



Gravity drives the evolution of infrared dark hubs: JVLA observations of SDC13

Gwenllian Williams

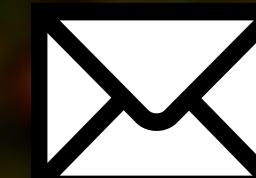
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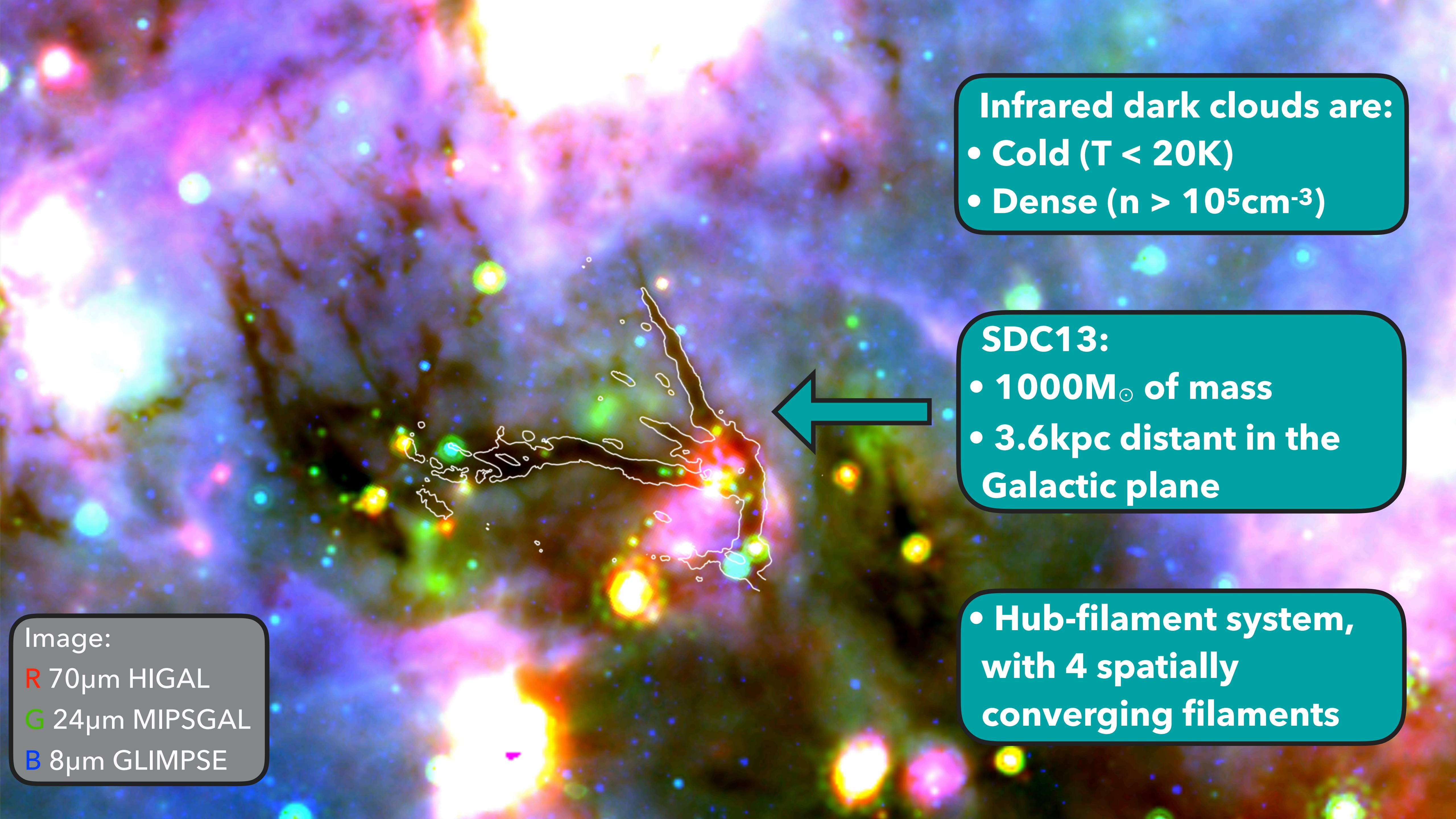


Image:

R 70μm HIGAL

G 24μm MIPS GAL

B 8μm GLIMPSE

Infrared dark clouds are:

- Cold ($T < 20K$)
- Dense ($n > 10^5 \text{cm}^{-3}$)

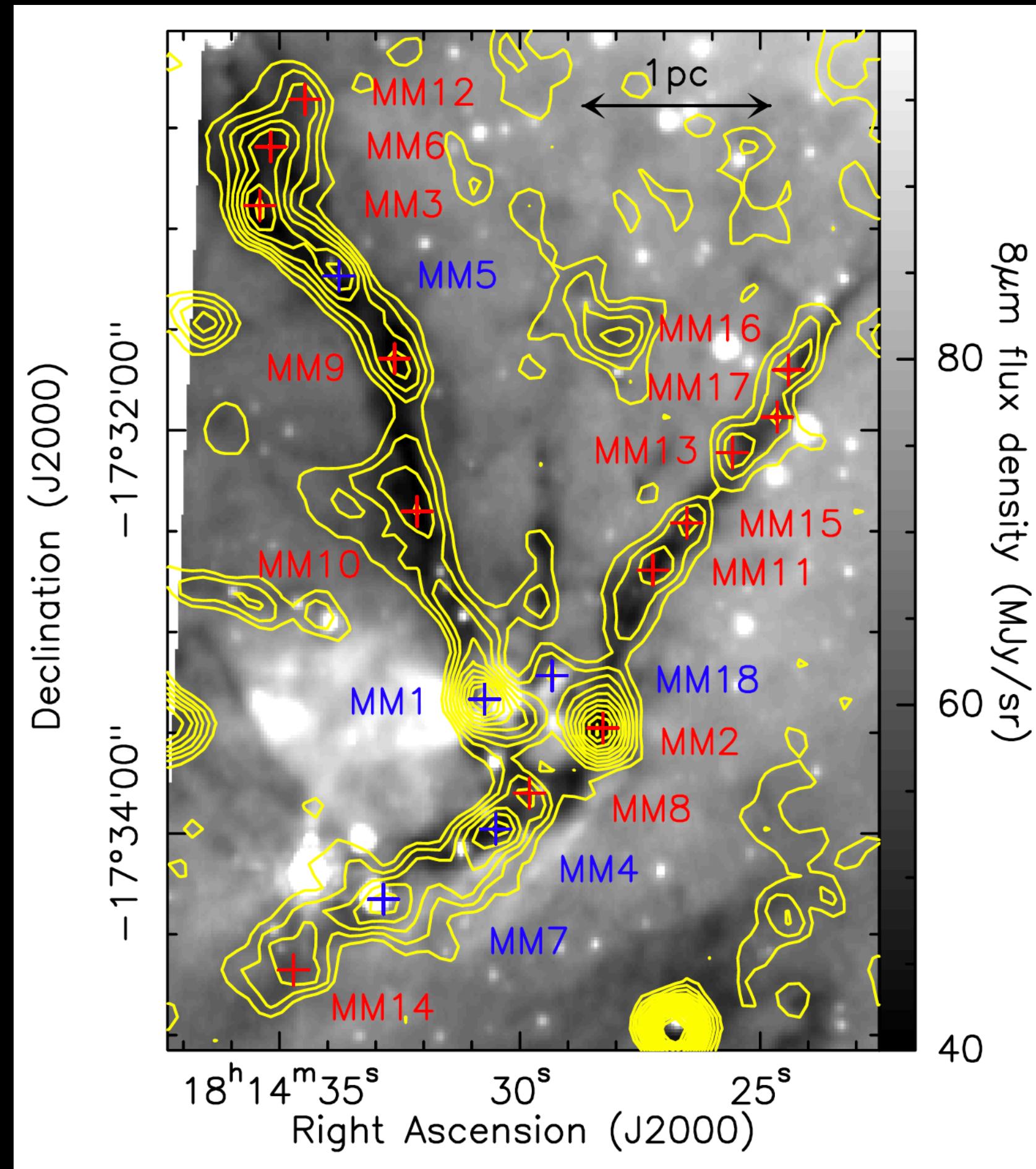
SDC13:

- $1000M_{\odot}$ of mass
- 3.6kpc distant in the Galactic plane

- Hub-filament system, with 4 spatially converging filaments

Previous work

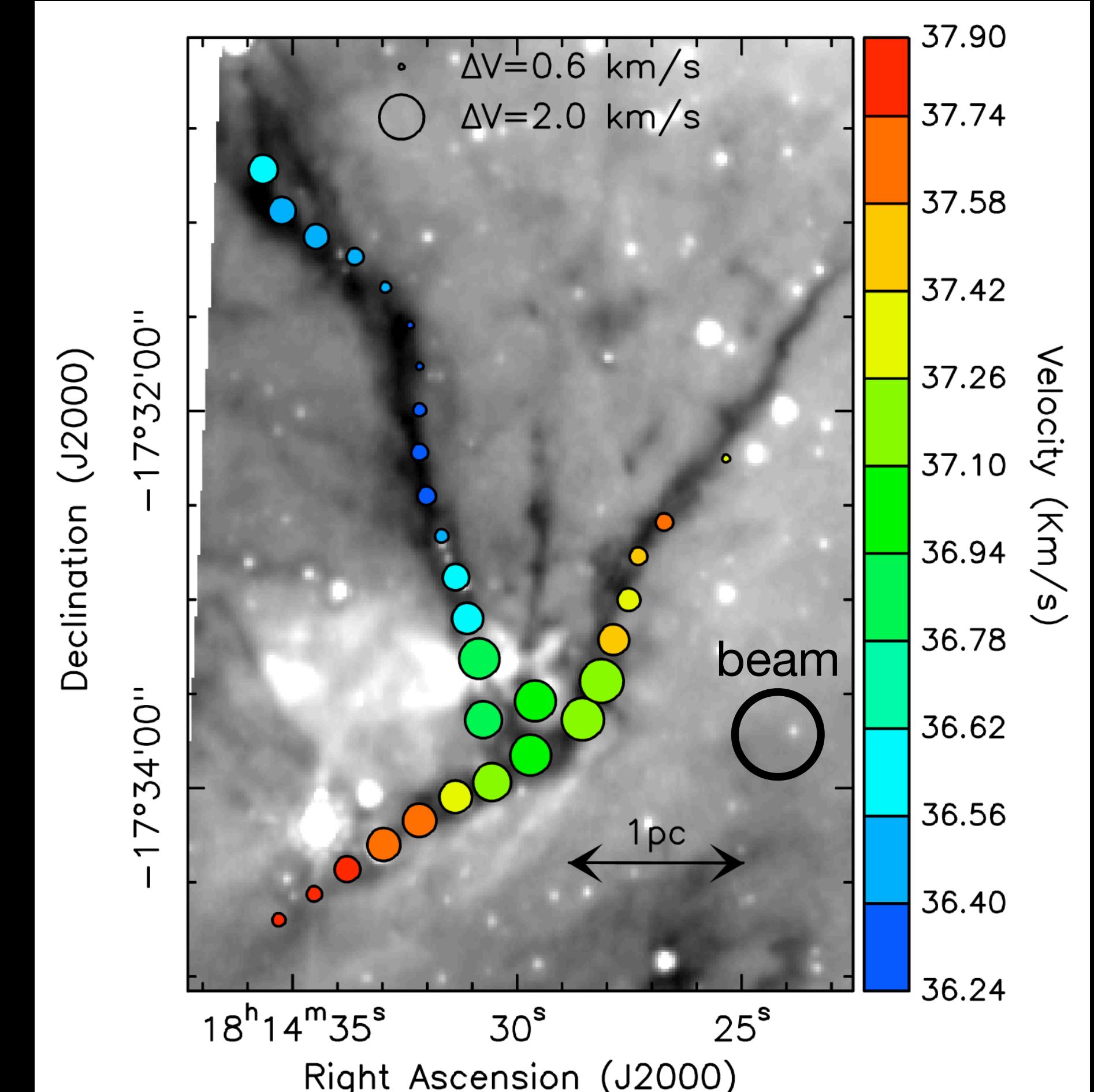
18 compact mm sources



+

 Starless + Protostellar

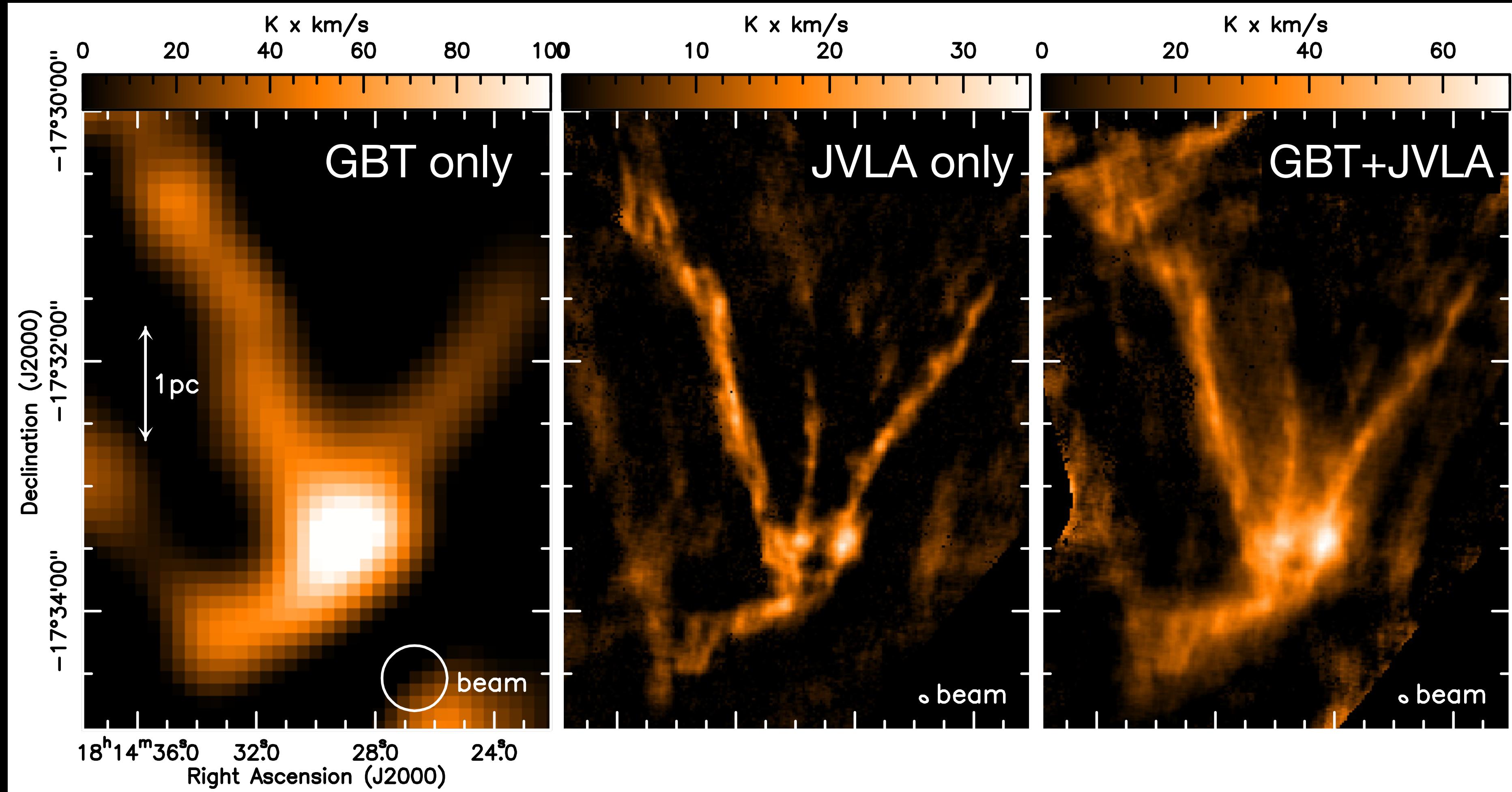
$\text{N}_2\text{H}^+(1-0)$ centroid velocity



