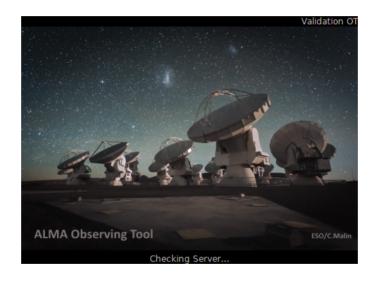
# The ALMA Observing Tool (OT)

(or the essential tool to prepare and submit a proposal for ALMA)



Ana Karla Díaz-Rodríguez





### **ALMA** timeline

- March: Call for proposals (CfP)
- April: Proposals due
- August: Grades for proposals are announced
- September: Phase 2 of proposal submission (review of the Scheduling Blocks); end of observations from the previous cycle; ACA supplemental CfP
- October: Observations start for new cycle; ACA proposals due

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   OT (Pl
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   (review of the Scheduling Blocks); end of observations
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   OT (if proposal accepted)

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March: Call for proposals (Cff

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October: Observations start for new cycle; ACA proposals due

Shifted 1 year

**COVID**-19

## Cycle 8 2021 ALMA timeline

- March 17: Call for proposals (CfP)
- April 21: Proposals due
- August: Grades for proposals are announced
- September 8: ACA supplemental CfP
- October 1: Observations start for Cycle 8 and continue with remainder of Cycle 7
- October 6: ACA proposals due

# Observing Tool (OT)

- Java desktop application (c8: own version of Java 11)
- Updated each cycle to match ALMA capabilities
- Works on Linux, Mac, Windows

Download and installing: ALMA OT Installer (recommended) or tarball (if anything goes wrong with the installer :( )

Documentation: OT Quickstart Guide, OT video tutorials, User Manual

### Lets propose!

#### **ALMA** proposal

Science justification + Technical Justification

Why?

4 or 6 pages

PDF file

(templates)

How?

OT

## Lets propose!

#### **ALMA** proposal

Science justification + Technical Justification

Why?

How?

**Science Goals** contain the technical specification of your project. Are limited to one correlator setup, one calibration strategy, and one set of Control and Performance parameters.

- Band 3 observations of one or more targets (1 SG)
- Band 3 observations at 0.02" and 4" (2 SG)
- Band 3, Band 4 and Band 5 observations (3 SG)

## Now, lets propose!

Example 1: Continuum observations of several targets (e.g. survey of protostars)

Example 2: Line observations of single target (e.g. study kinematics of a protostellar disk)

Example 3: Mosaic observations (e.g. mapping a starforming filament)

## Now, lets propose!

Example 1: Continuum observations of several targets (e.g. survey of protostars)

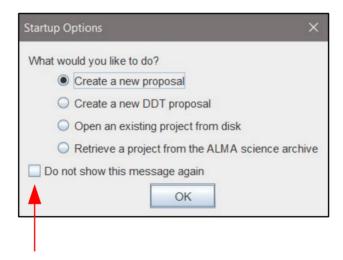
Example 2: Line observations of single target (e.g. study kinematics of a protostellar disk)

Example 3: Mosaic observations (e.g. mapping a starforming filament)

O. Register with the ALMA Science Portal (PI and co-Is)

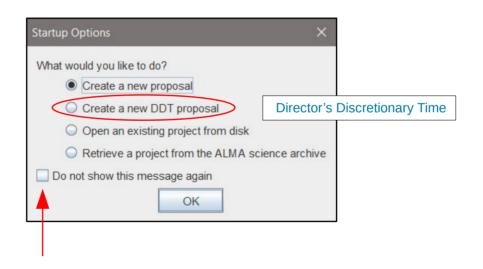
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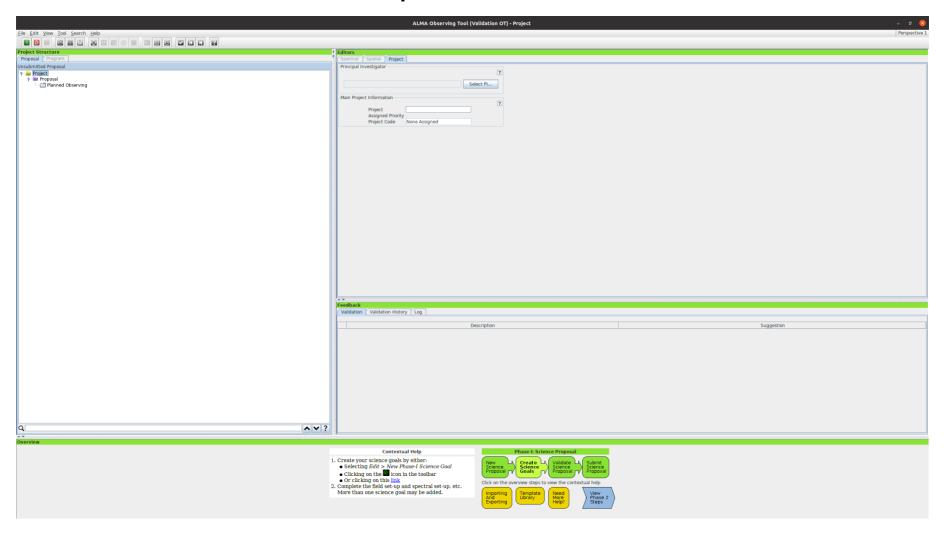
Mac (double-click): ALMA-OT.app

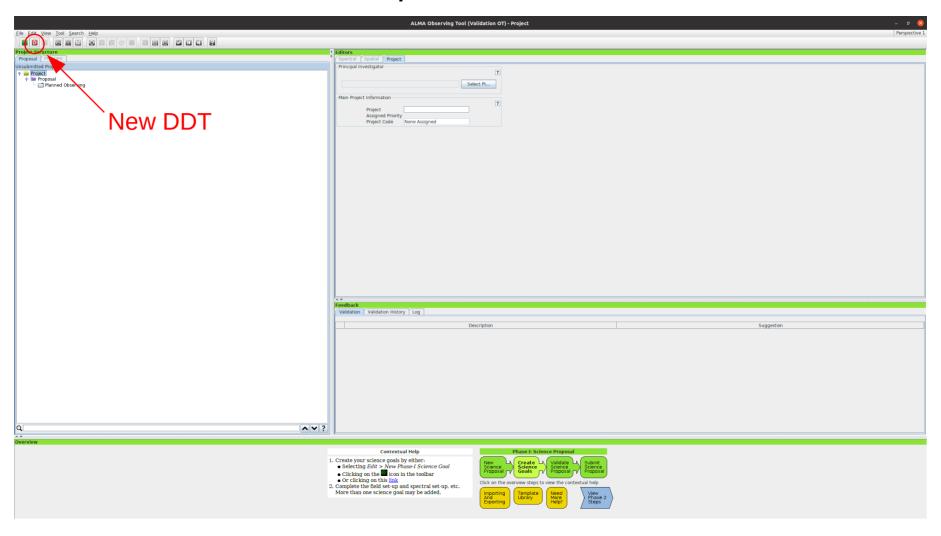


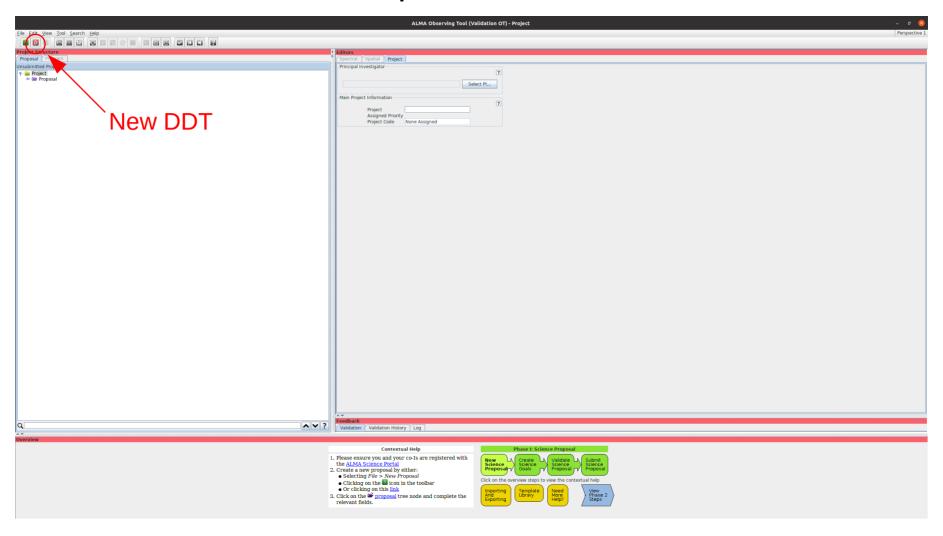
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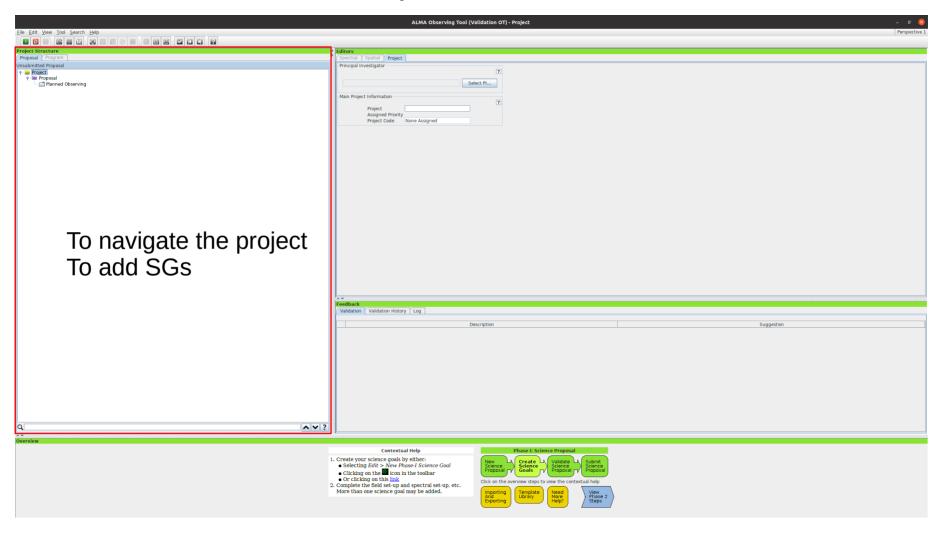
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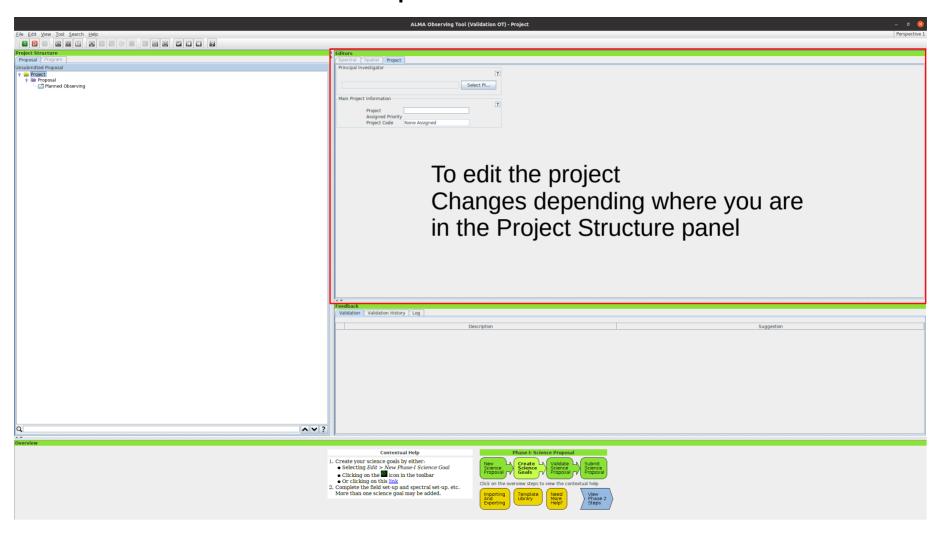


