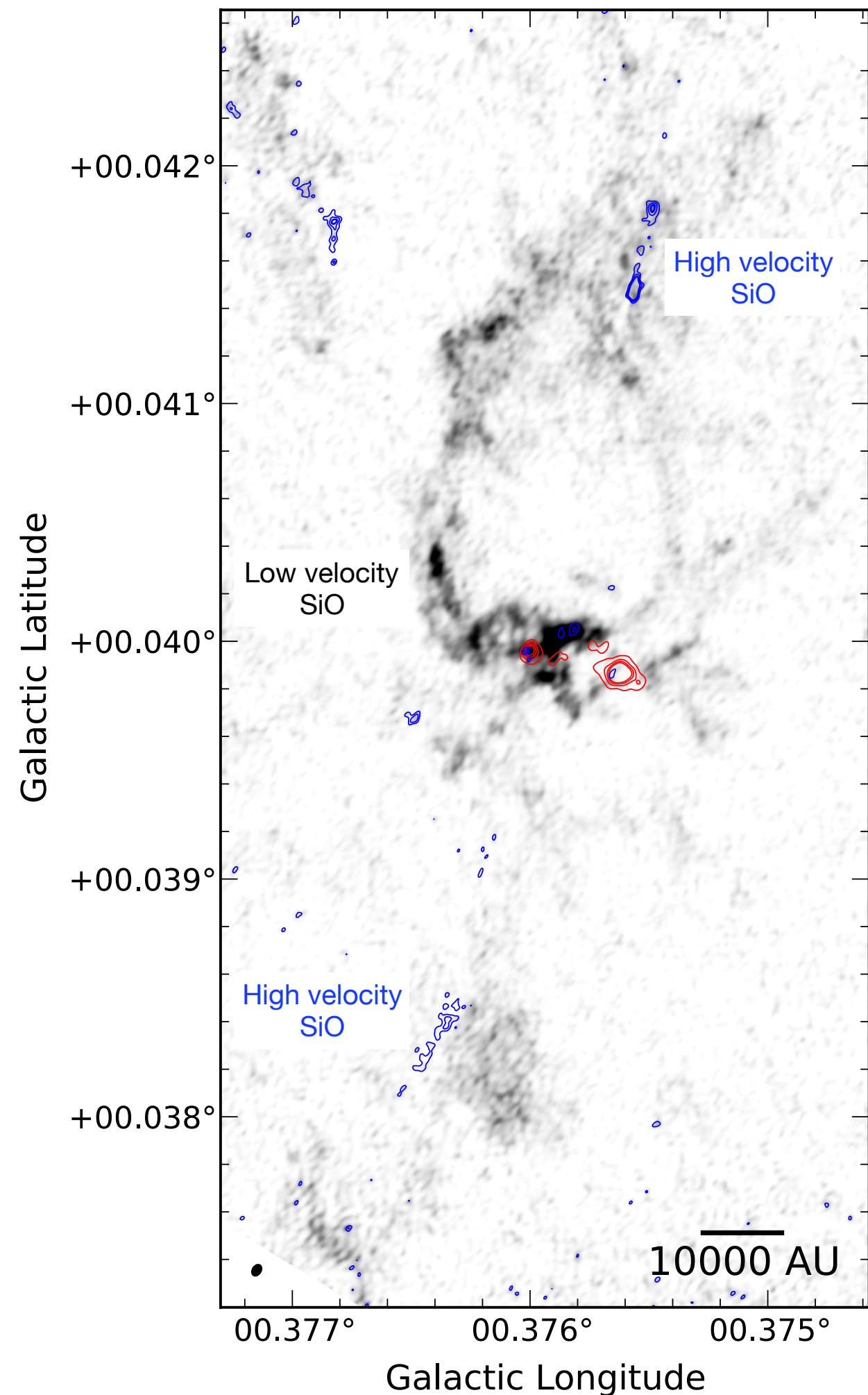
- Observed with ALMA in band 6 (~ 230 GHz)
 - Angular resolution $\sim 0.13''$ (1000 AU)
 - Velocity resolution ~ 0.7 km/s, targeting CH_3CN , SiO , H_2CO , ^{13}CO & many more
- Also observed dust ridge cloud **c** in band 3 at $\sim 0.07''$ (600 AU)