Ordered velocity structure

New data: combining GBT and JVLA data

Integrated intensity maps of $\text{NH}_3(1,1)$

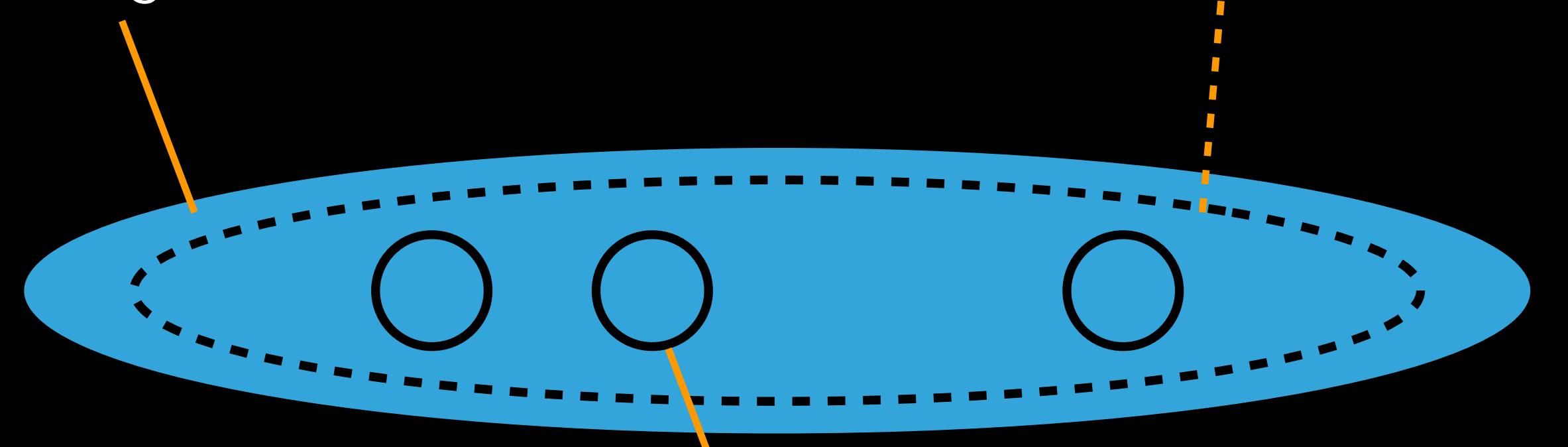


- Resolution of $\sim 4''$, $\sim 0.07\text{pc}$
- Able to resolve mean core sizes of $\sim 0.1\text{pc}$

Fragmentation

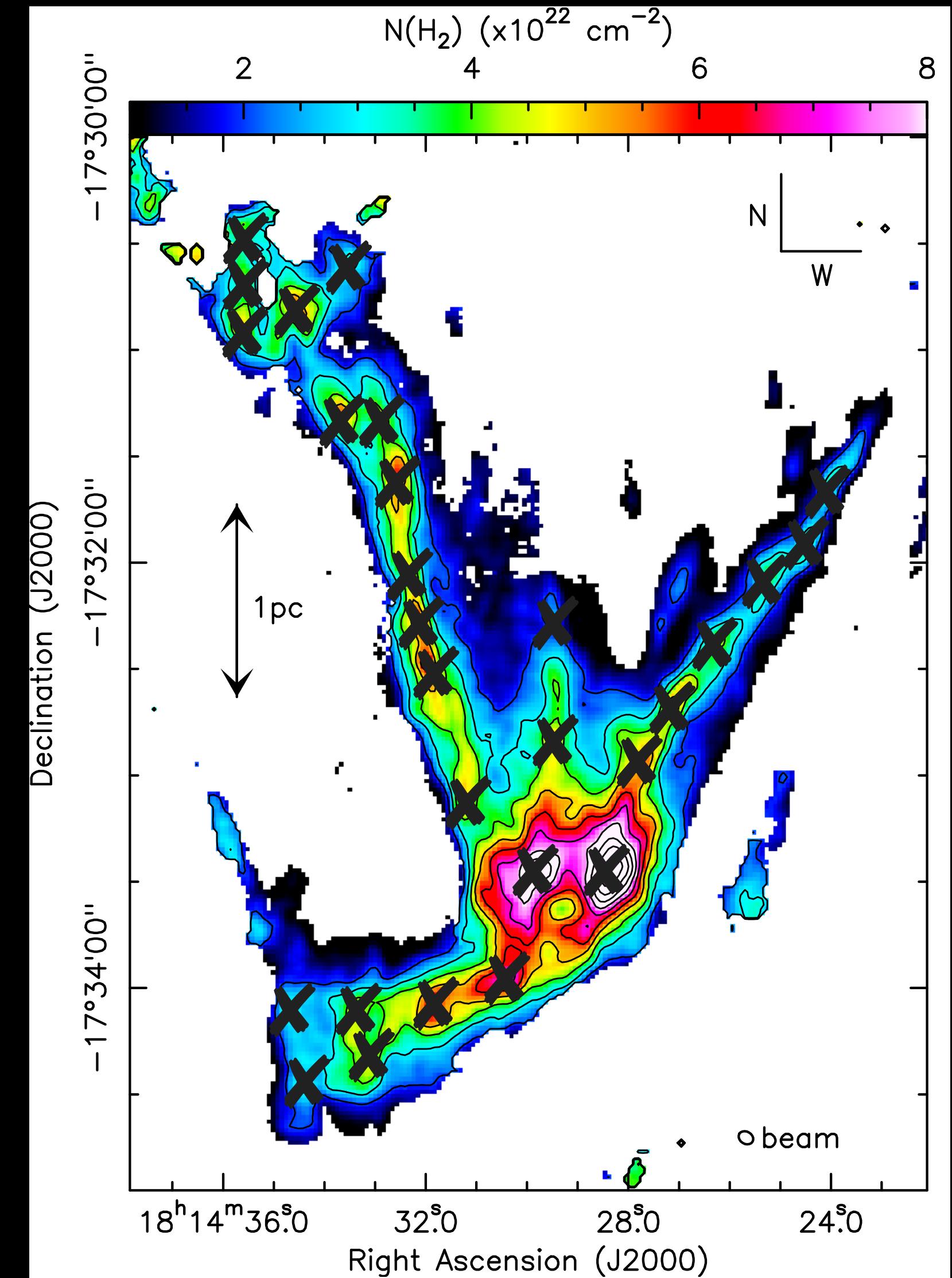
- Dendrogram analysis used to identify structures:

Filament - Base of dendrogram tree



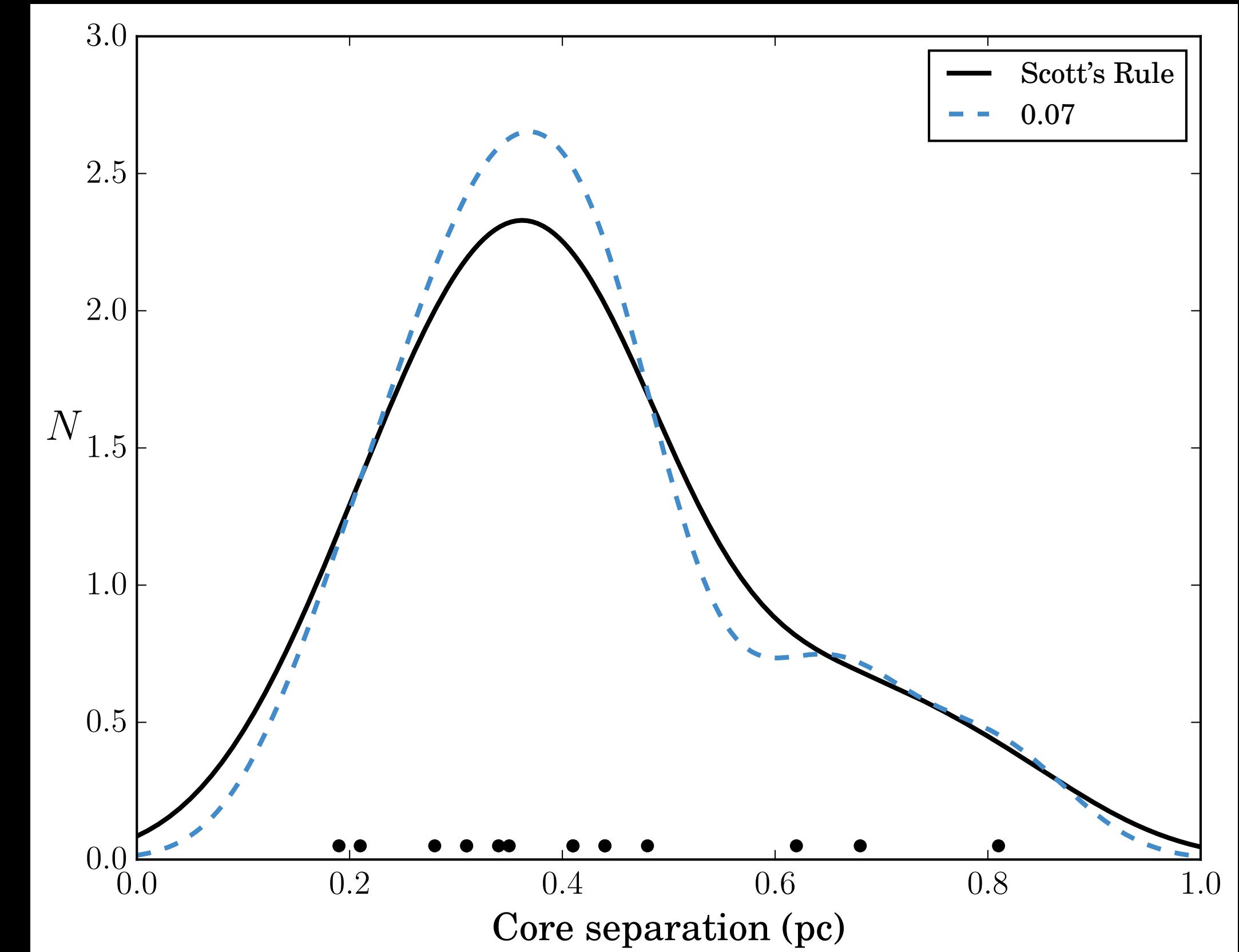
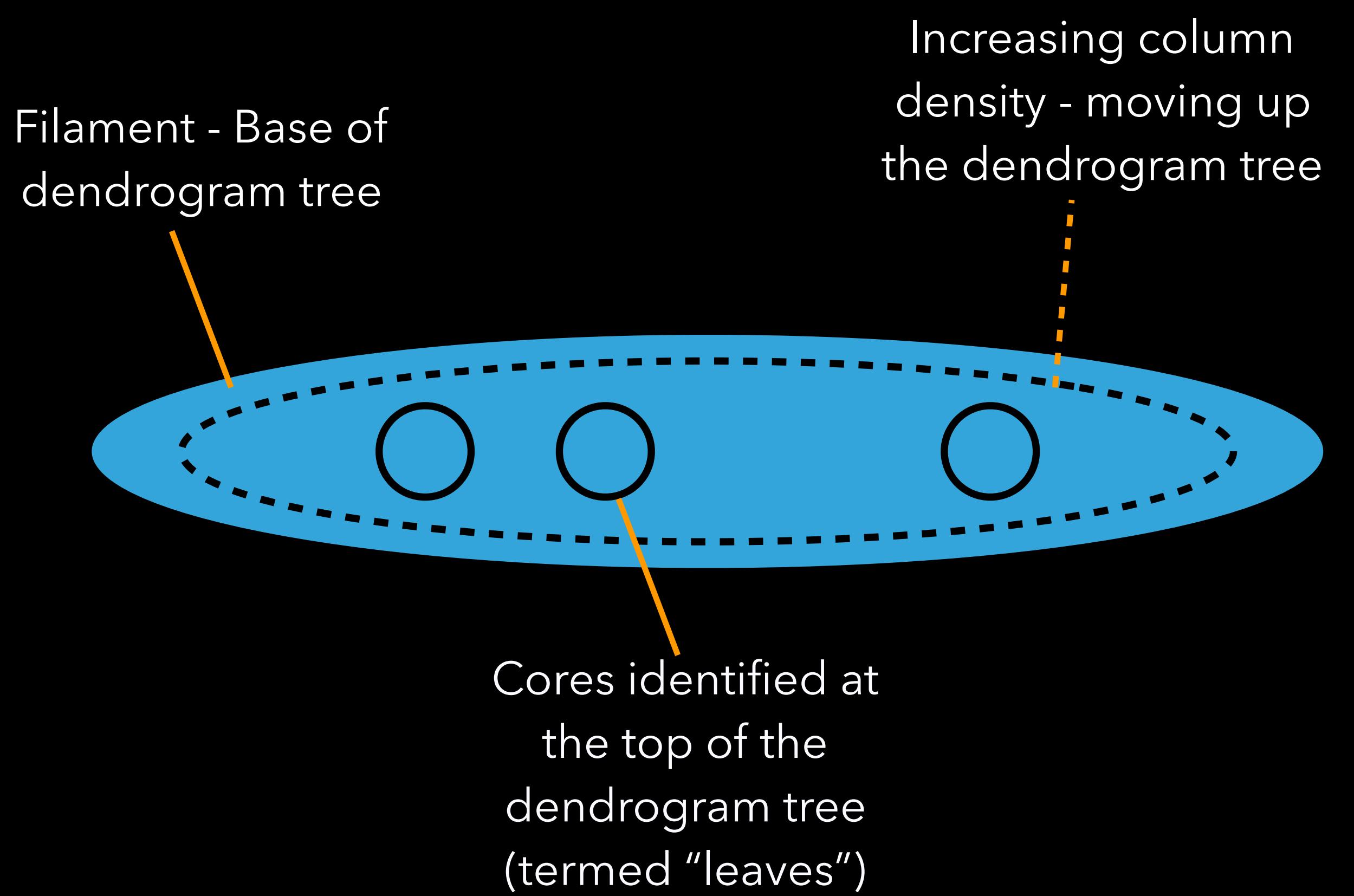
Increasing column density - moving up the dendrogram tree