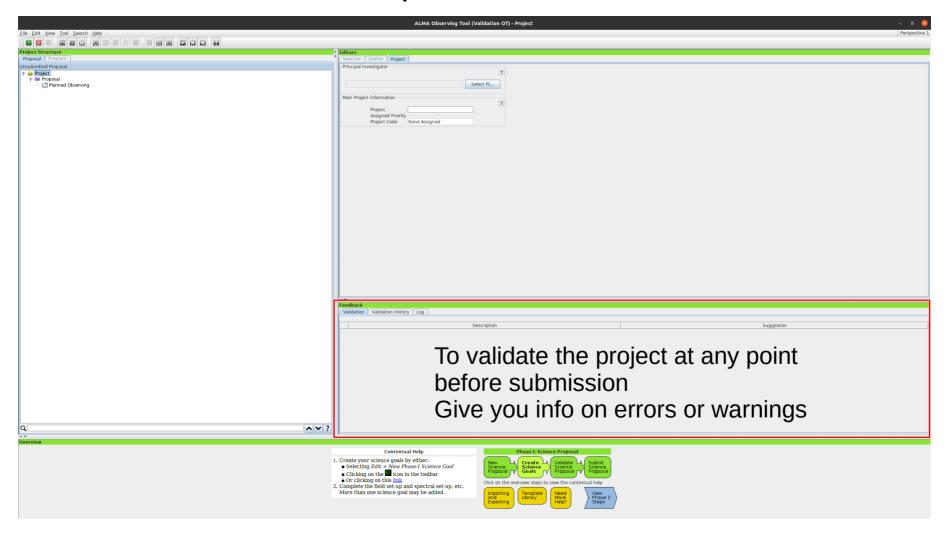


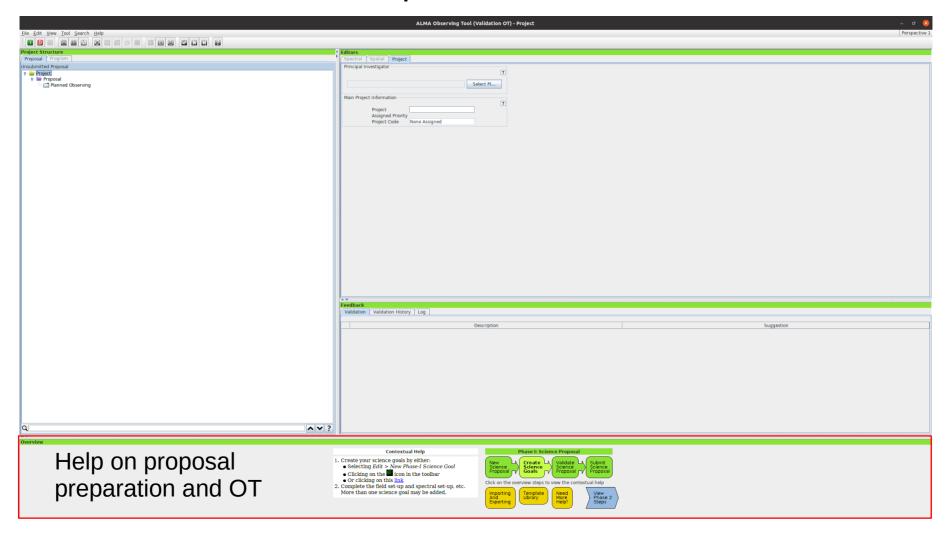


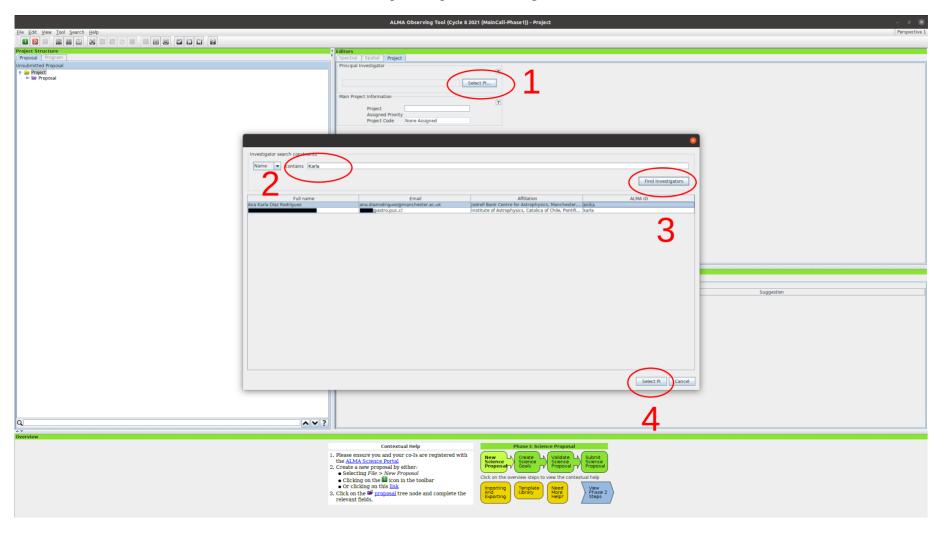


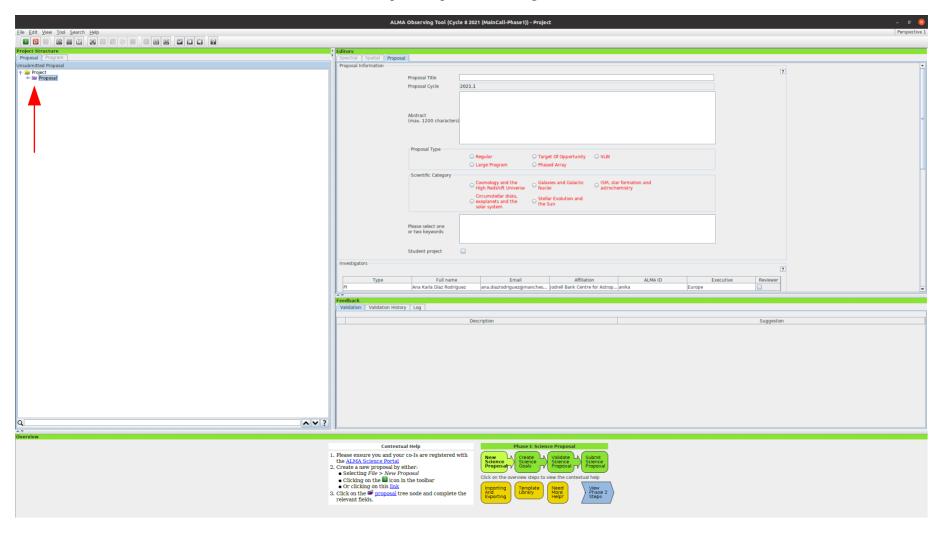


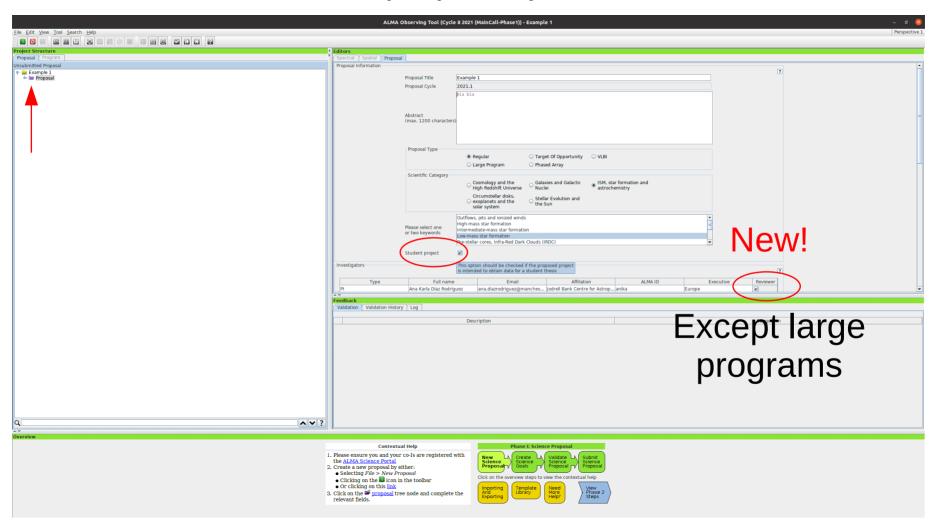


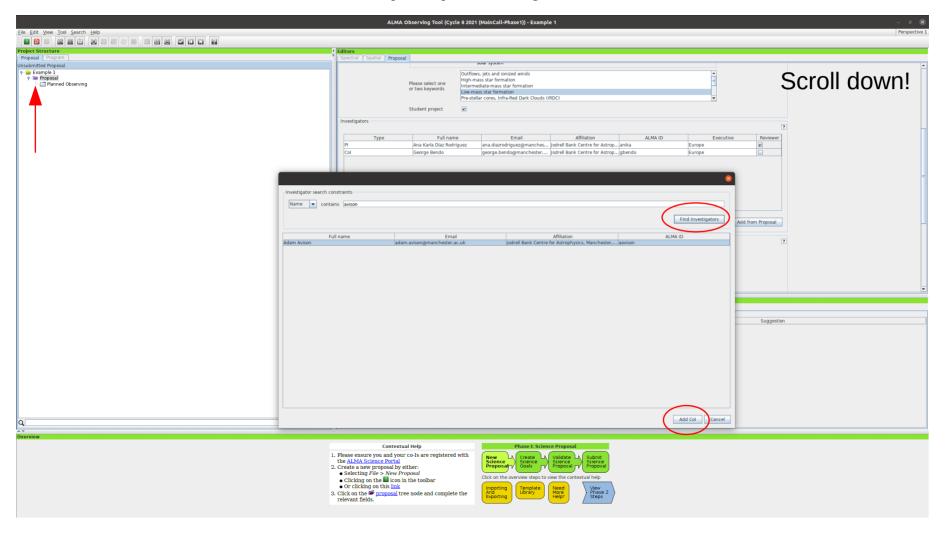


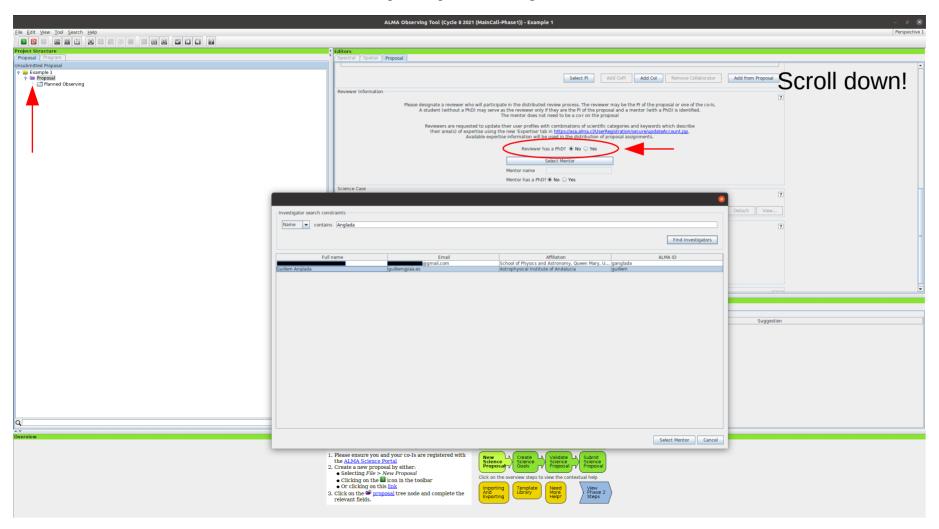


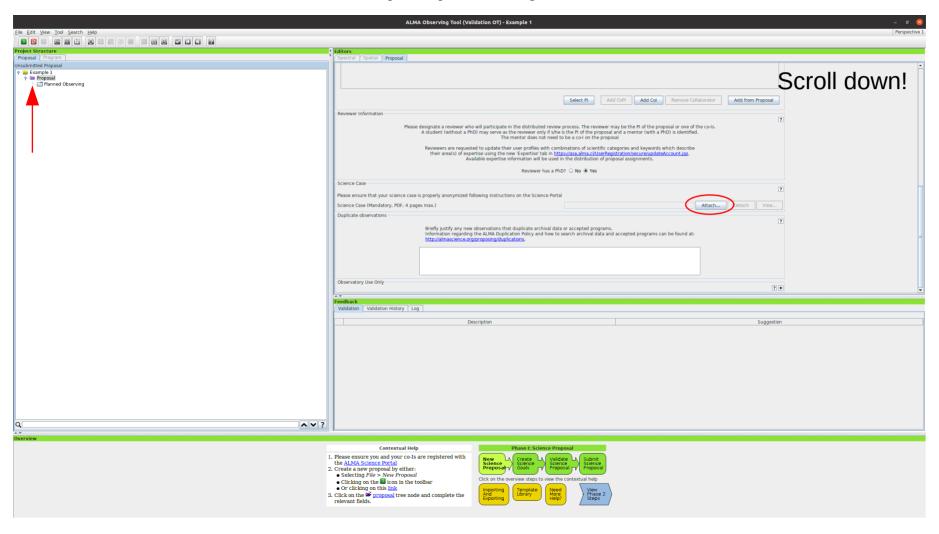


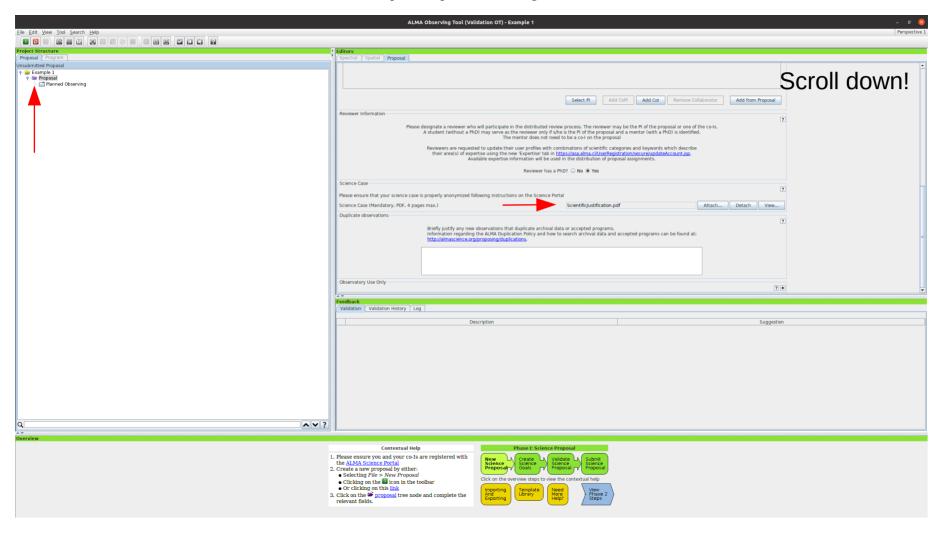


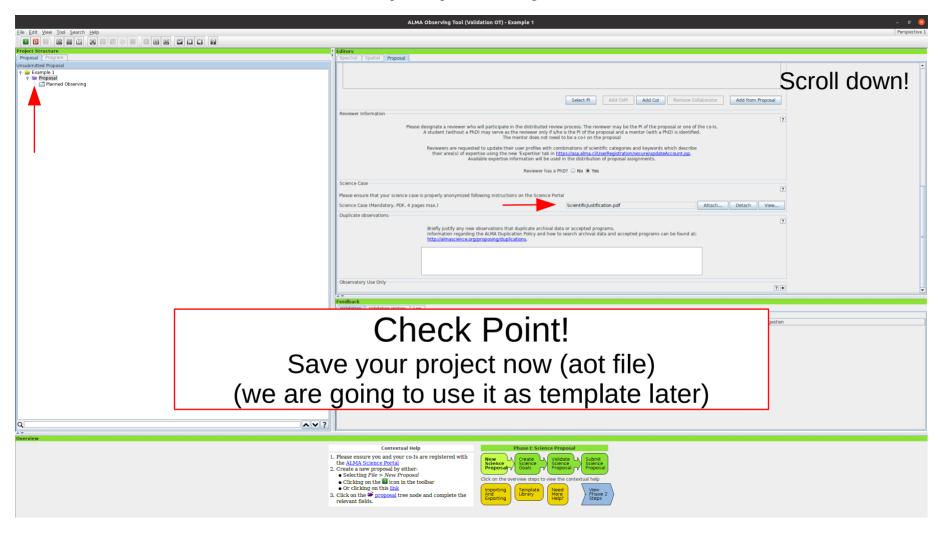








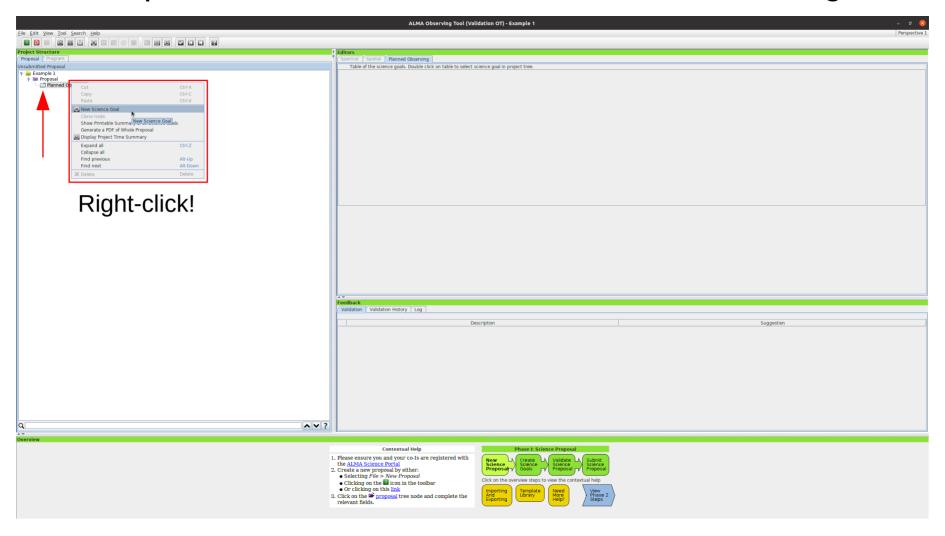


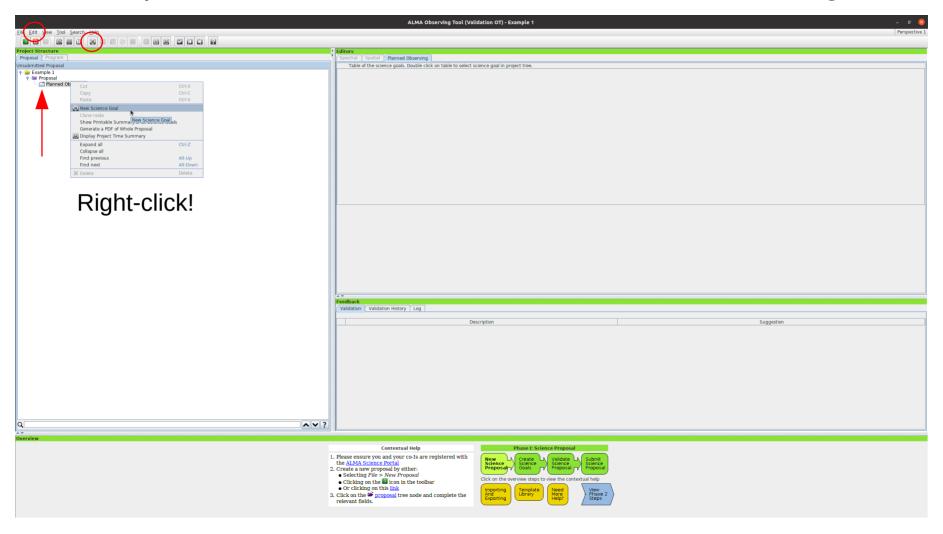


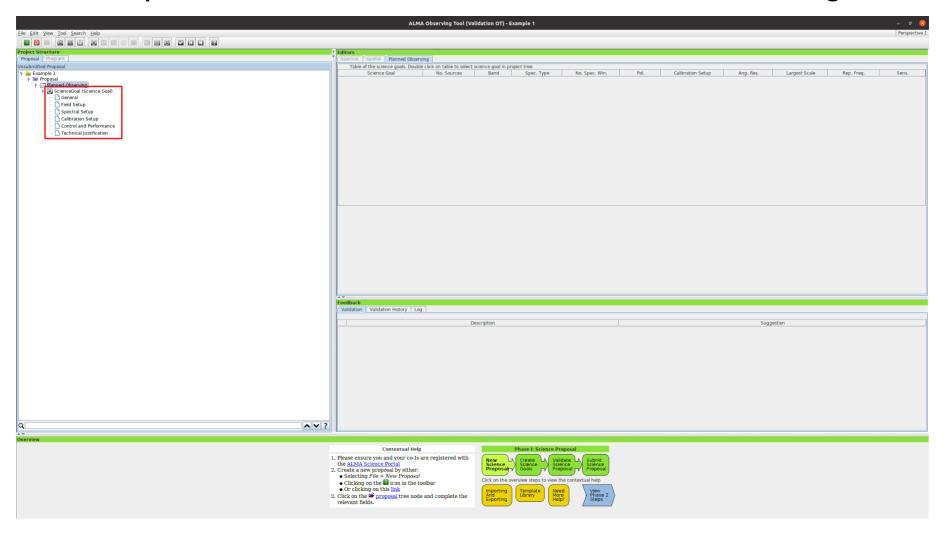
### Now, lets propose!