ALMA follow-up observations — dust ridge cloud C

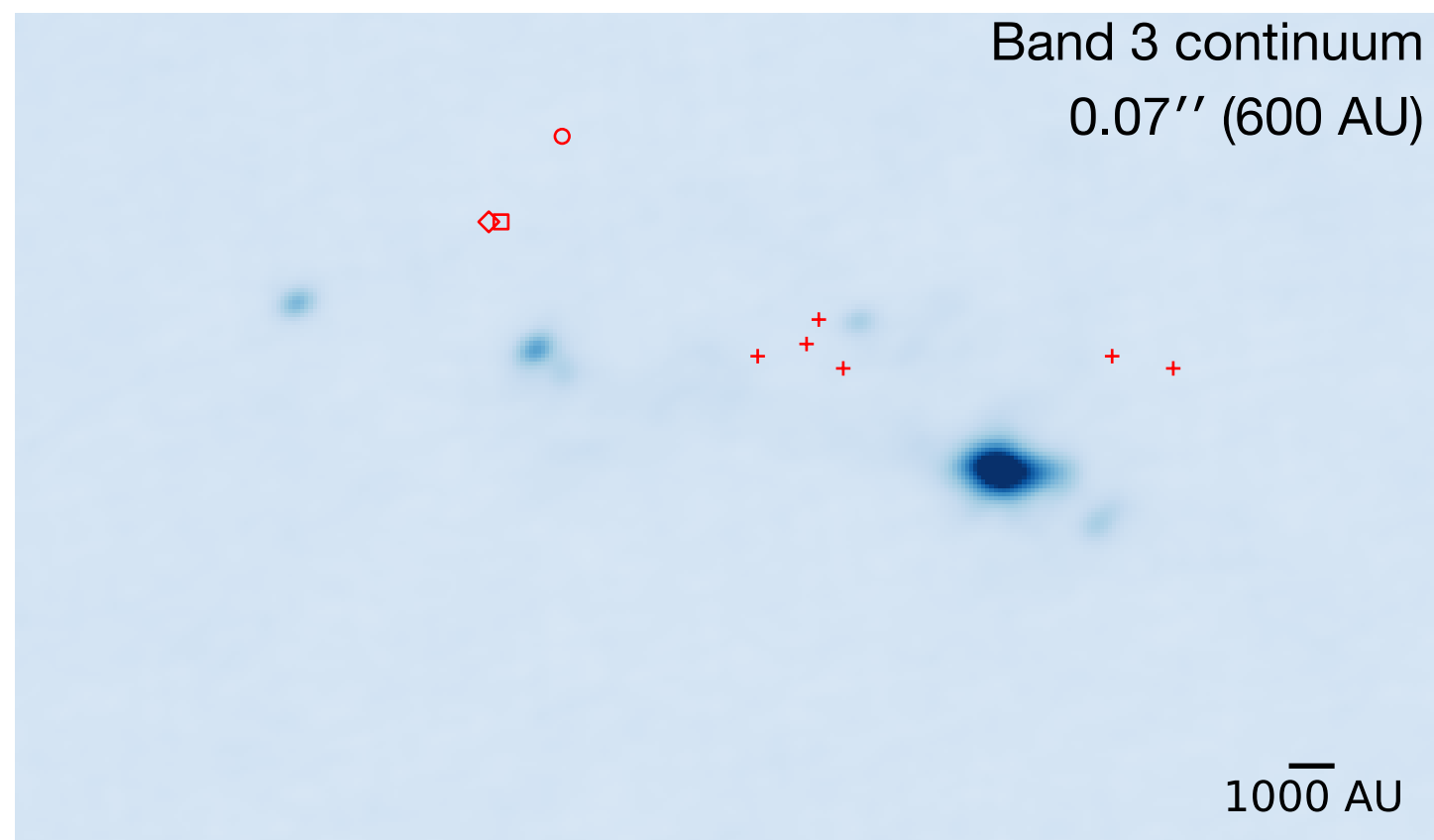
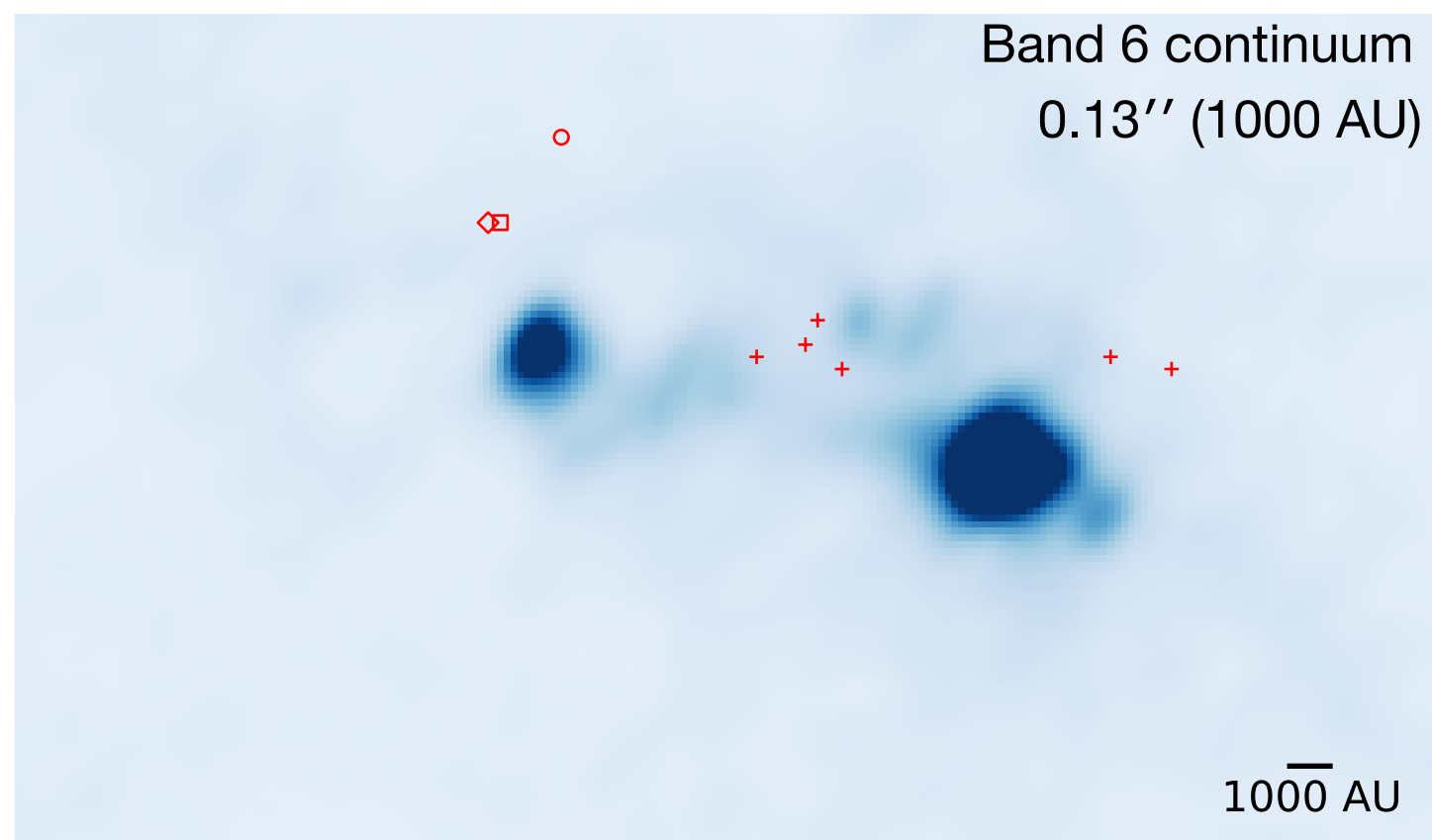