Cores identified at the top of the dendrogram tree (termed "leaves")



Fragmentation

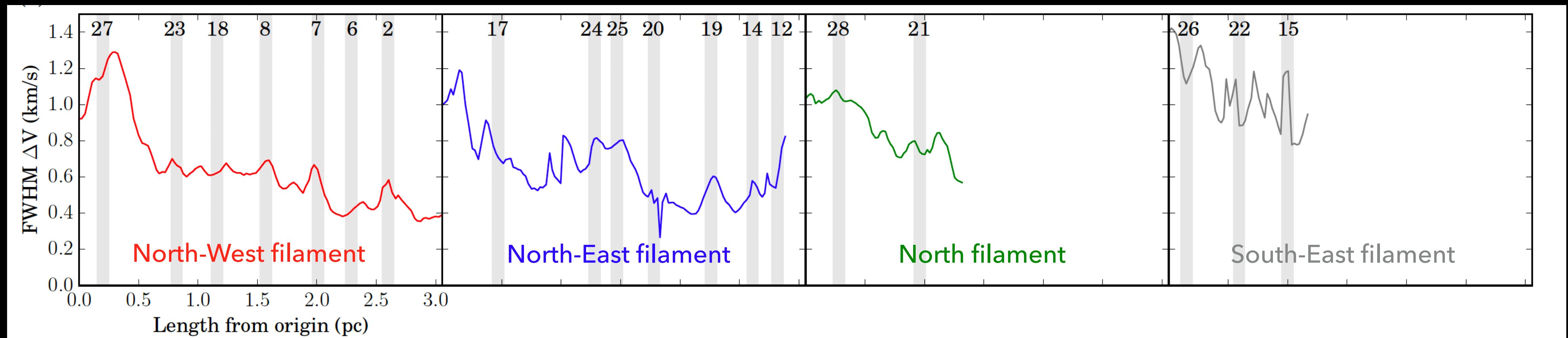
- Dendrogram analysis used to identify structures:



Regular core separation of
 0.37 ± 0.17 pc

Velocity width

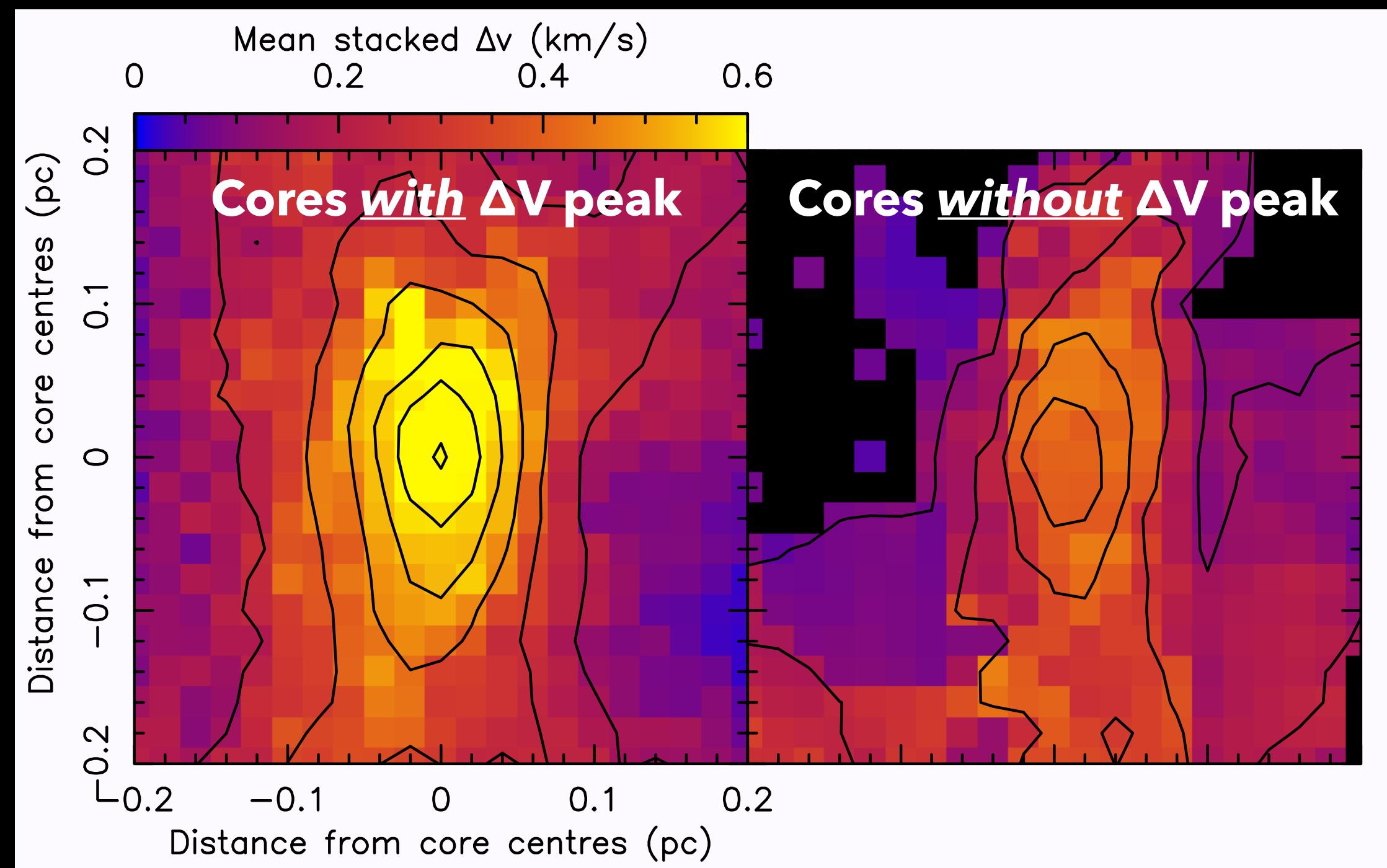
Velocity width profiles along the spines of all four filaments