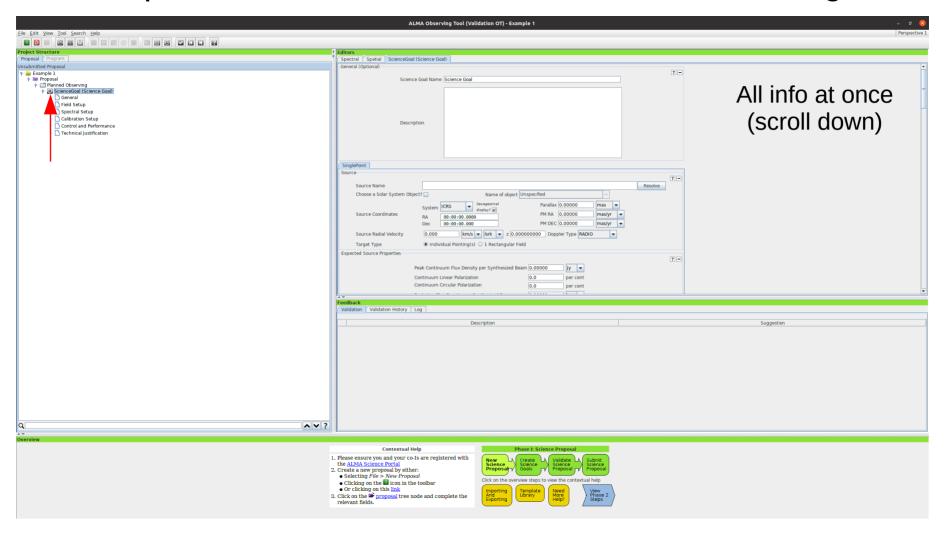
Example 1: Continuum observations of several targets (e.g. survey of protostars)