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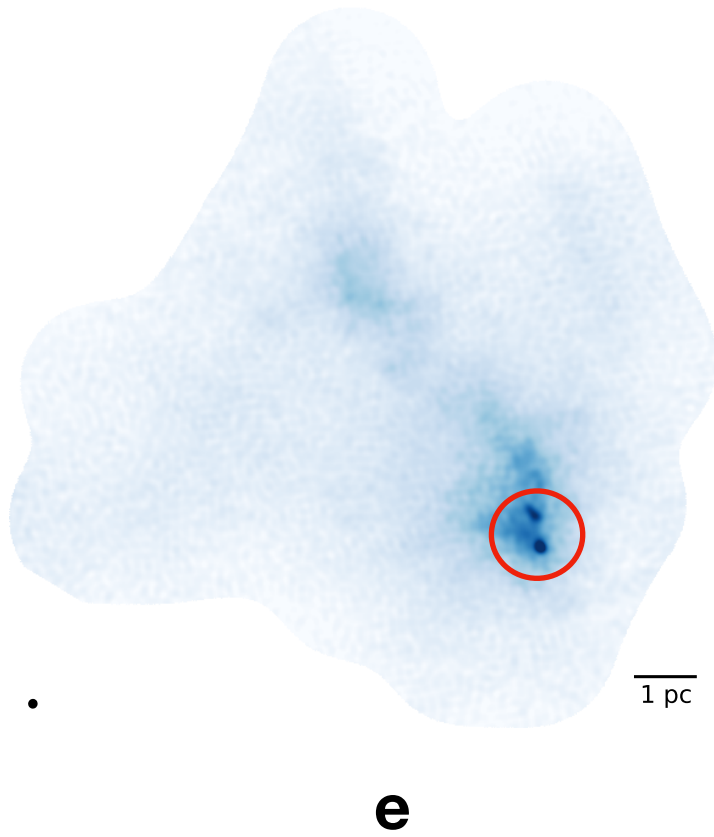