- Velocity width peaks at the position of two-thirds of starless cores

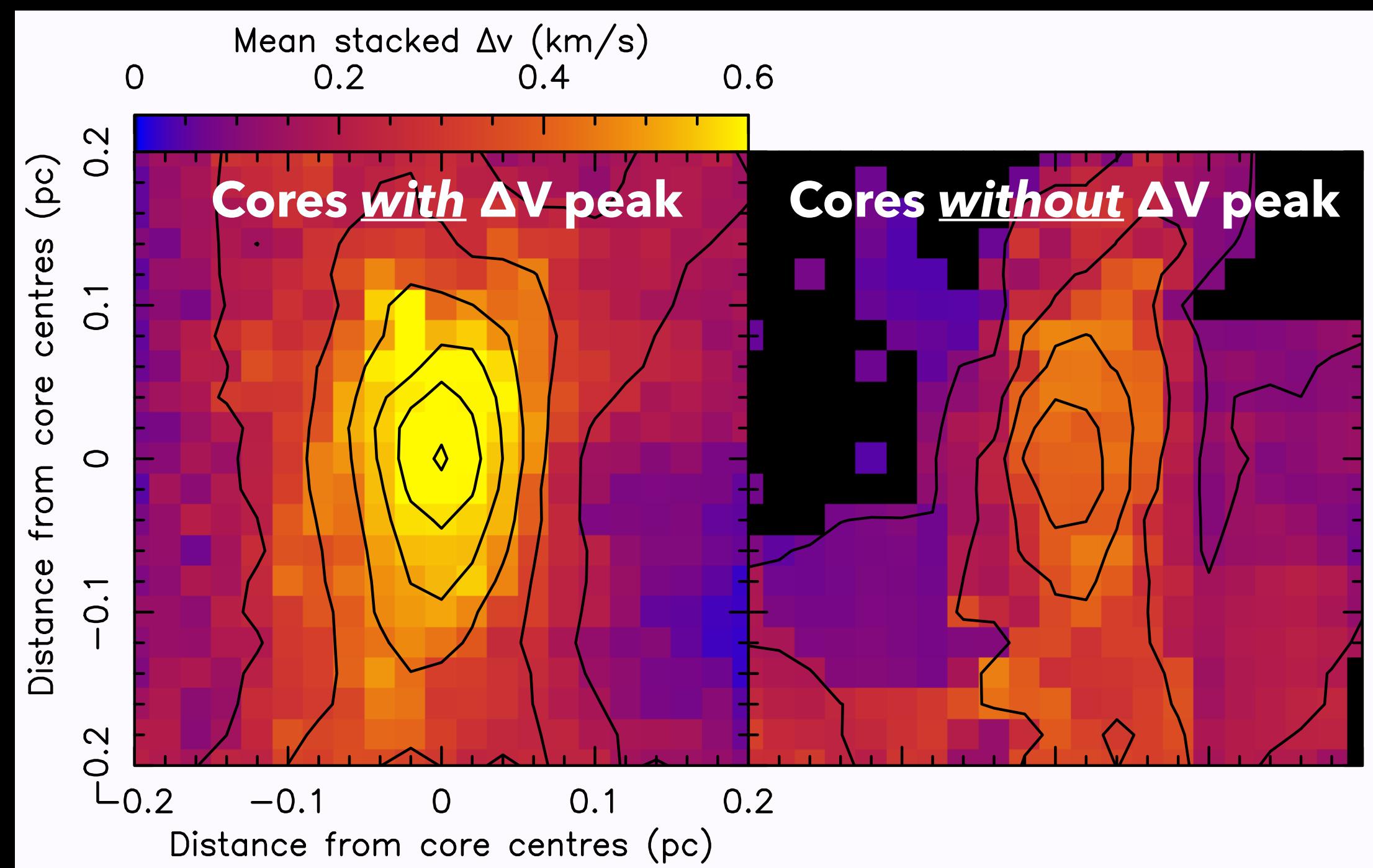
Velocity width

Mean velocity width of the cores

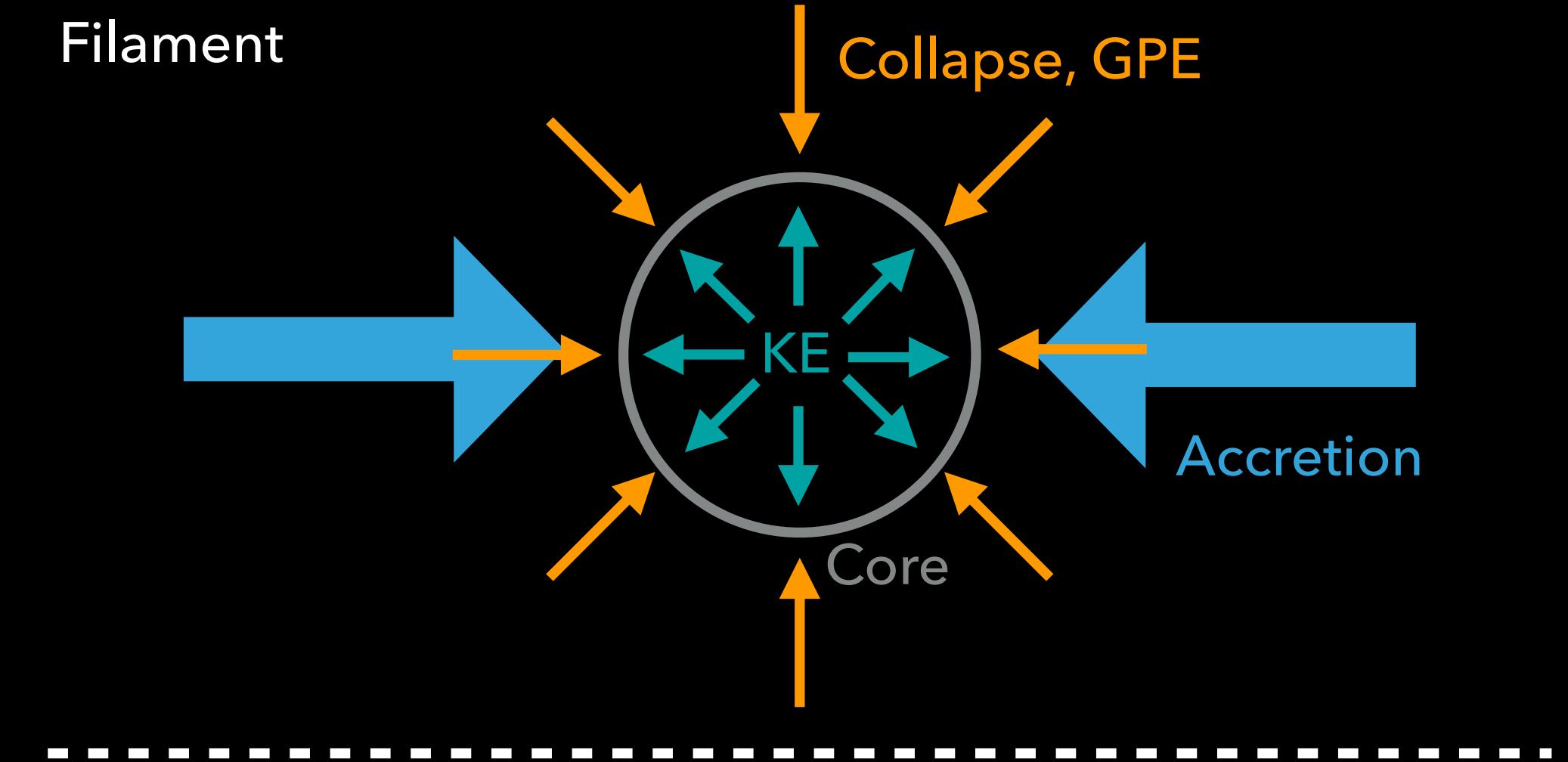


Velocity width

Mean velocity width of the cores



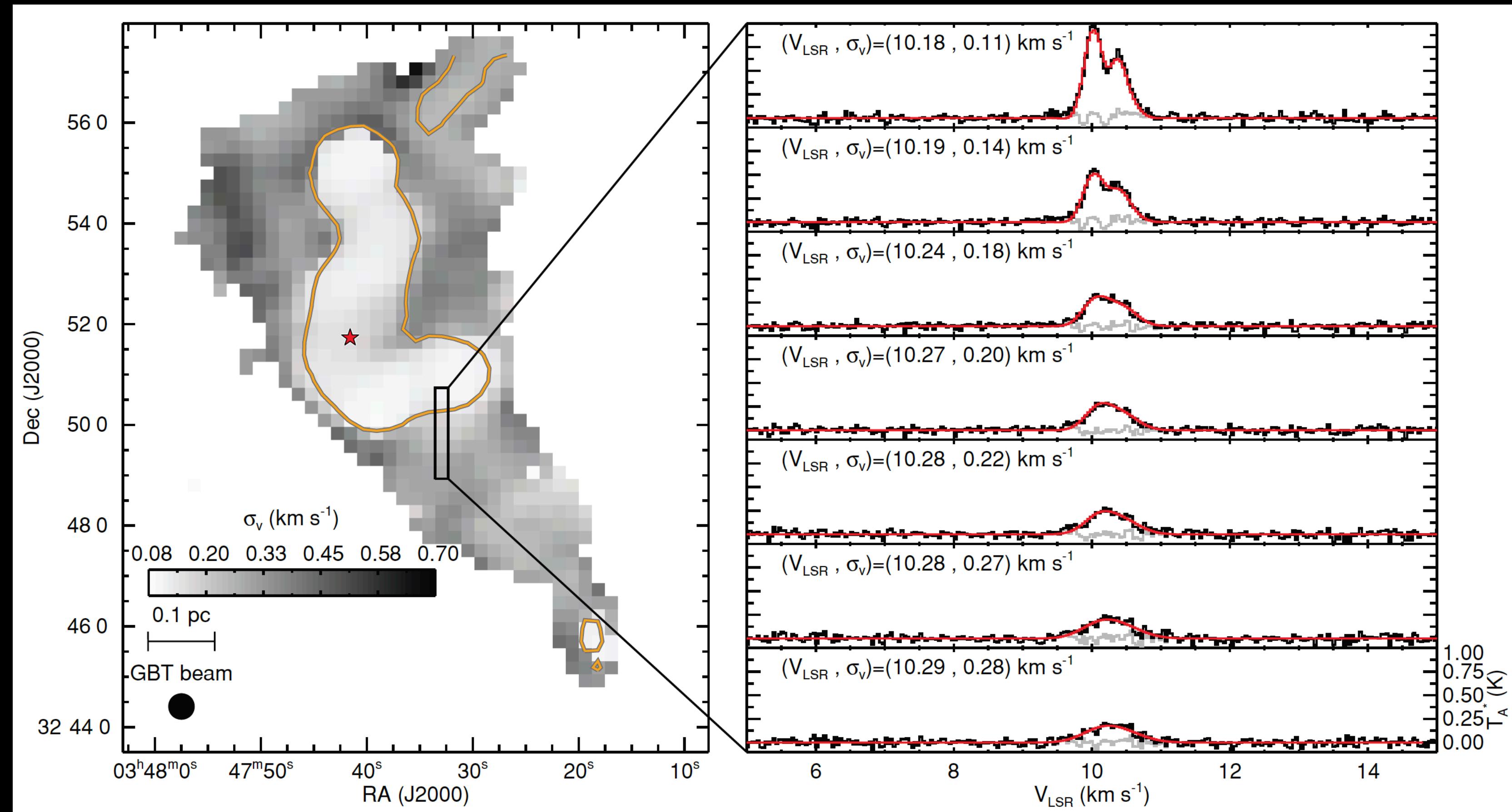
Interpretation:



Extra “turbulent” support of cores, allowing for the assembly of super-Jeans cores

Velocity width

B5 prestellar core in Perseus - Pineda et al. (2010)



Velocity width narrows at the position of this low mass, nearby core

Virial analysis

Given the virial equation:

$$\alpha_{vir} = 2 \frac{E_k}{E_g}$$

- $\alpha_{vir} > 2$, unbound
- $\alpha_{vir} < 2$, gravitationally bound
- $\alpha_{vir} \sim 1$, hydrostatic equilibrium
- $\alpha_{vir} < 1$, gravitationally unstable