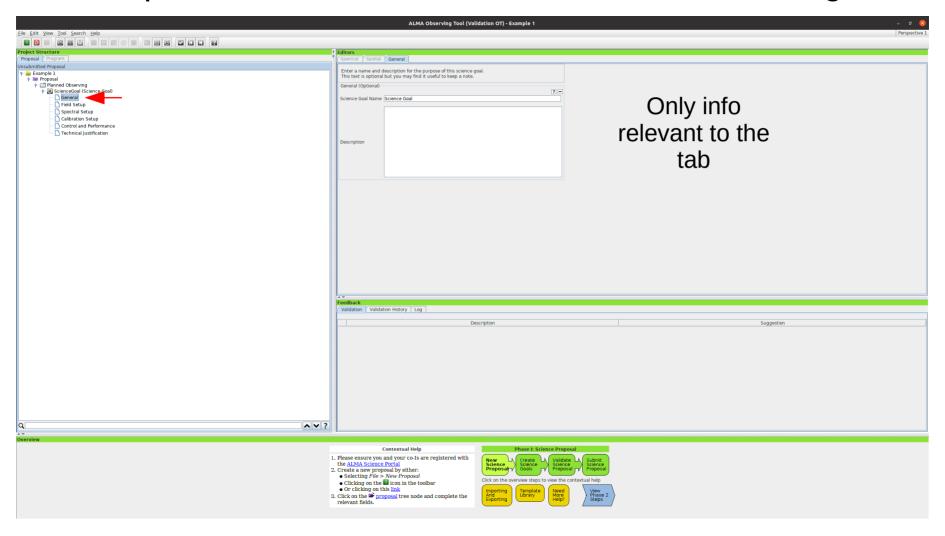
1 Science Goal with multiple targets

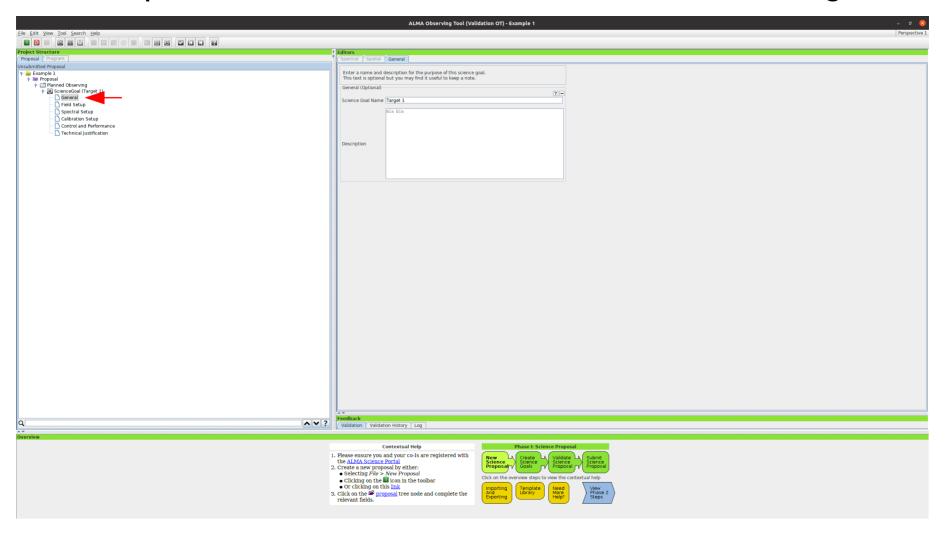


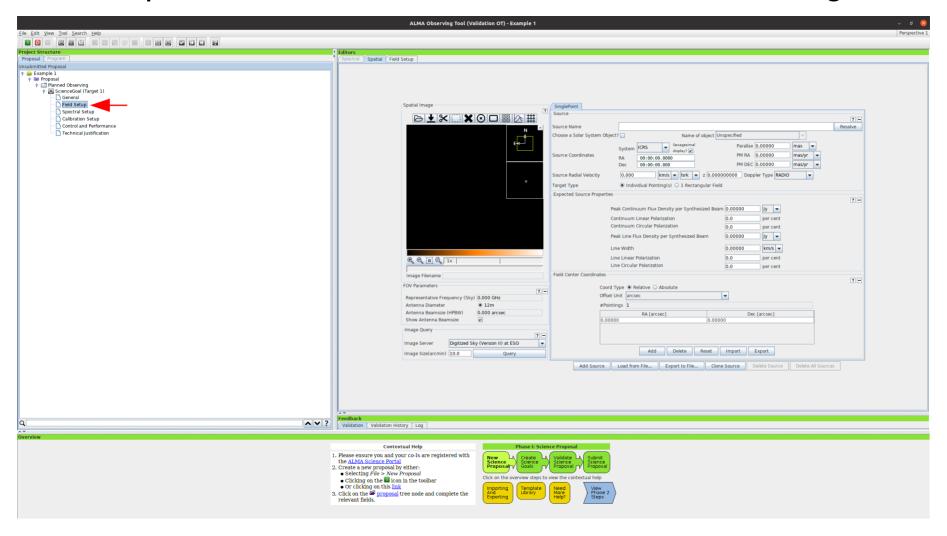


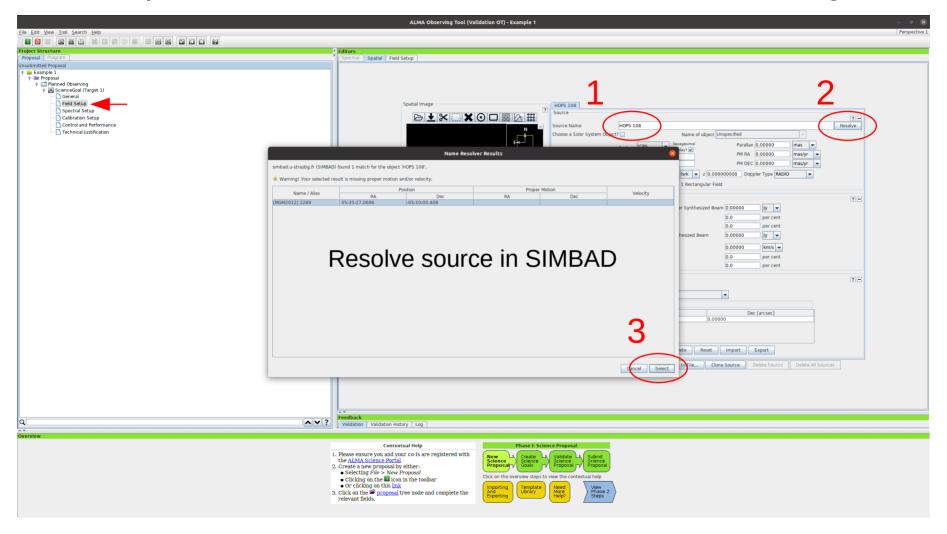


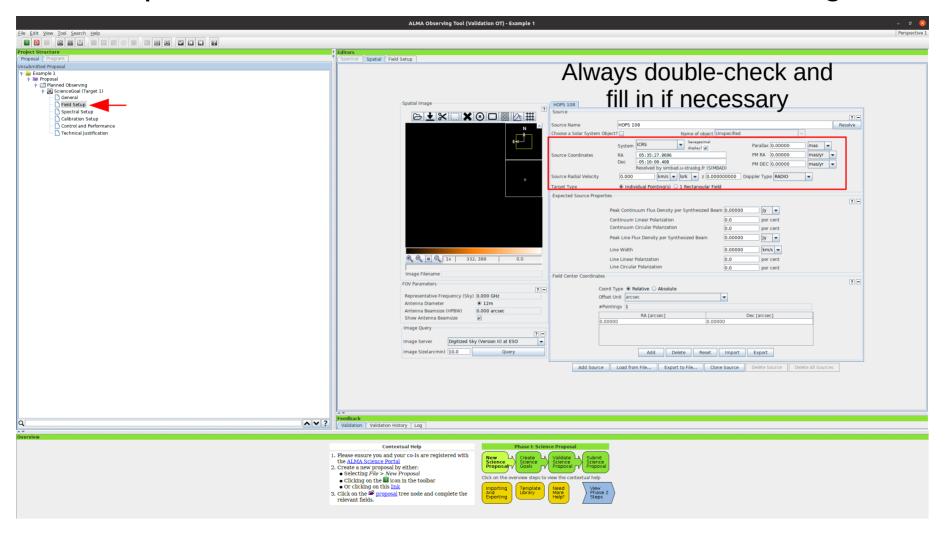


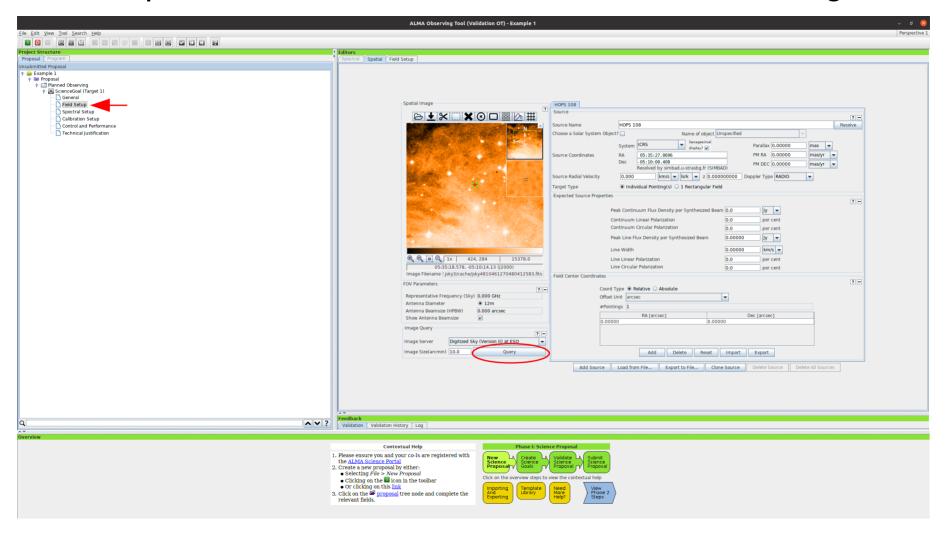


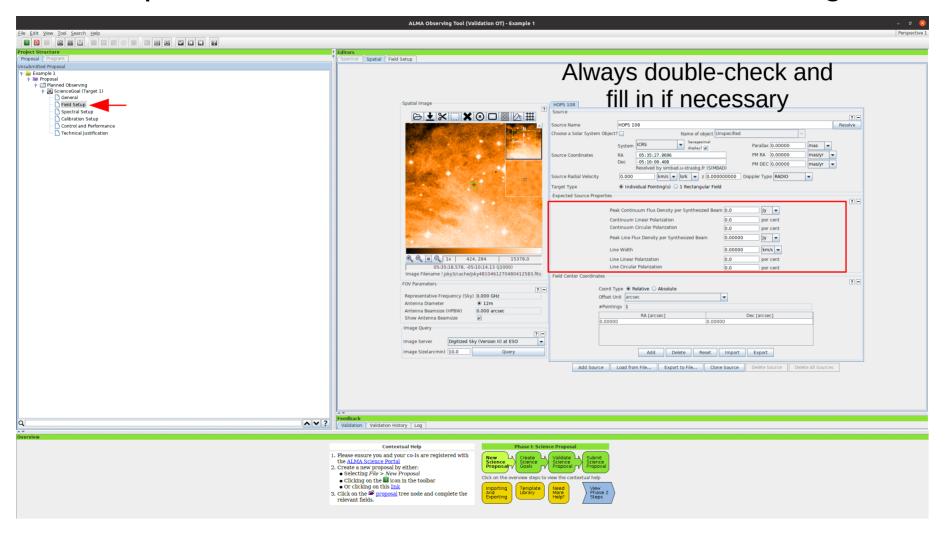


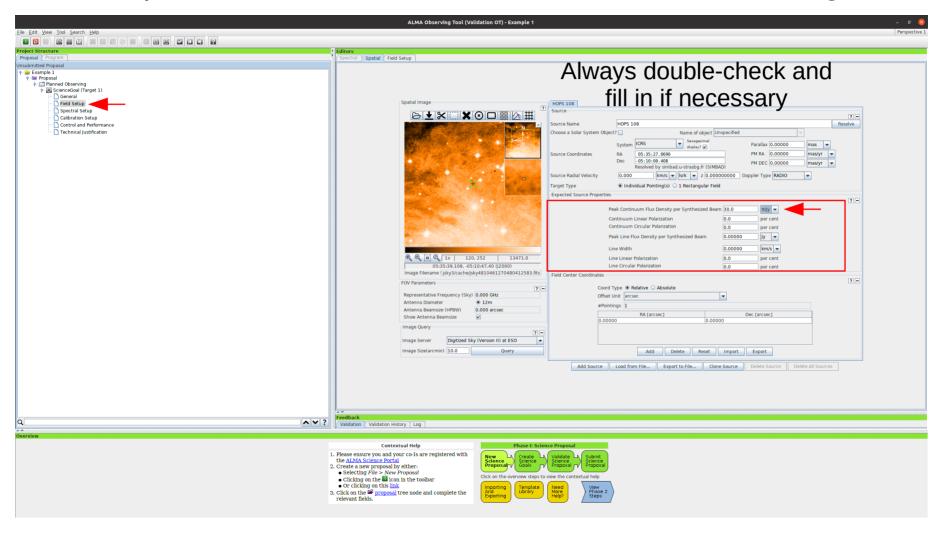


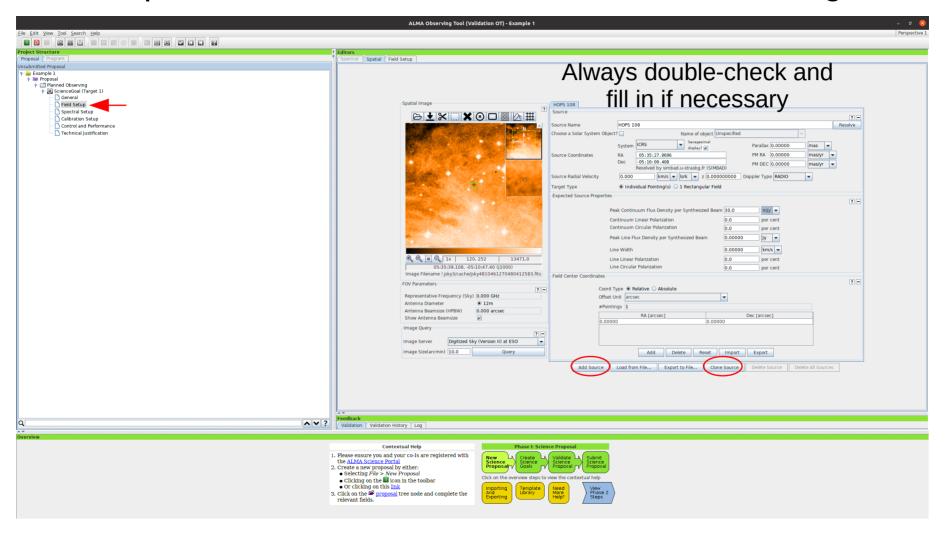


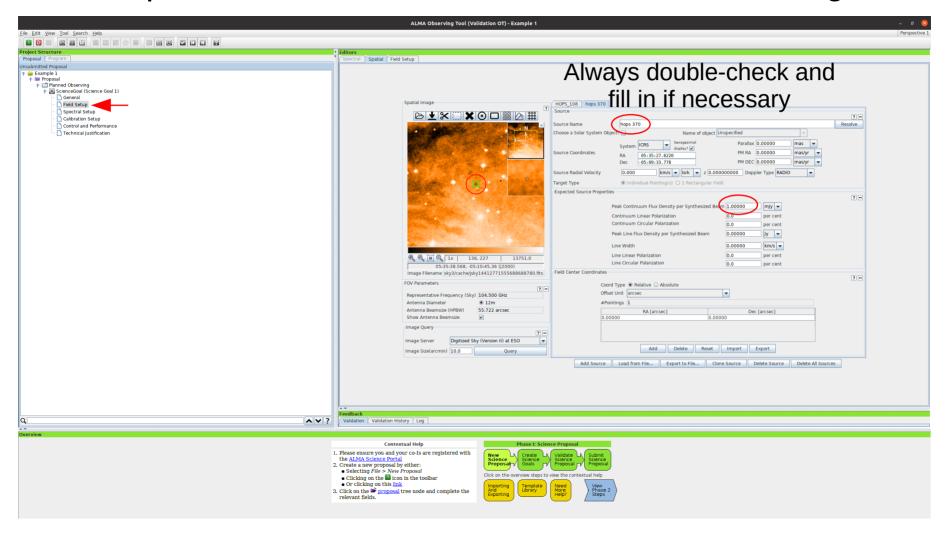


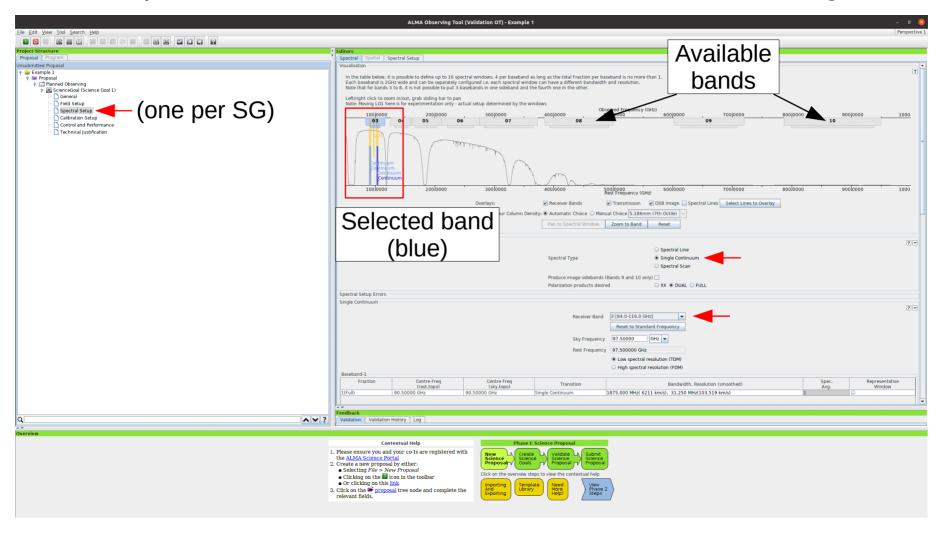


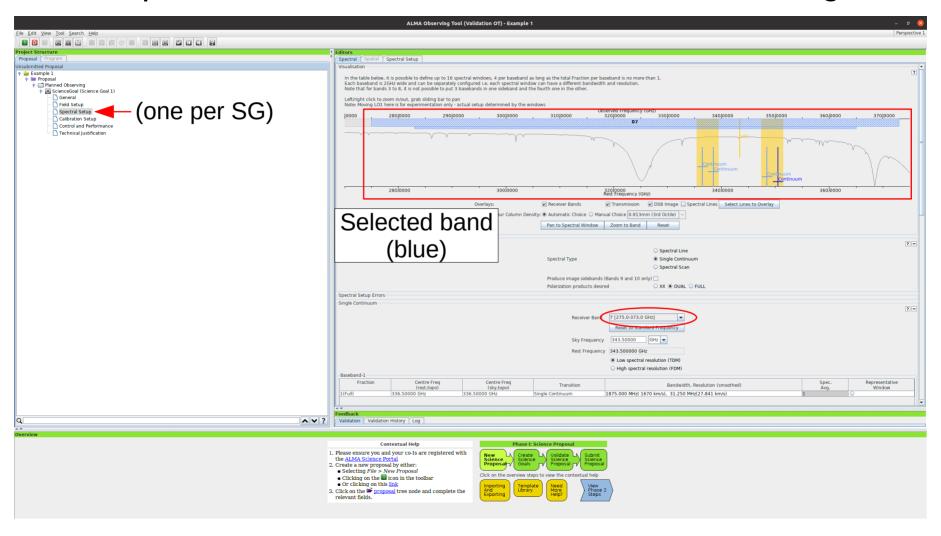


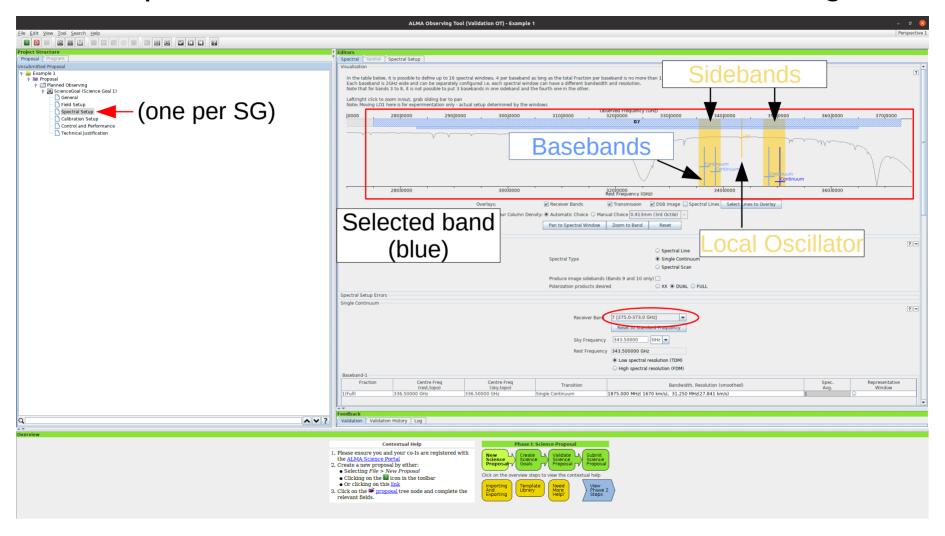


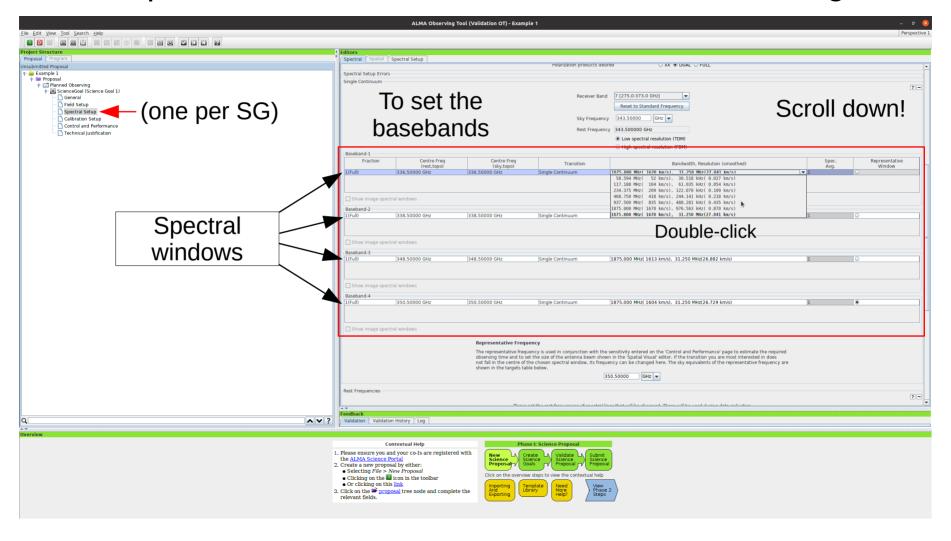


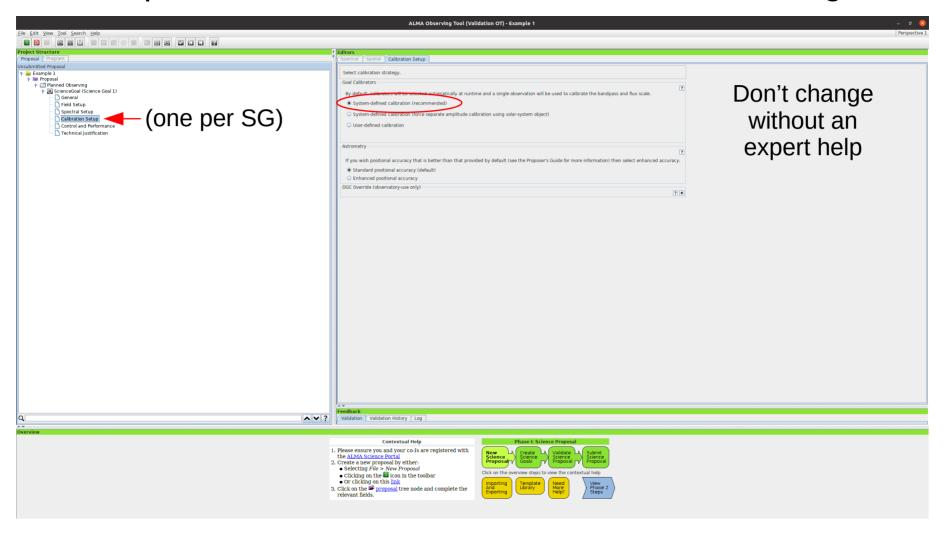