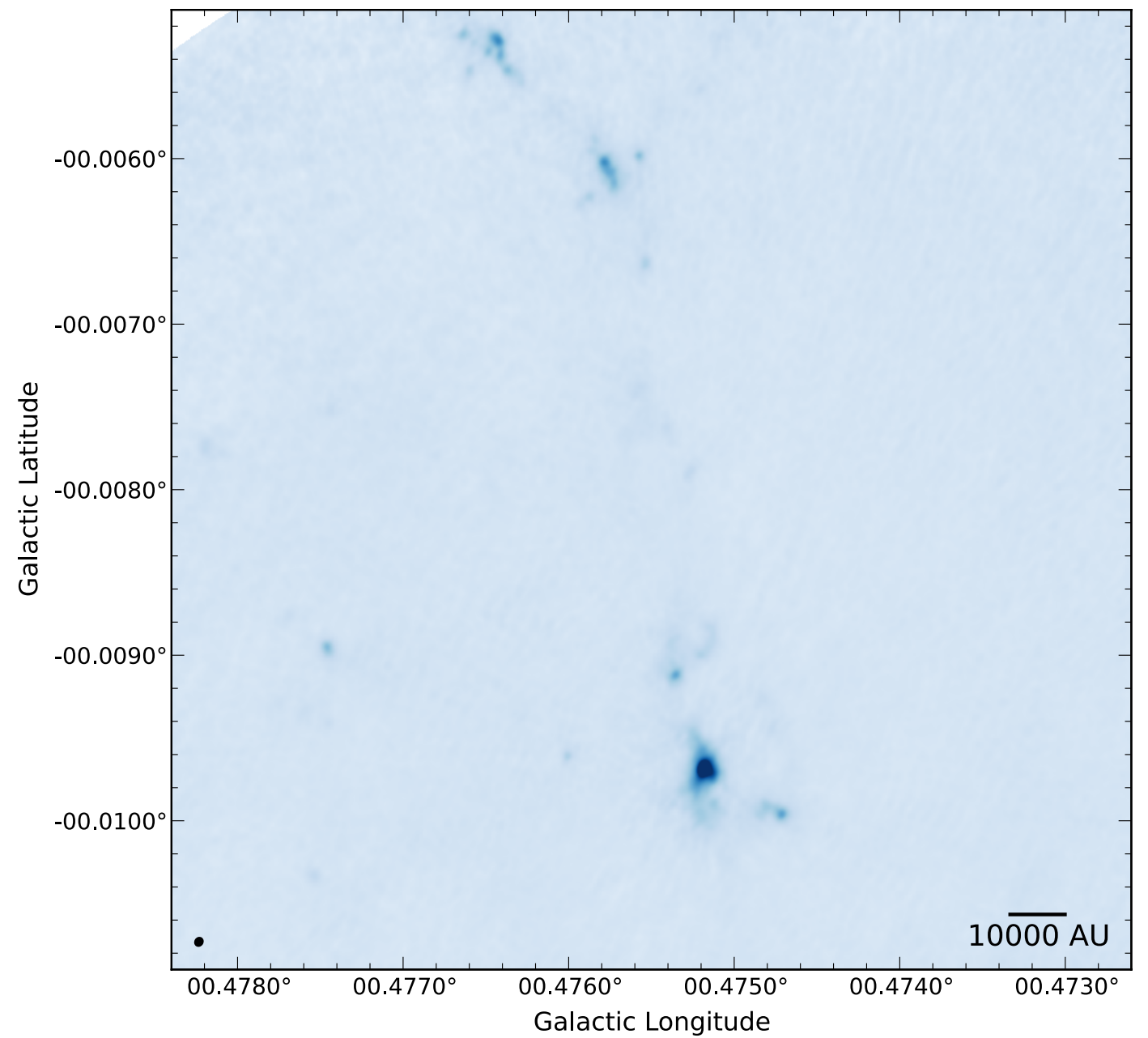
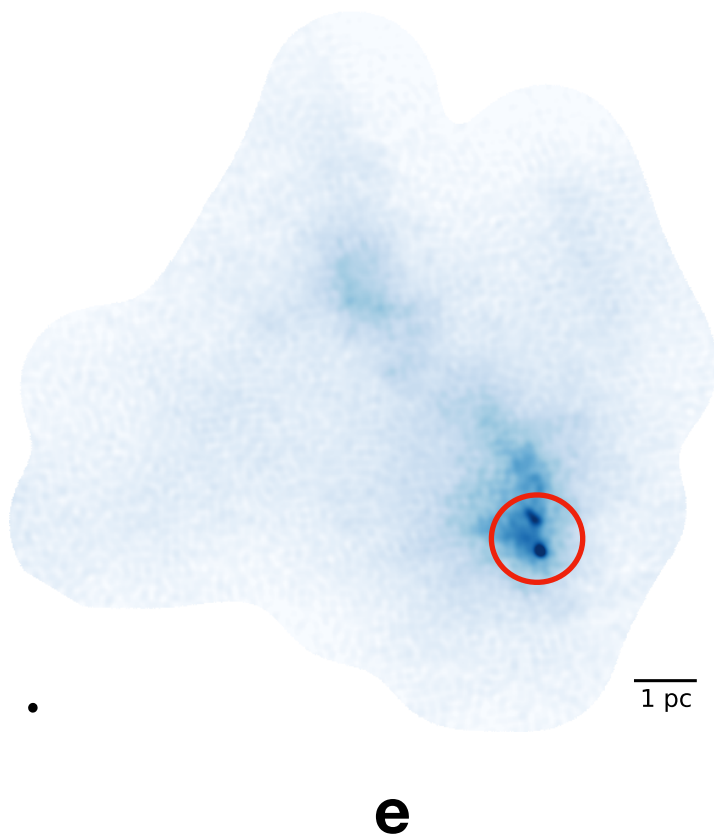
s — dust ridge cloud C