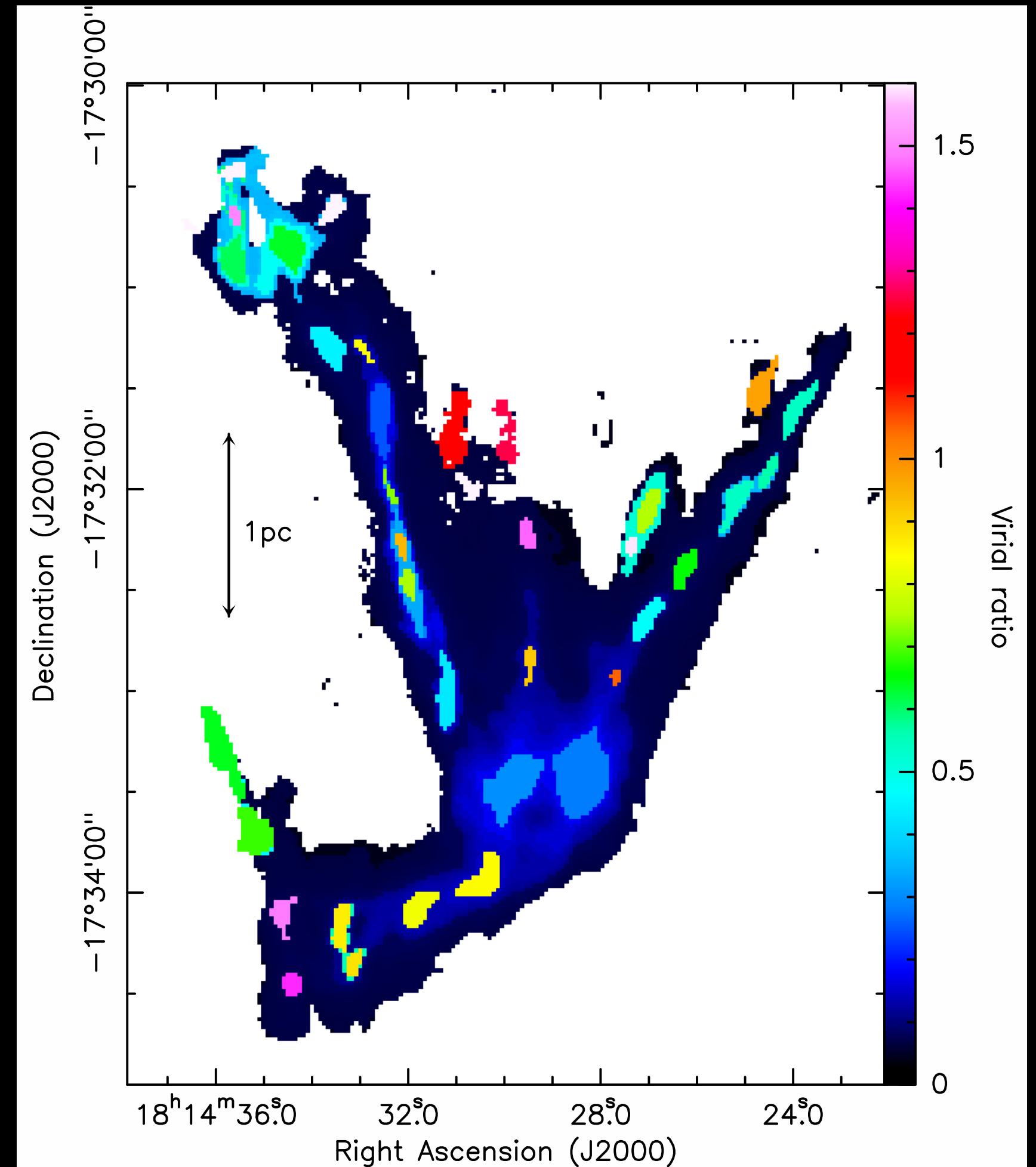
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- $\alpha_{vir} < 1$, gravitationally unstable
- SDC13 appears gravitationally unstable on all scales
- Increase of virial ratio on small, core scales

Map of the virial parameter



Conversion of GPE to KE

Kinetic energy, E_k :

$$E_k = E_{k,0} + \Delta E_k$$

Fractional increase in kinetic energy:

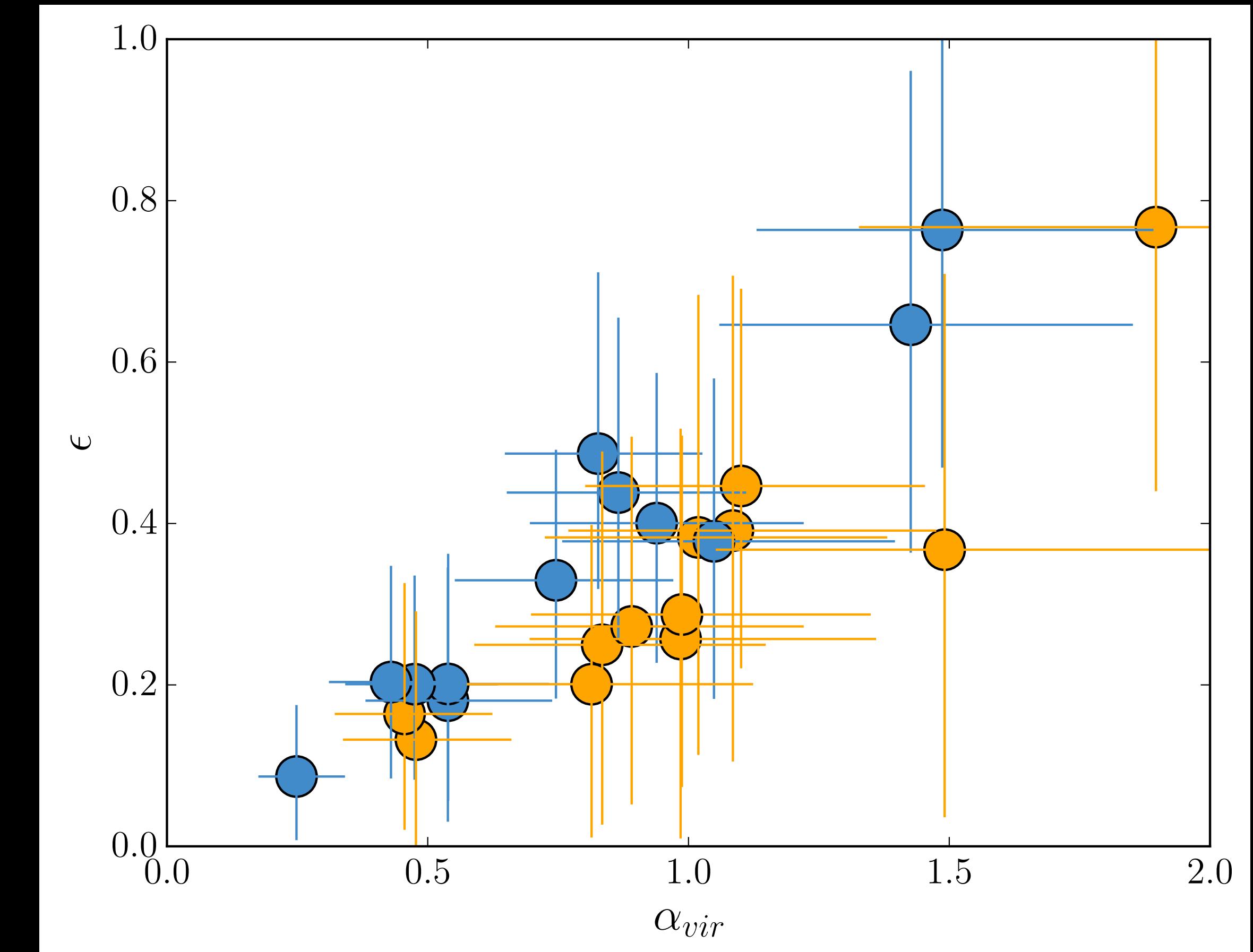
$$\Delta E_k = \epsilon (E_g - E_{g,0})$$

Given the virial equation:

$$\alpha_{vir} = 2 \frac{E_k}{E_g}$$

Substituting and rearranging gives:

$$\alpha_{vir} = 2 \frac{E_{k,0}}{|E_g|} + 2\epsilon \left(1 - \frac{|E_{g,0}|}{|E_g|} \right)$$



● = JVLA+GBT data, ● = JVLA only data.
 $\epsilon_{\text{mean}} = 35\%$

Conclusions

- A new high-angular/spectral resolution study
- Regularly spaced cores, fragmented along the filaments
- Velocity width broadens at the positions of starless cores
- SDC13 is gravitationally unstable on all scales
- Conversion of GPE into KE is on average 35% efficient in cores