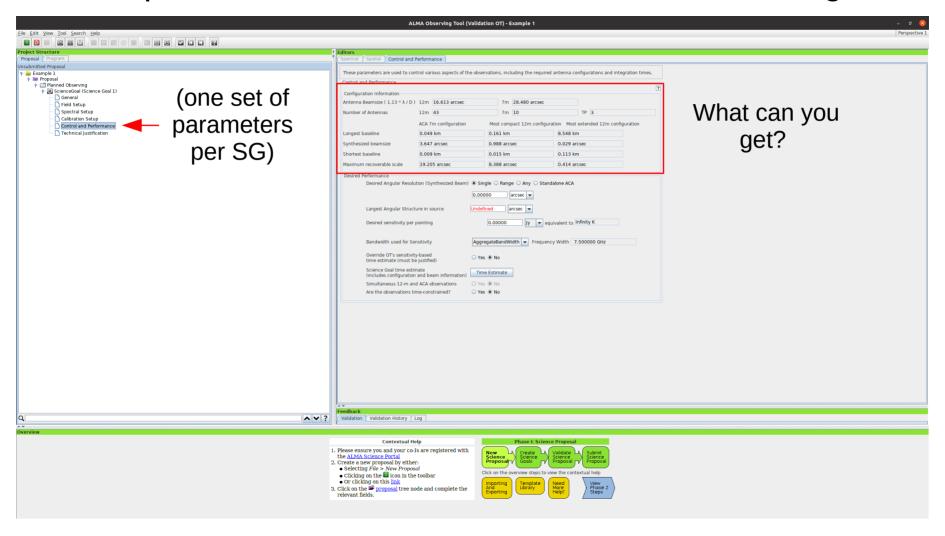


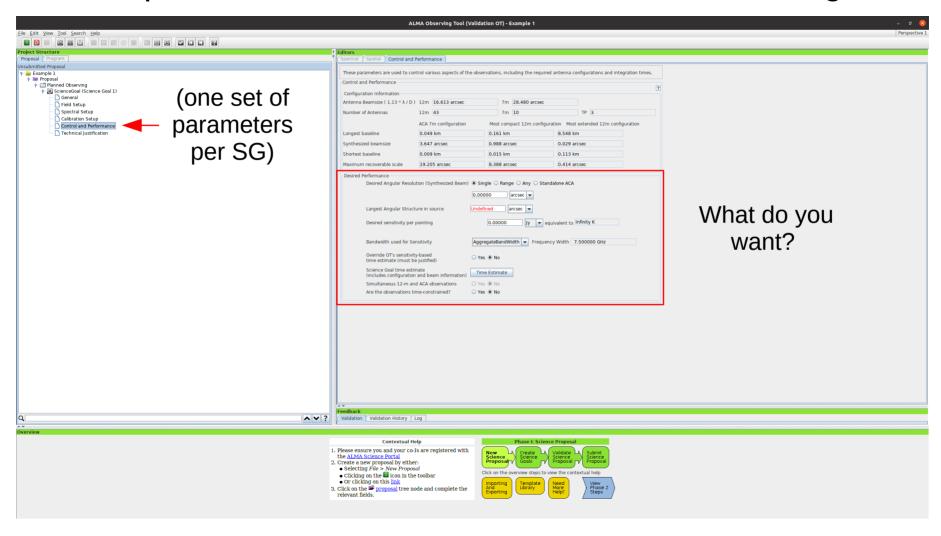


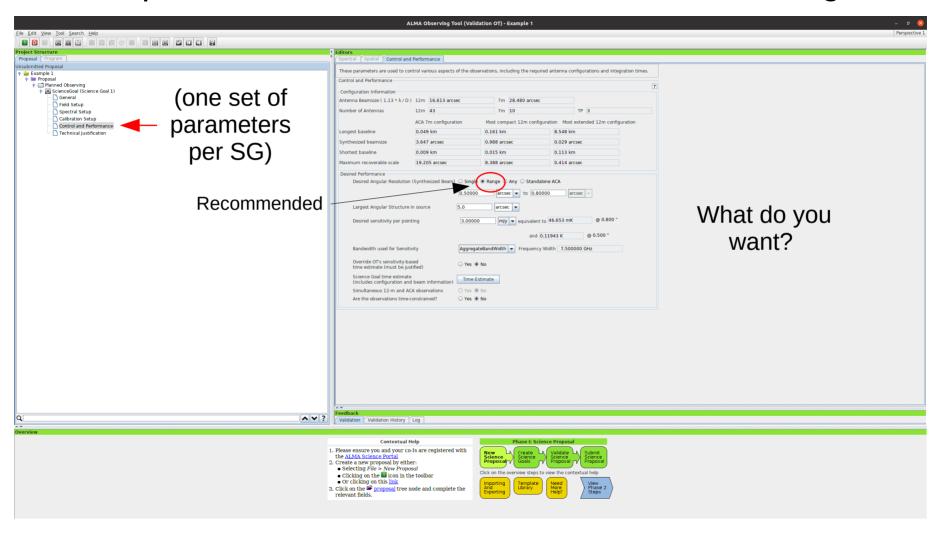


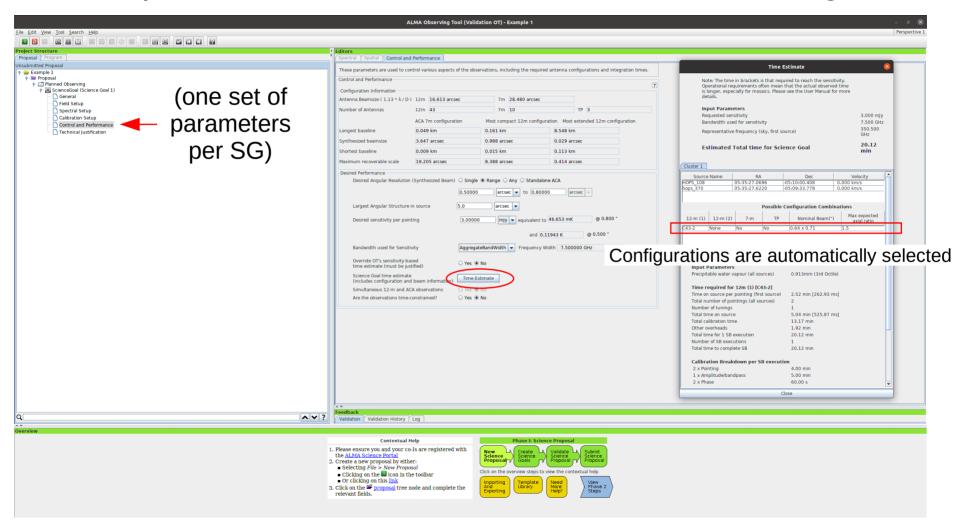


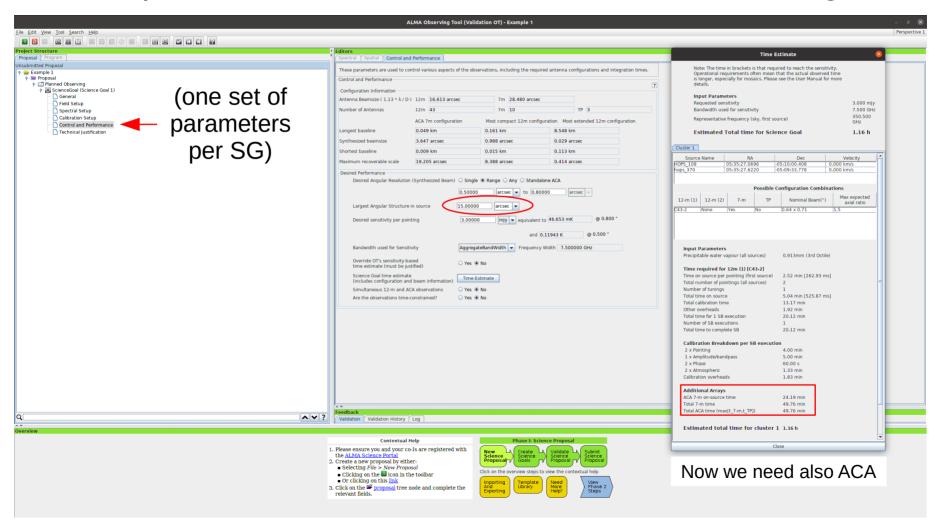


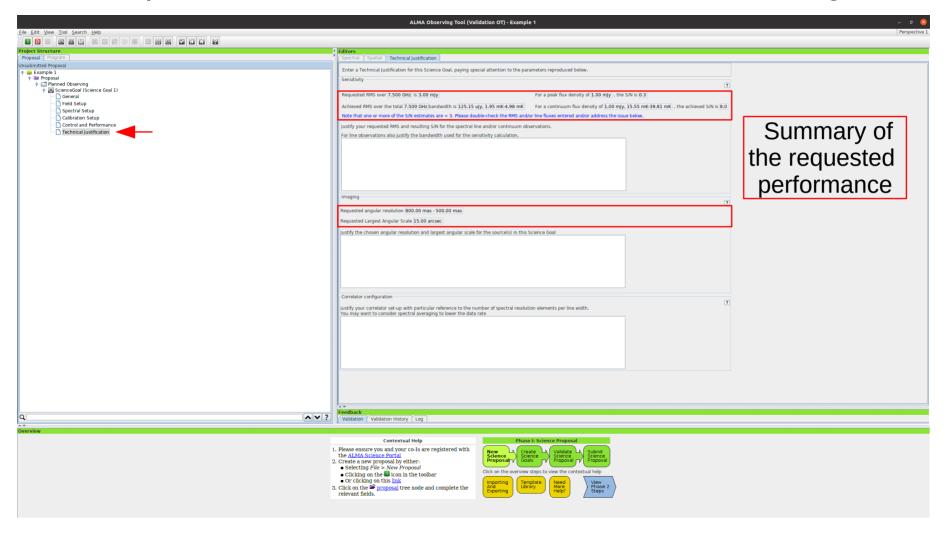


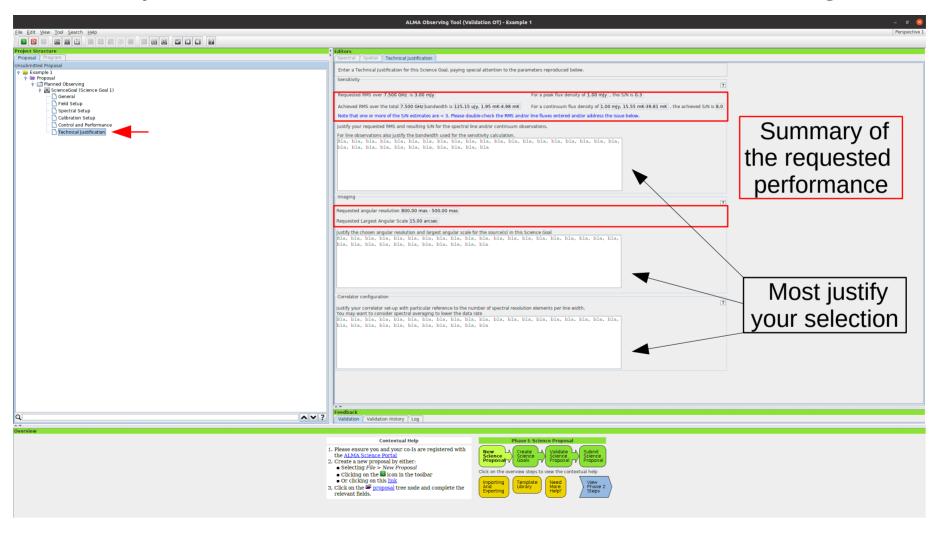


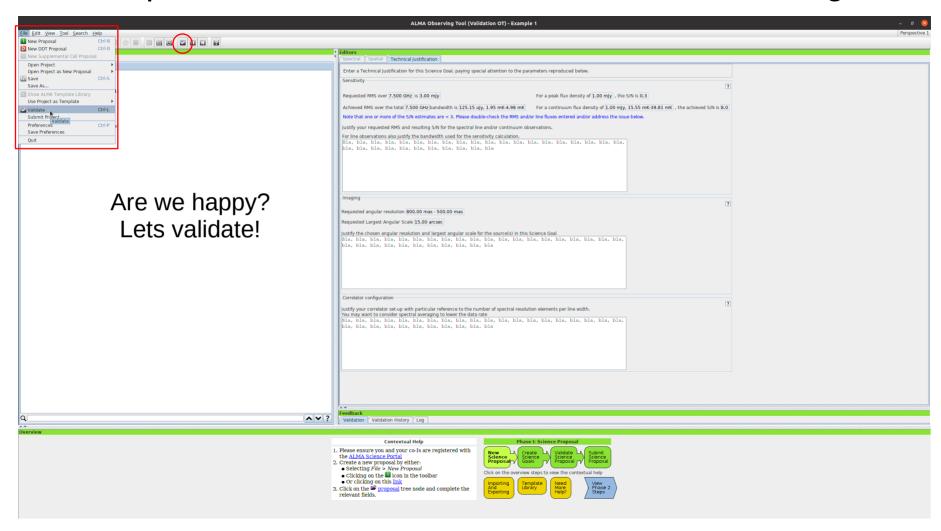


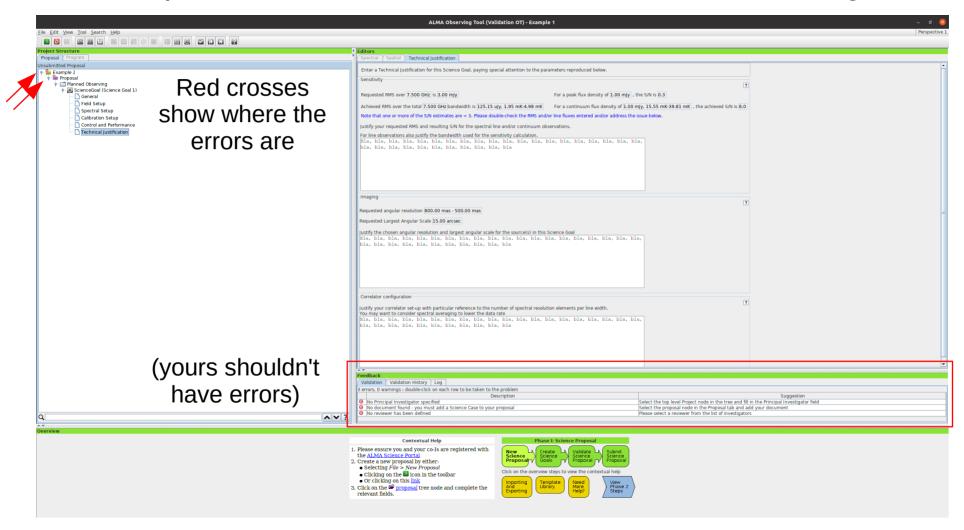


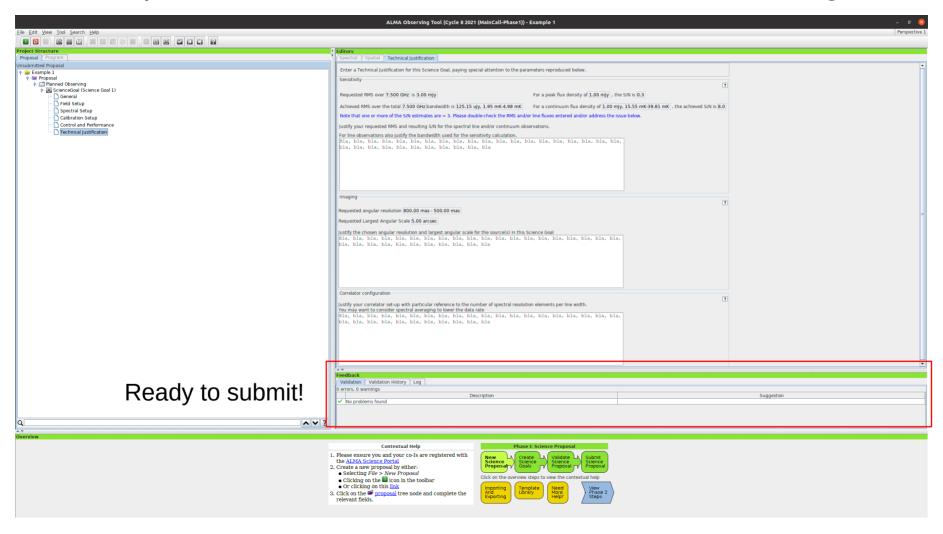


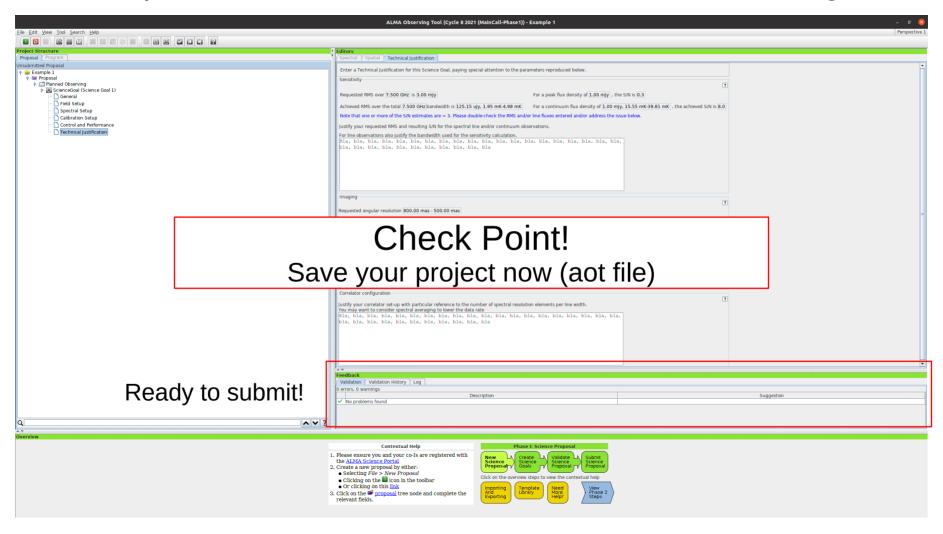


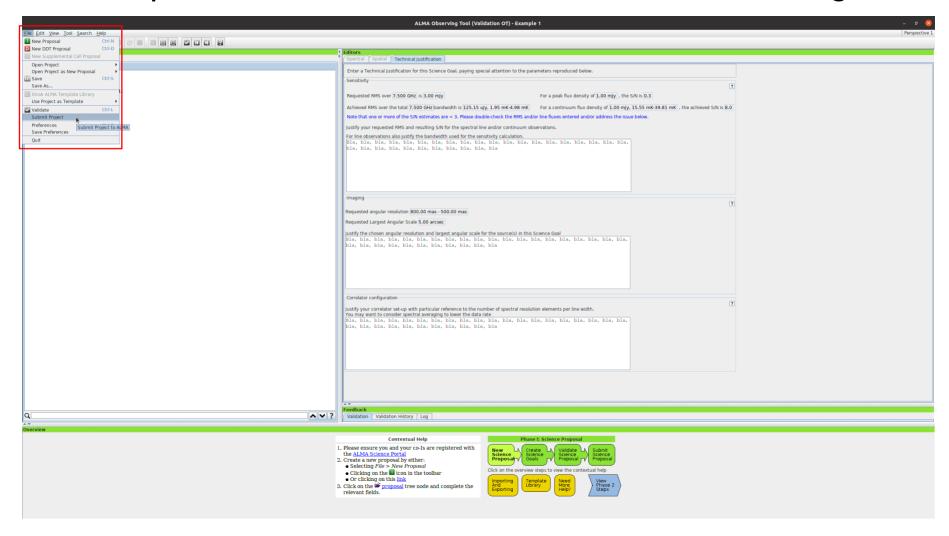


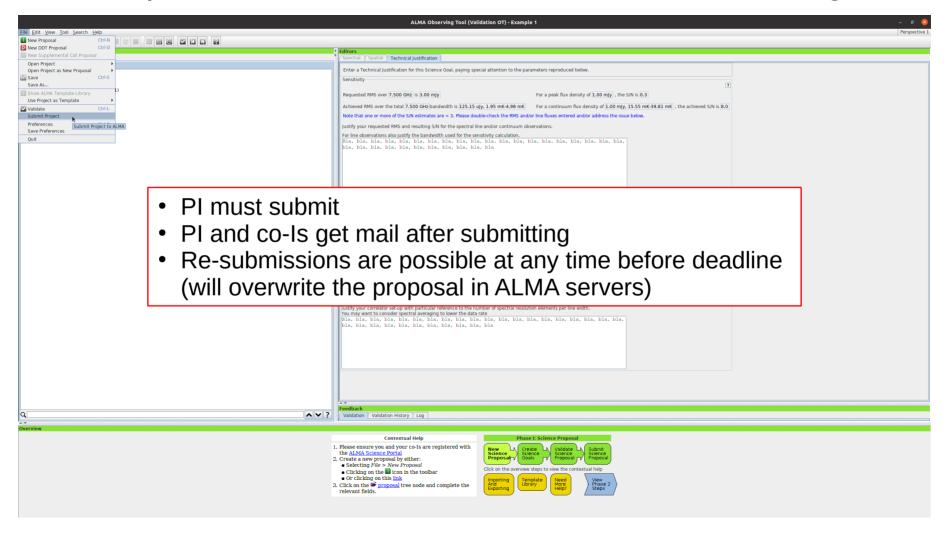


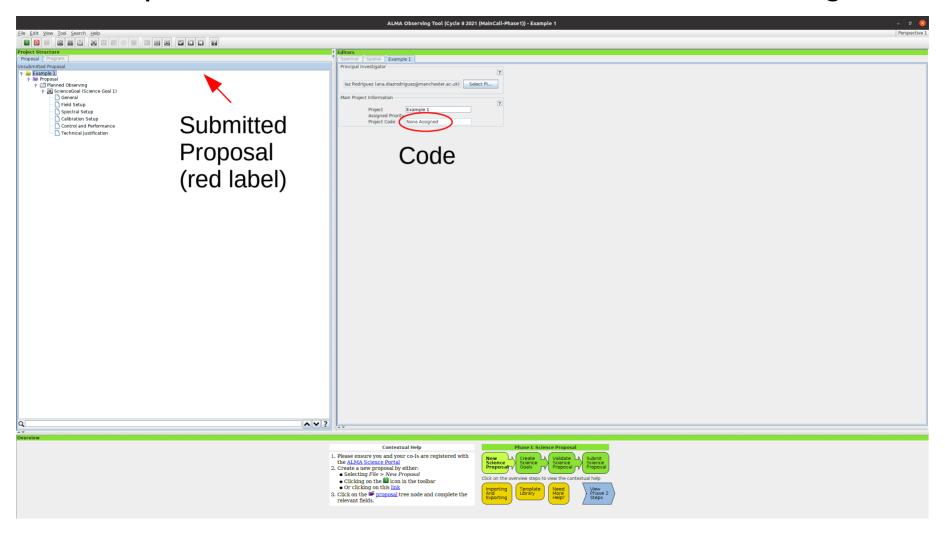


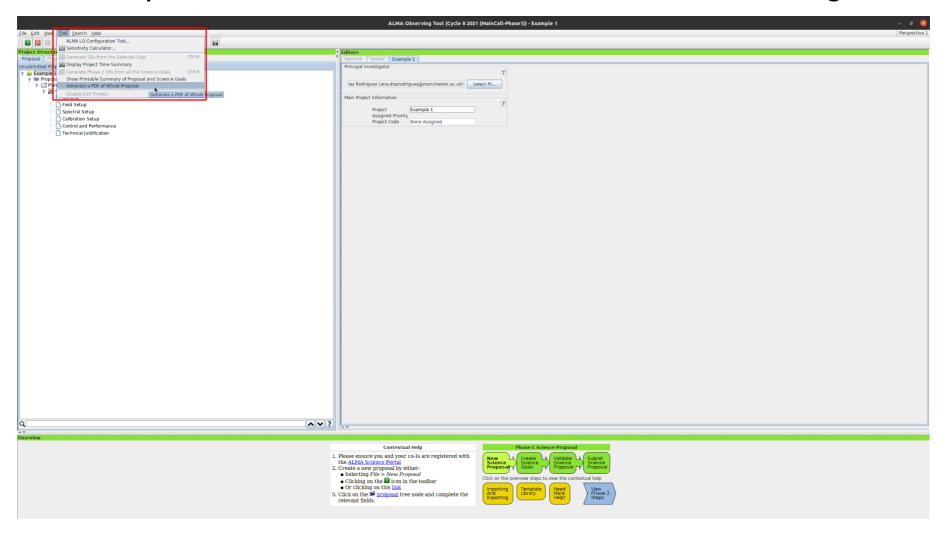


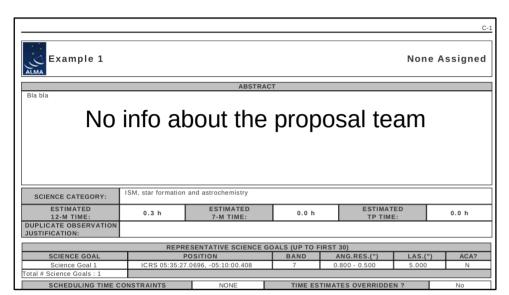