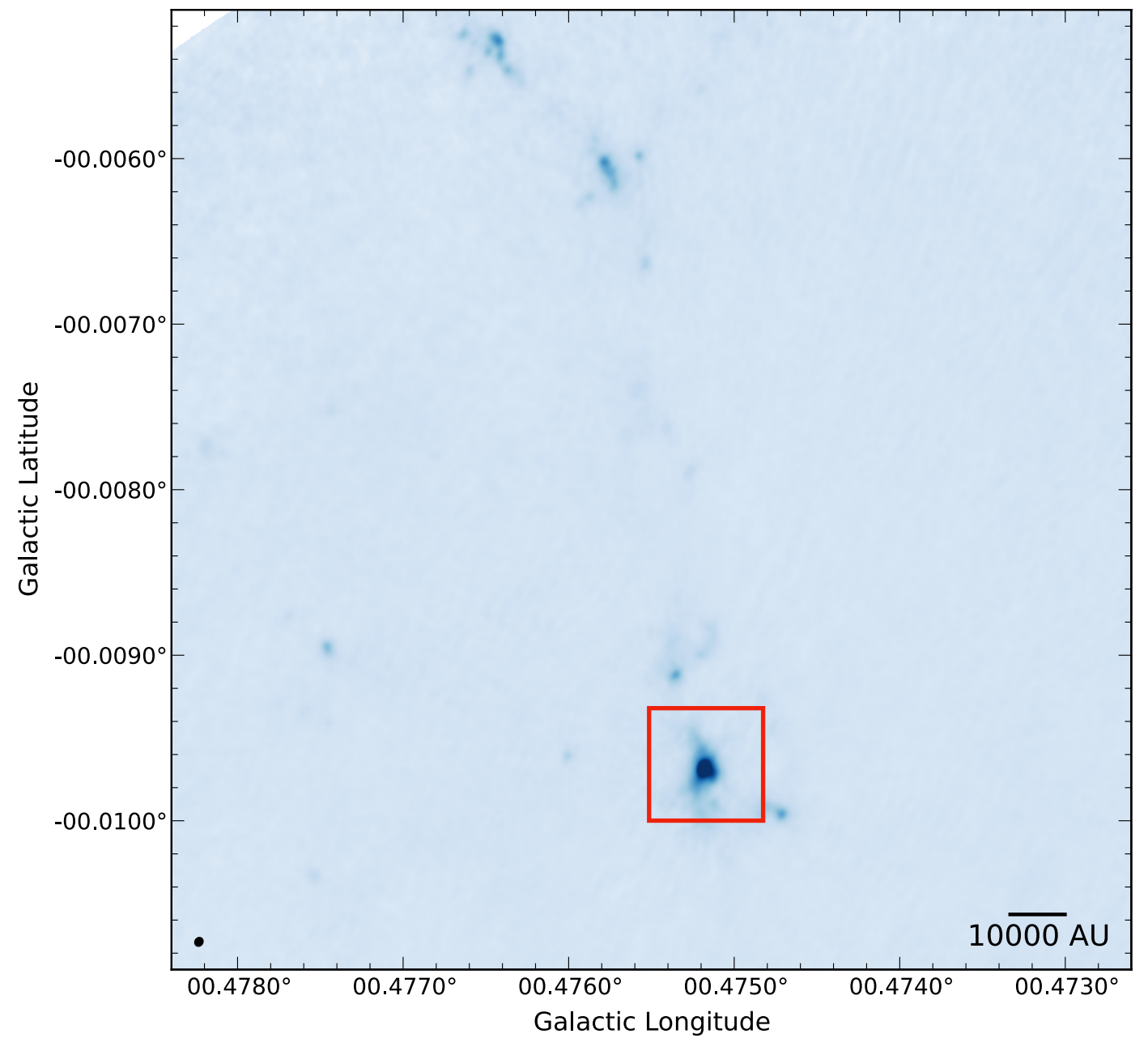
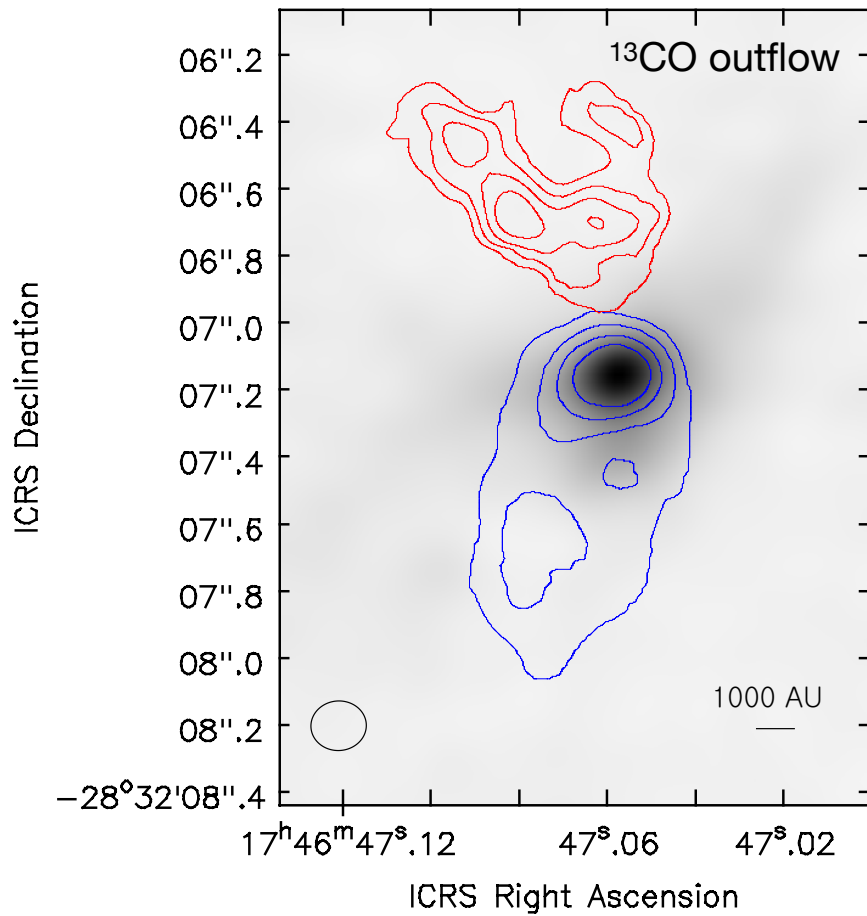
ALMA follow-up observations — dust ridge cloud E