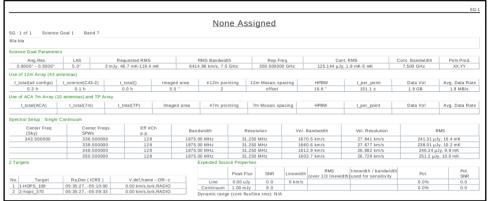






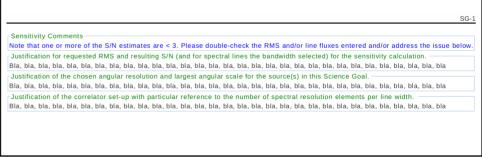






Share PDF (and aot file) with Co-PIs





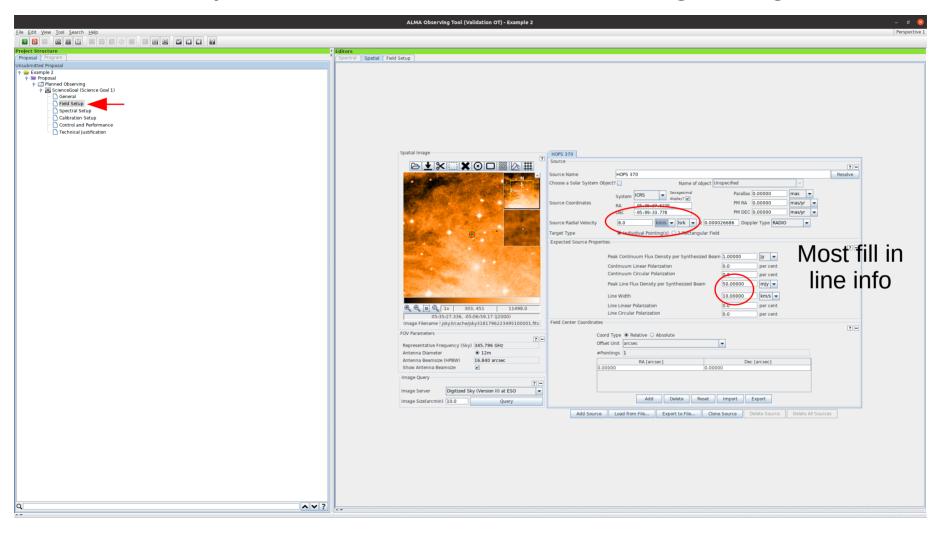
# Lets propose!

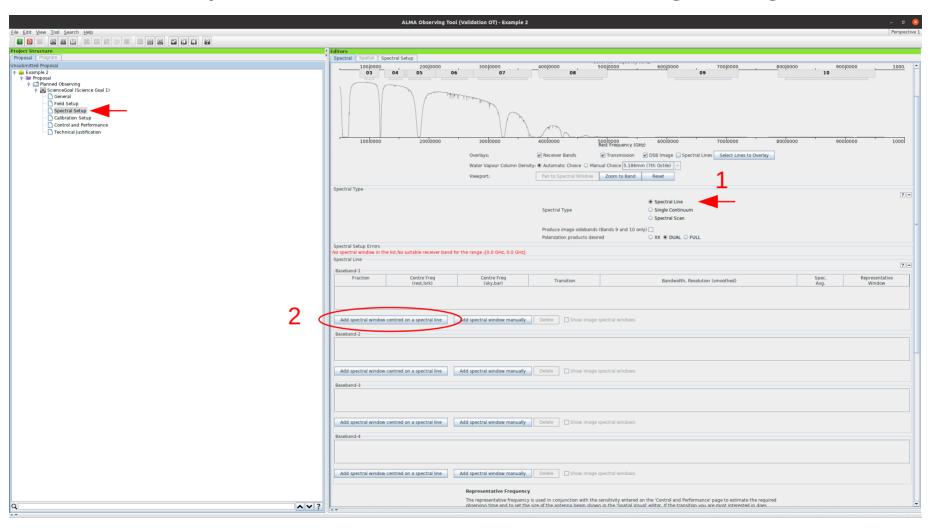
Example 2: Line observations of single target (e.g. study kinematics of a protostellar disk)

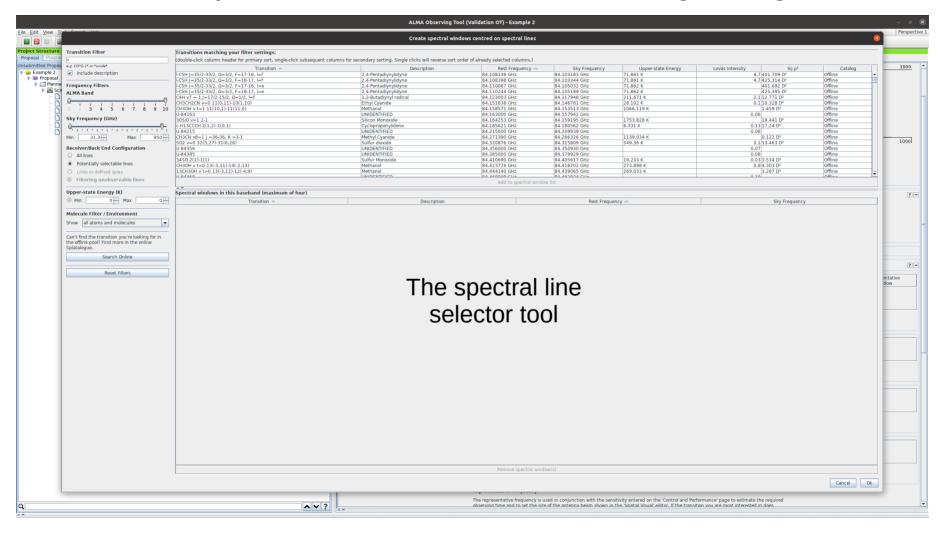
1 Science Goal with 1 target

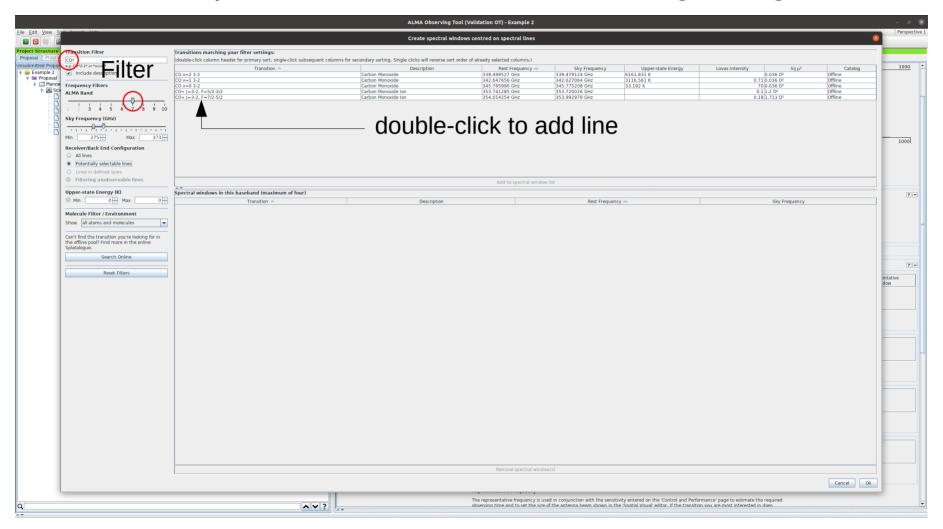
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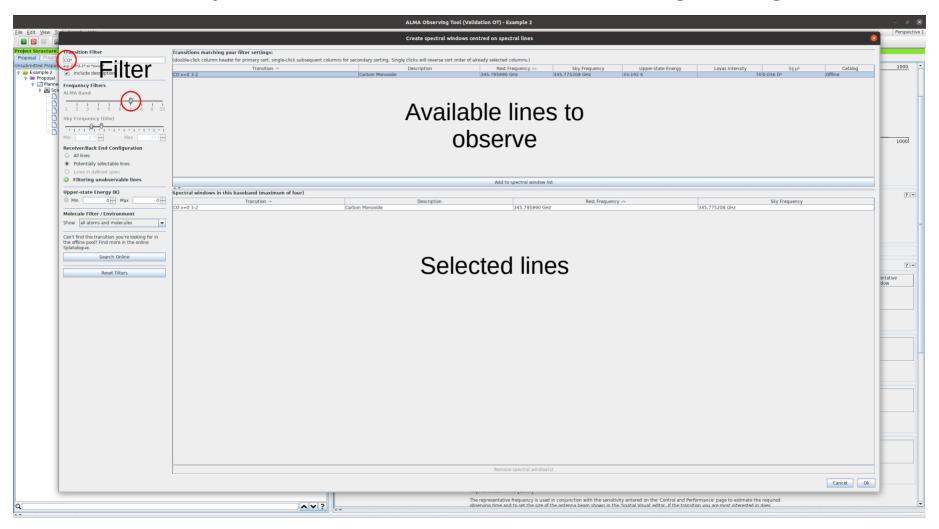
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- 2) Add Science Goal
- 3) Fill in General and Source Setup fields
- 4) Save it as Example 2

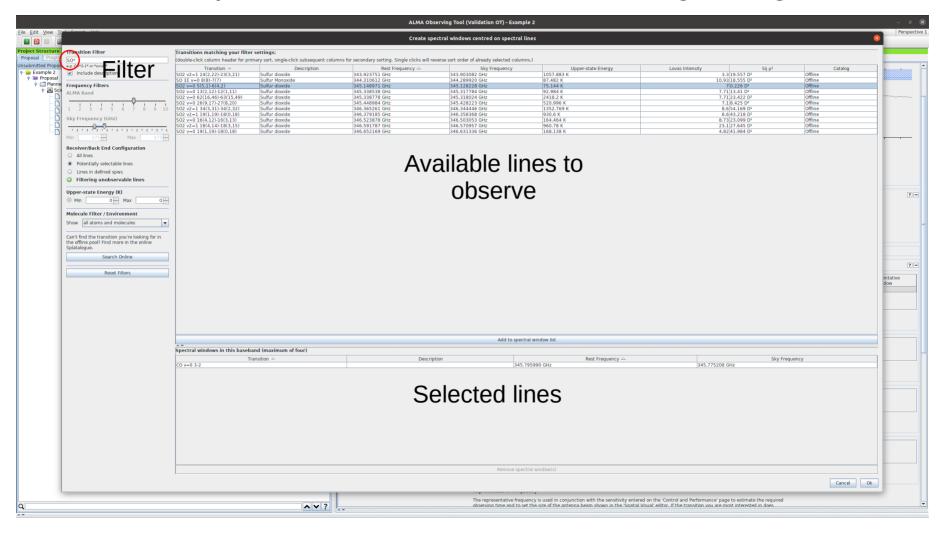


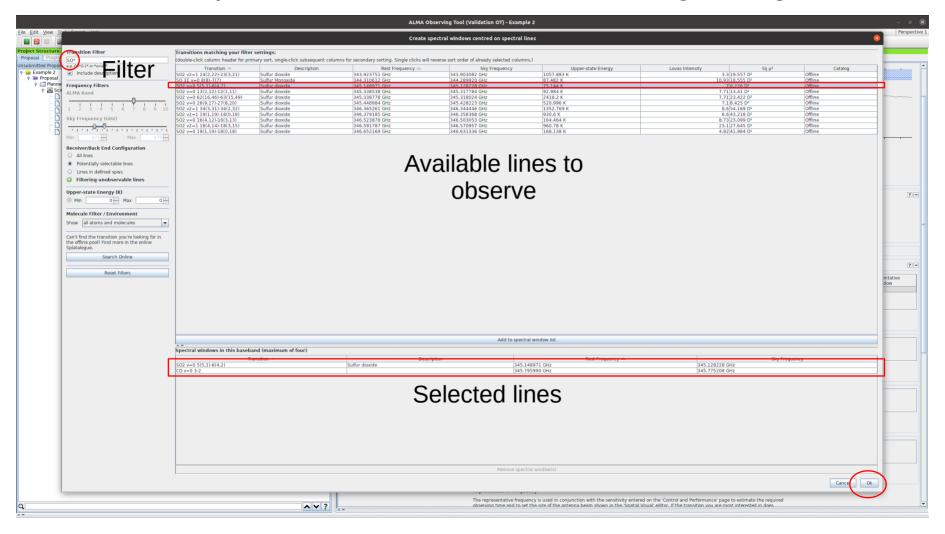


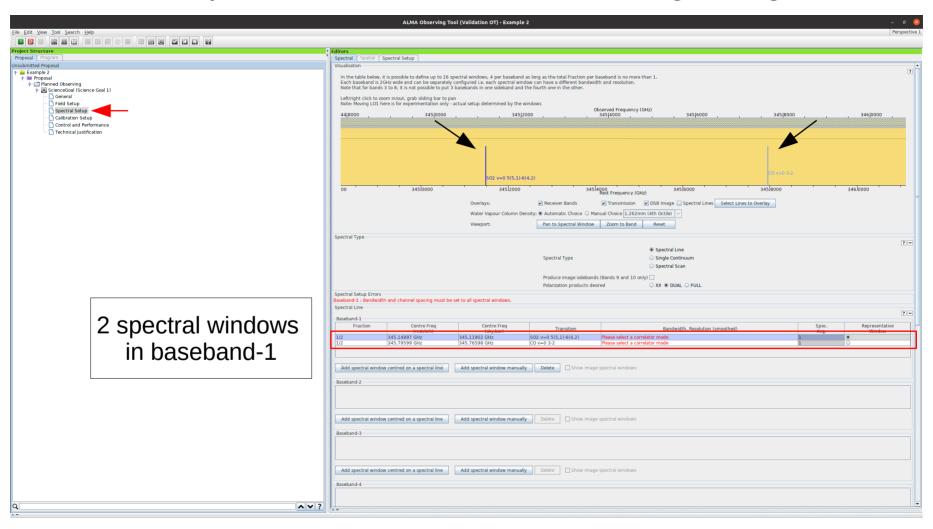


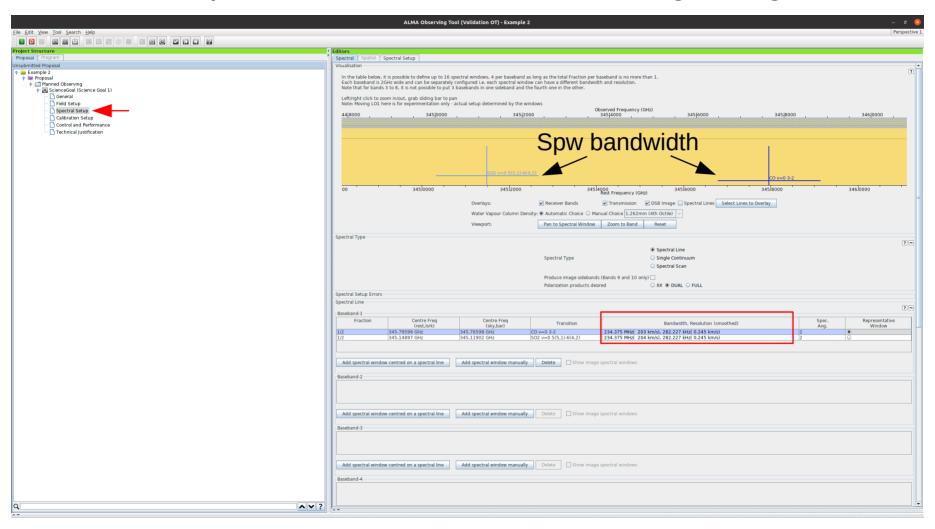


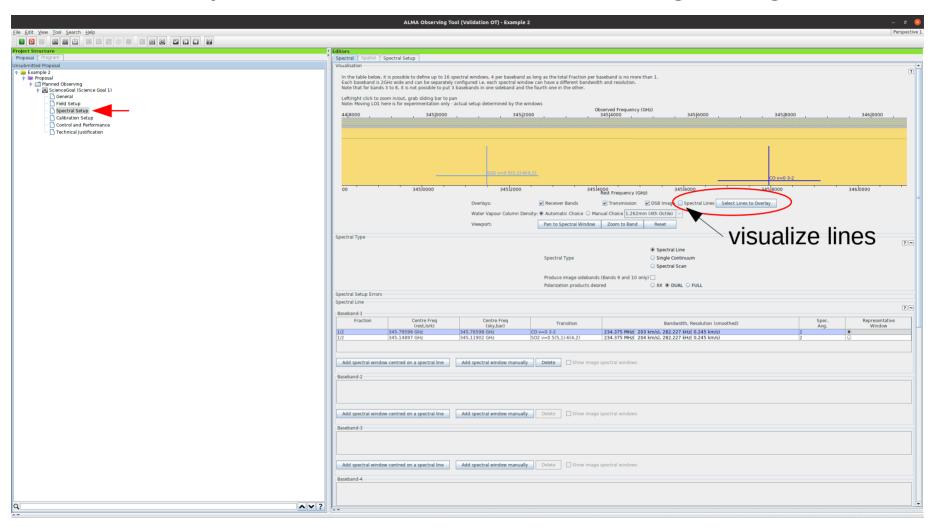


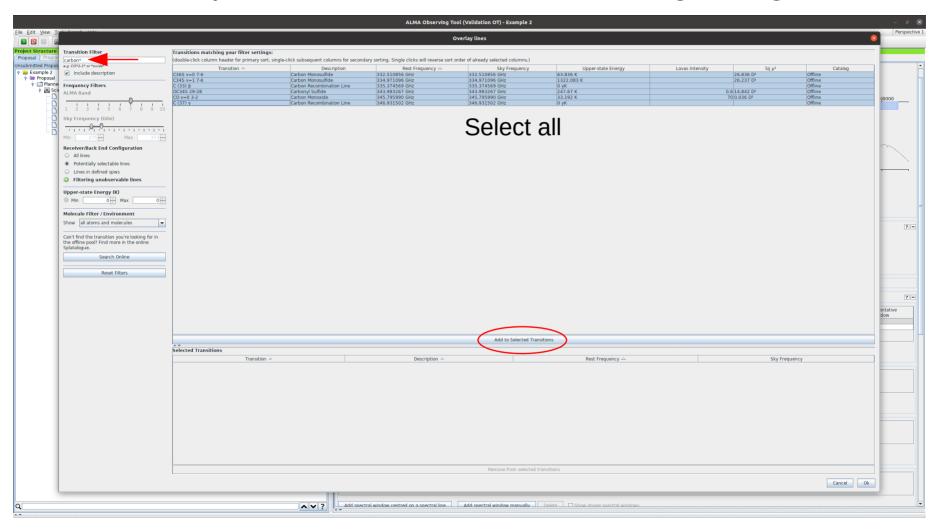


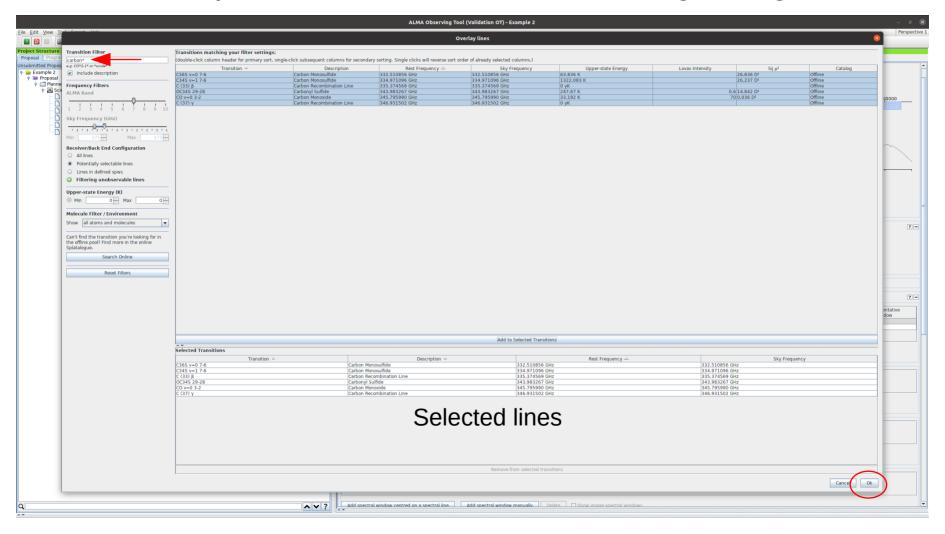


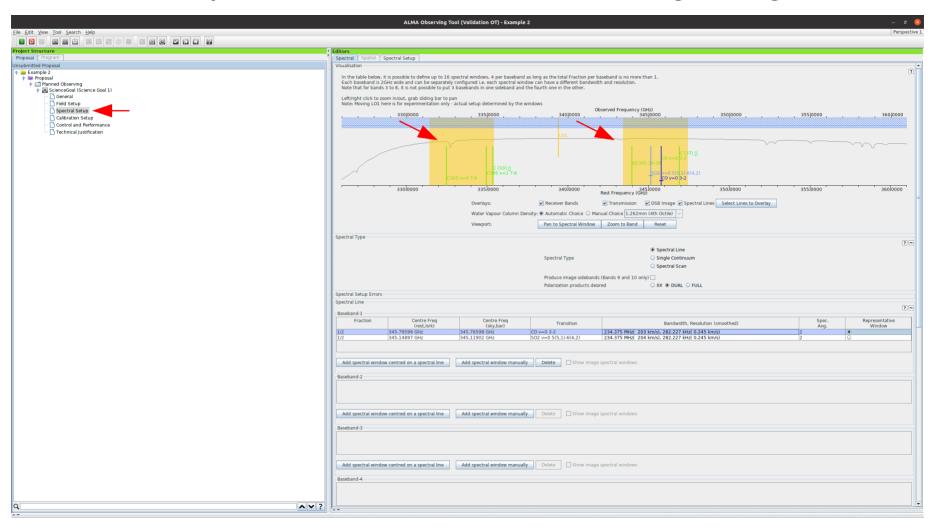


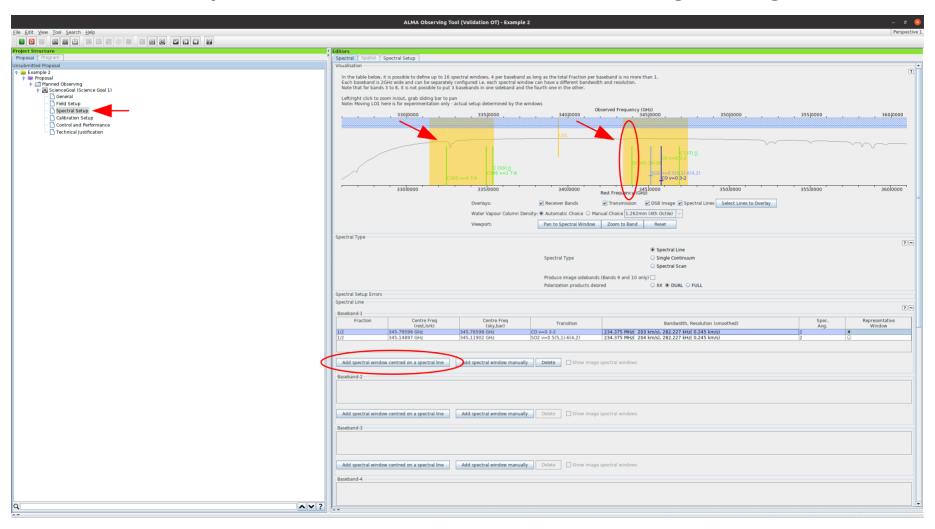


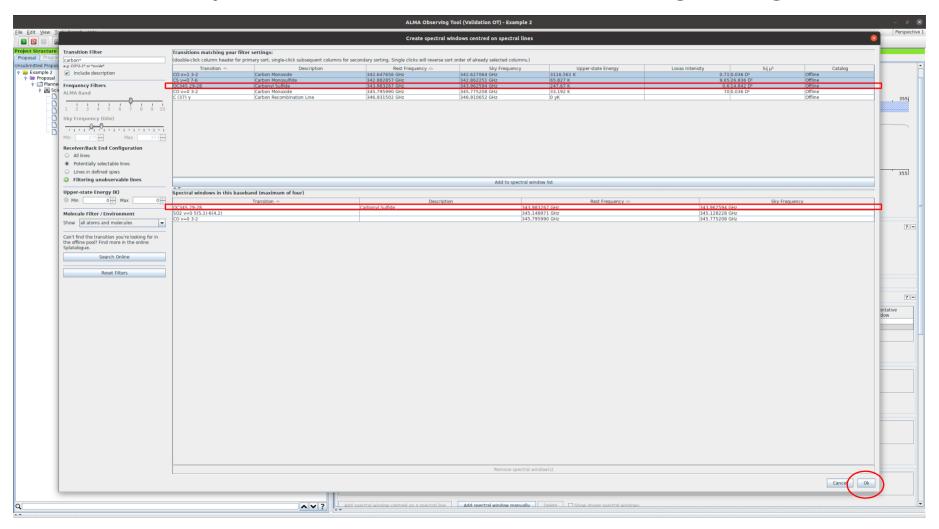


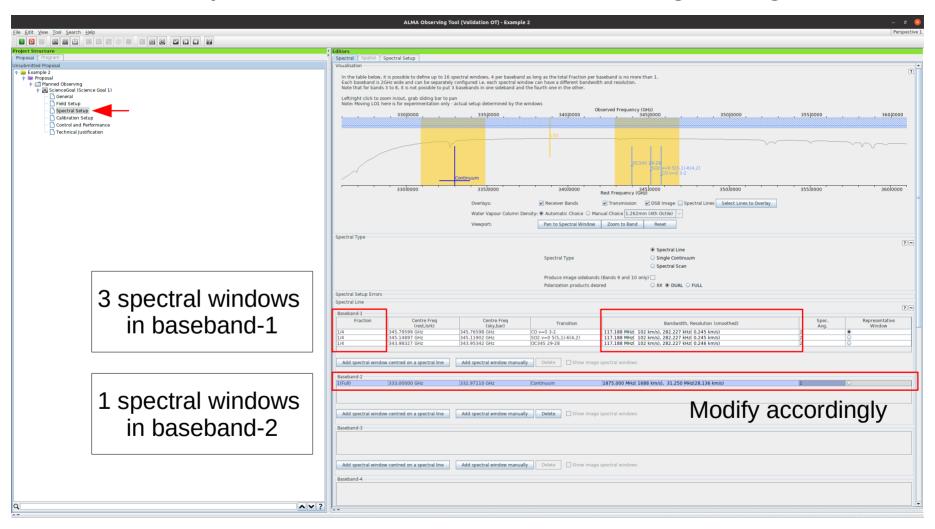


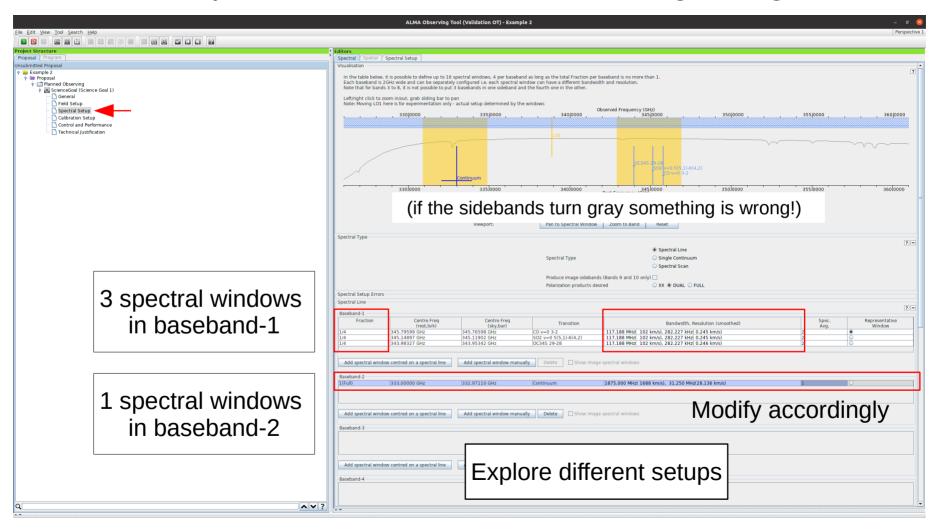