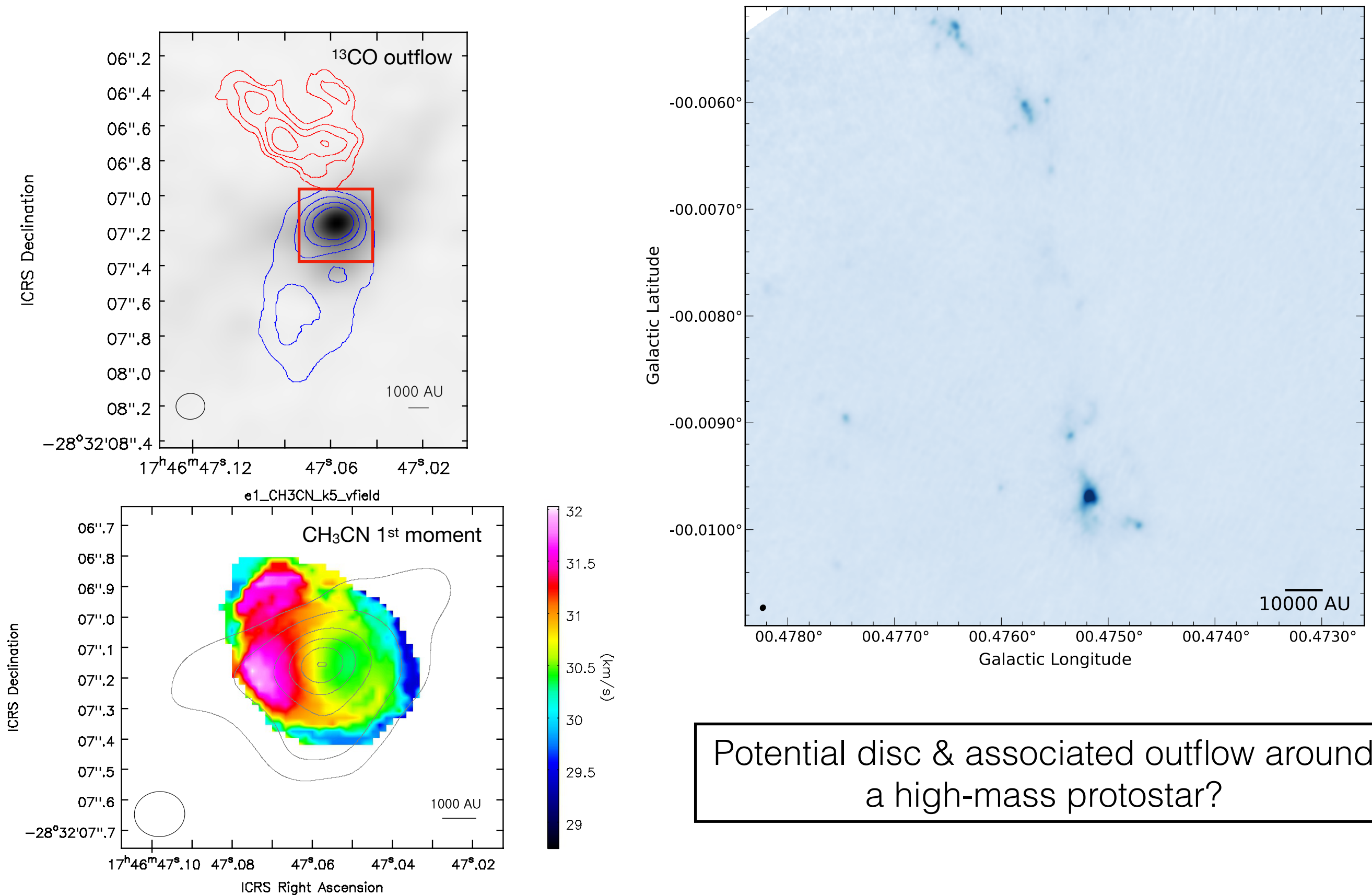
ALMA follow-up observations — dust ridge cloud E



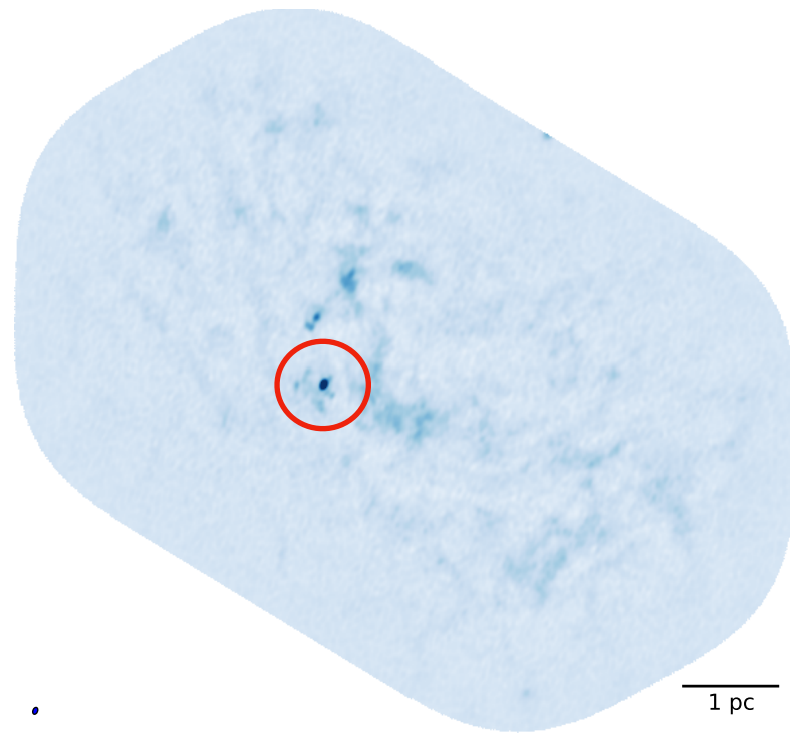
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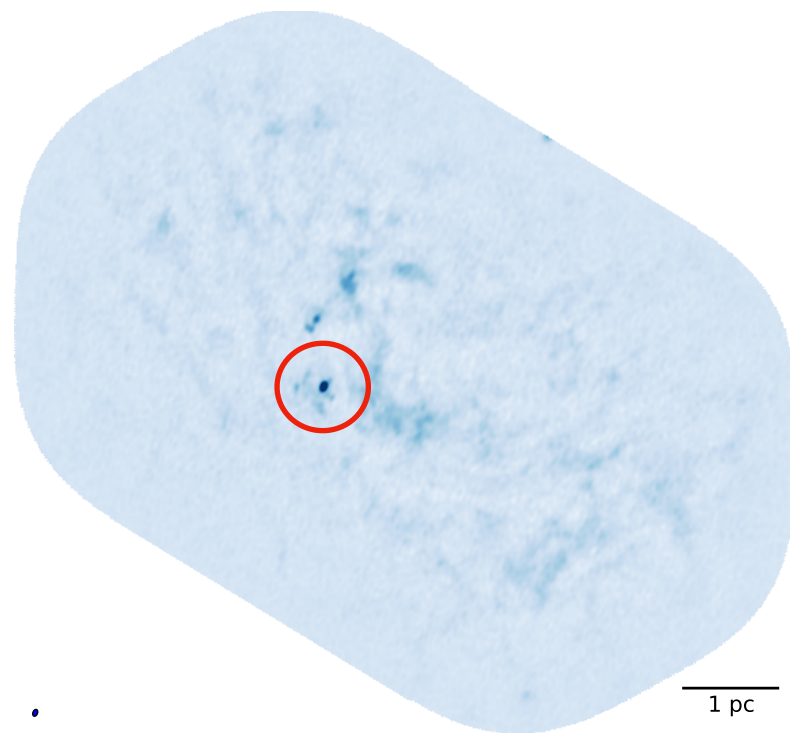


ALMA follow-up observations — the Brick ‘maser core’

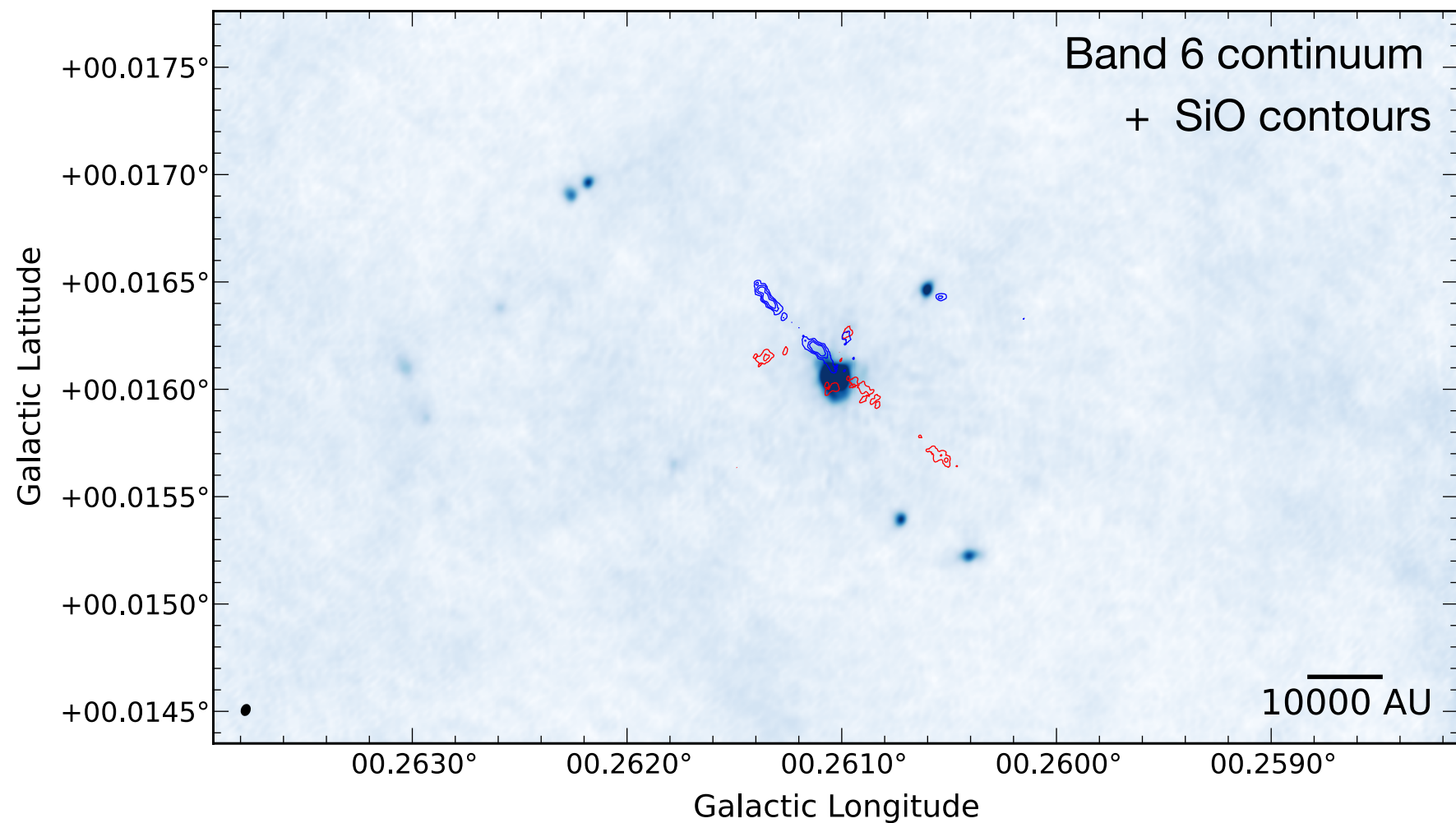


Brick

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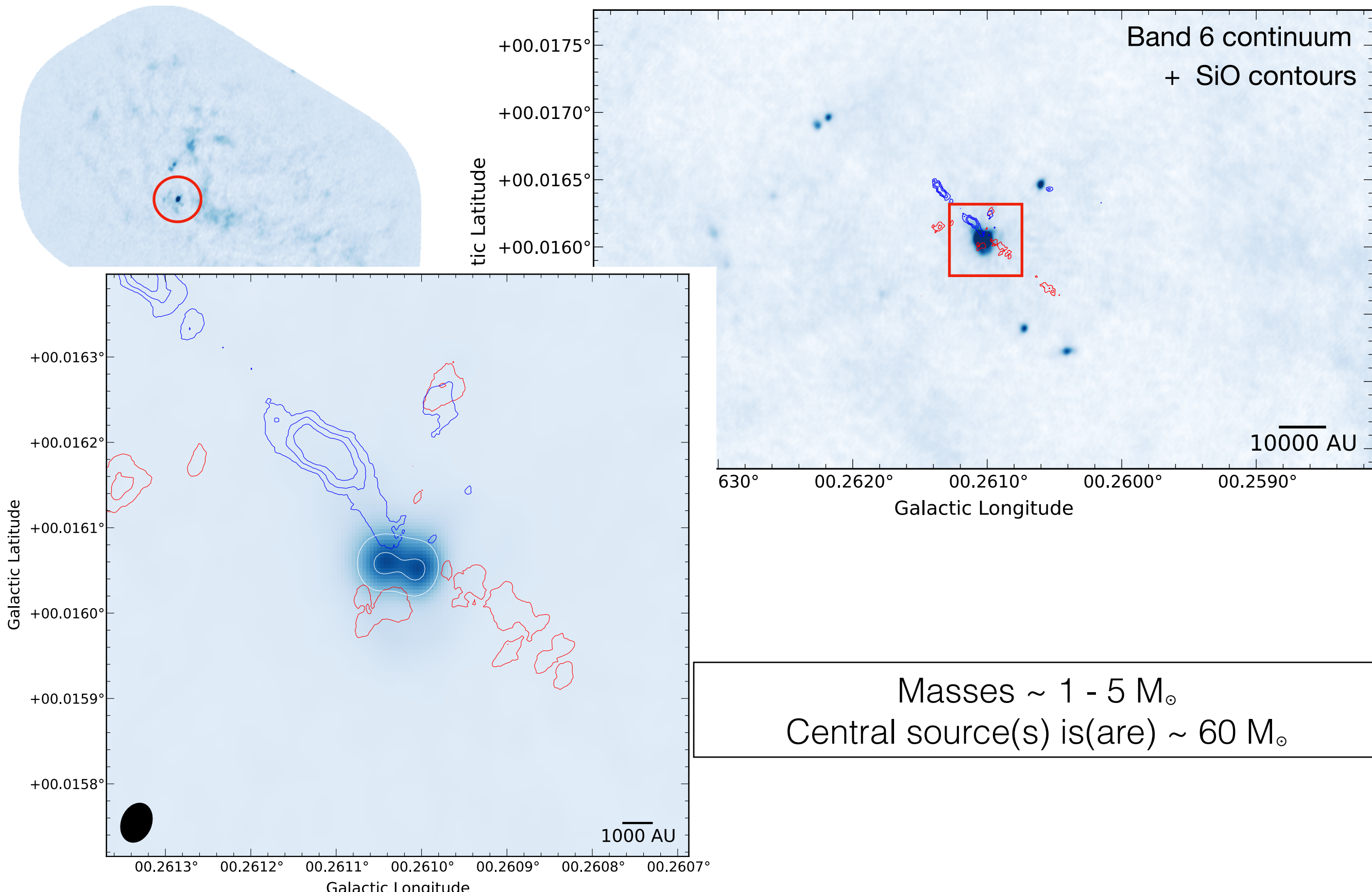


Brick



Masses $\sim 1 - 5 M_{\odot}$
Central source is $\sim 60 M_{\odot}$

ALMA follow-up observations — the Brick ‘maser core’



Summary

- SMA *CMZoom* observations find a population of embedded dense cores in the Galactic centre dust ridge —
 - > 80% of the cores are quiescent, despite densities $> 10^6 \text{ cm}^{-3}$
 - Further evidence for environmentally-dependent critical density threshold for SF
 - Two new sites of embedded high-mass star formation
- Follow-up ALMA observations reveal detailed sub-structure on $< 1000 \text{ AU}$ scales, showing significant fragmentation, suggesting that the low- and high-mass sources are forming simultaneously
- All 3 targeted massive cores are driving bipolar outflows in SiO and ^{13}CO — this is the first direct evidence of outflows in the CMZ in young (pre-UCHII) protostars
- One source in cloud E shows tantalising evidence of disc-like kinematic structure — if confirmed, this would be the first direct detection of a (massive) disc in the CMZ