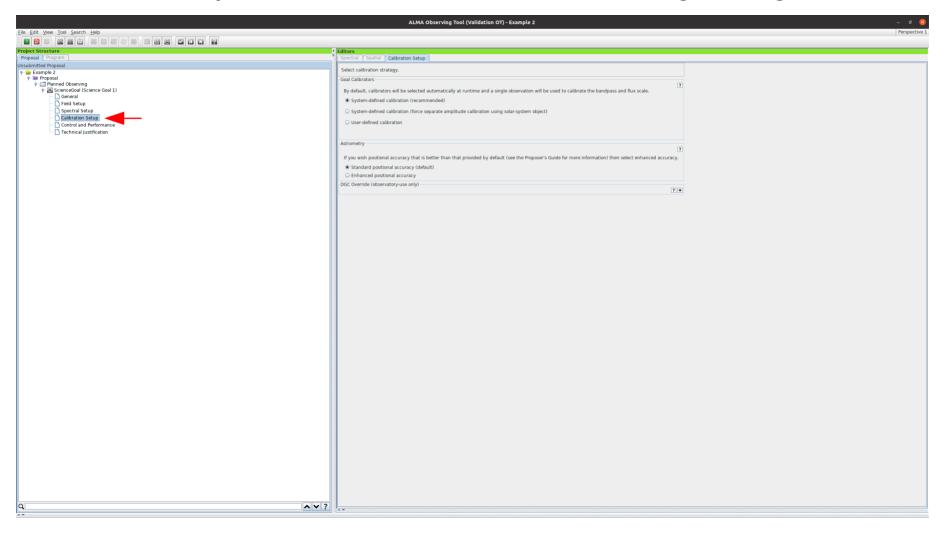


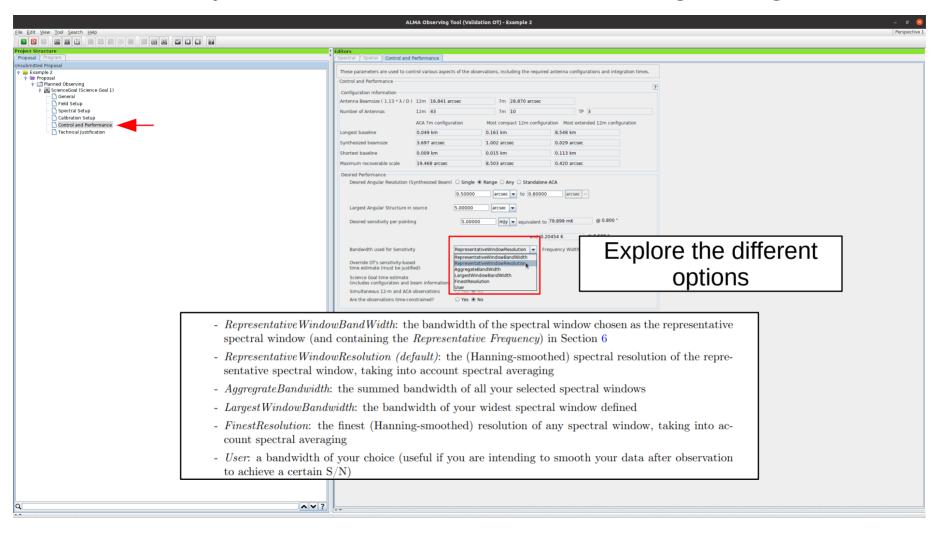


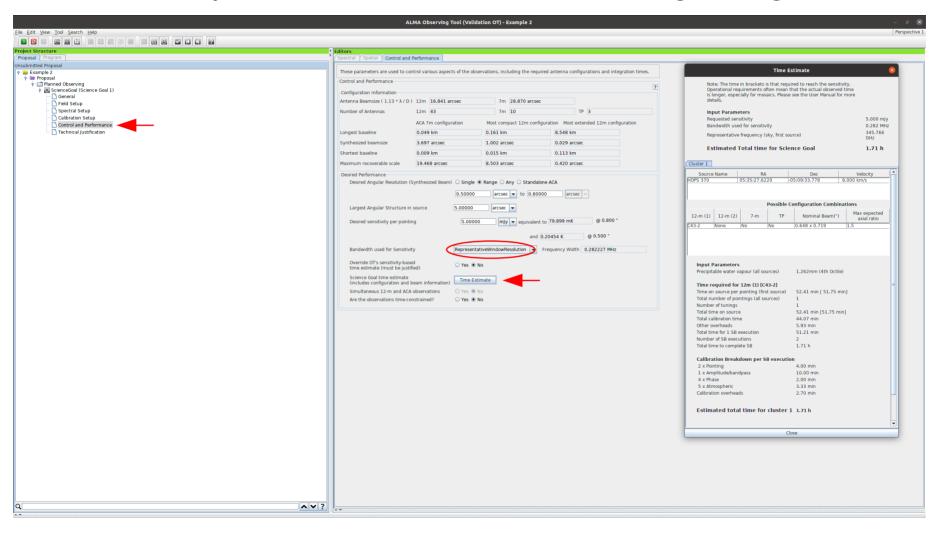


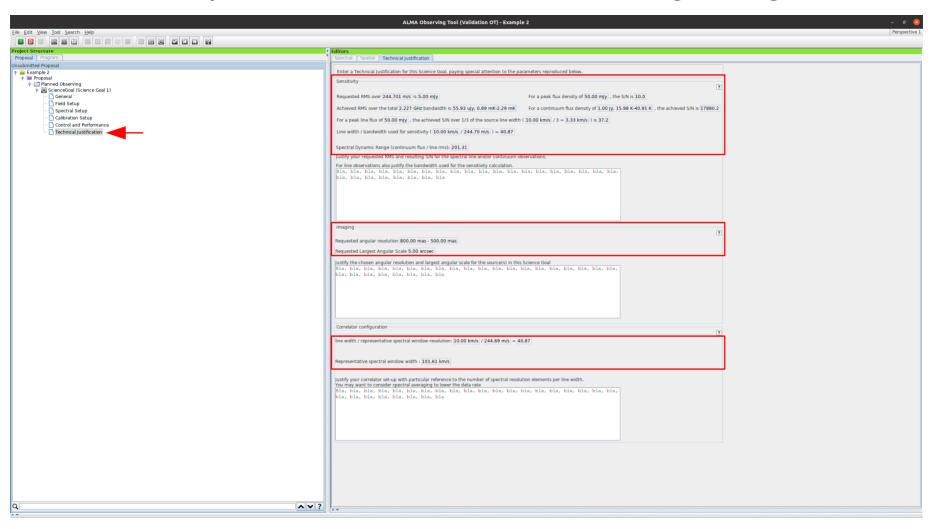


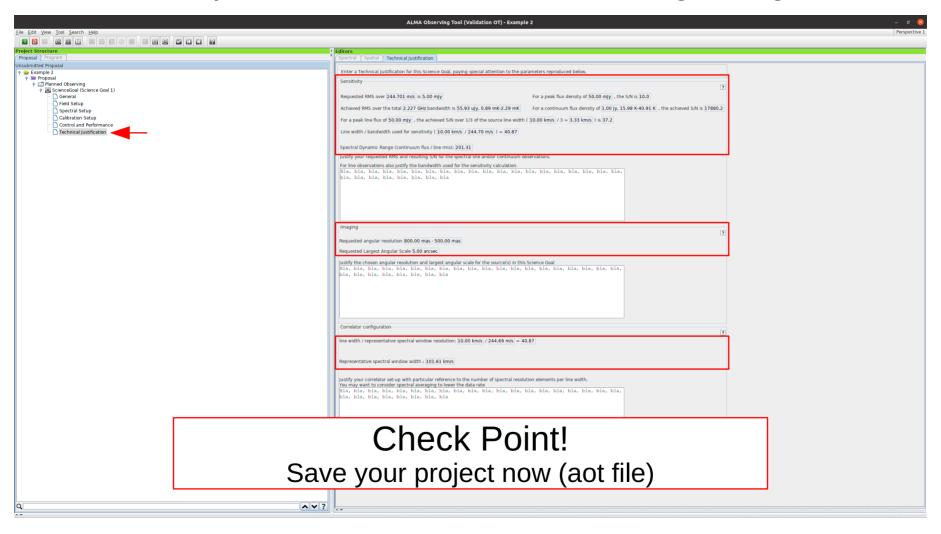












# Lets propose!

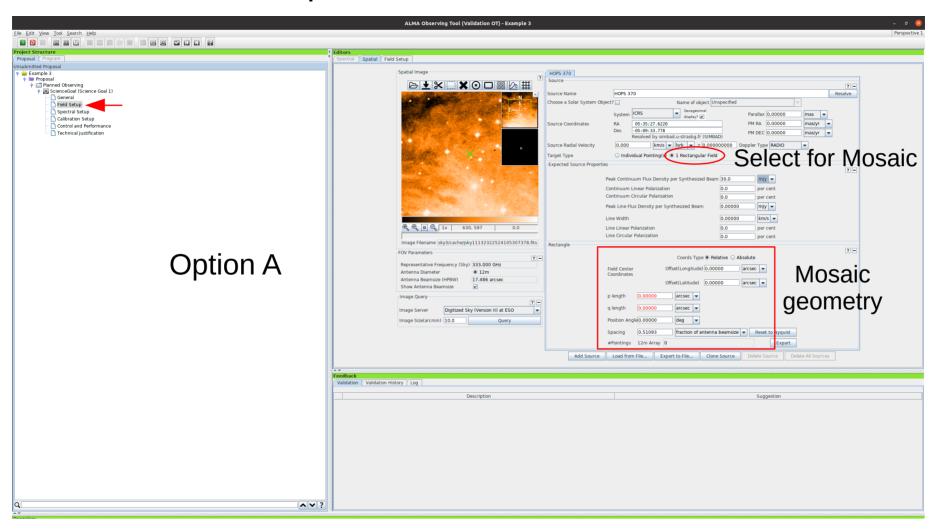
(Phase 1)

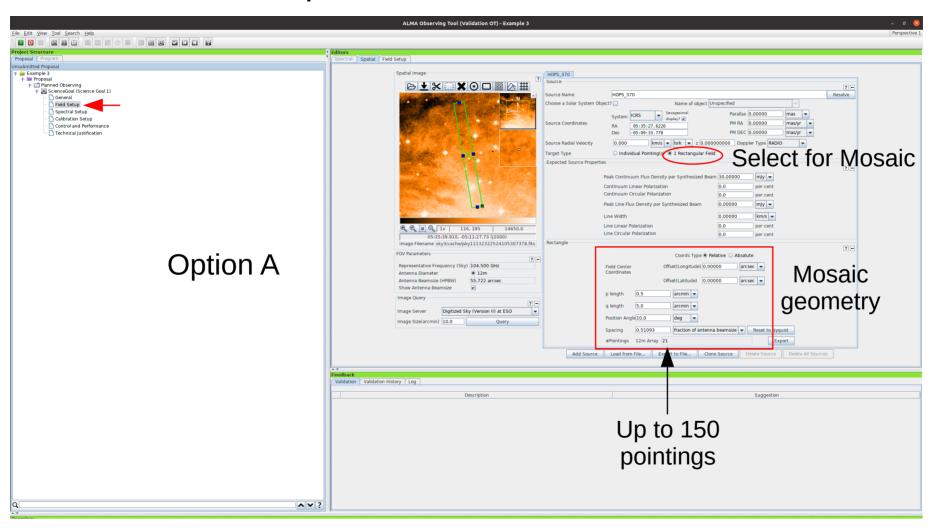
Example 3: Mosaic observations (e.g. mapping a starforming filament)

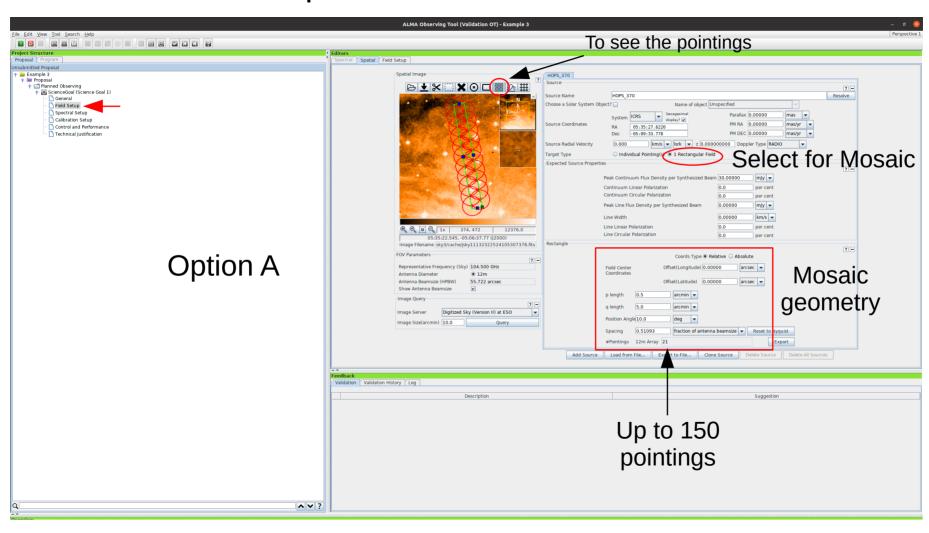
1 Science Goal with 1 mosaic target

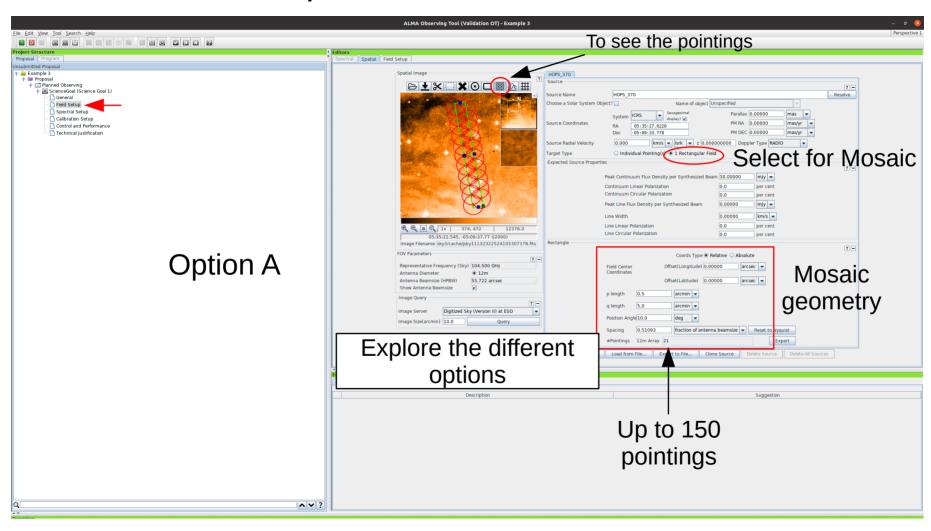
#### To do:

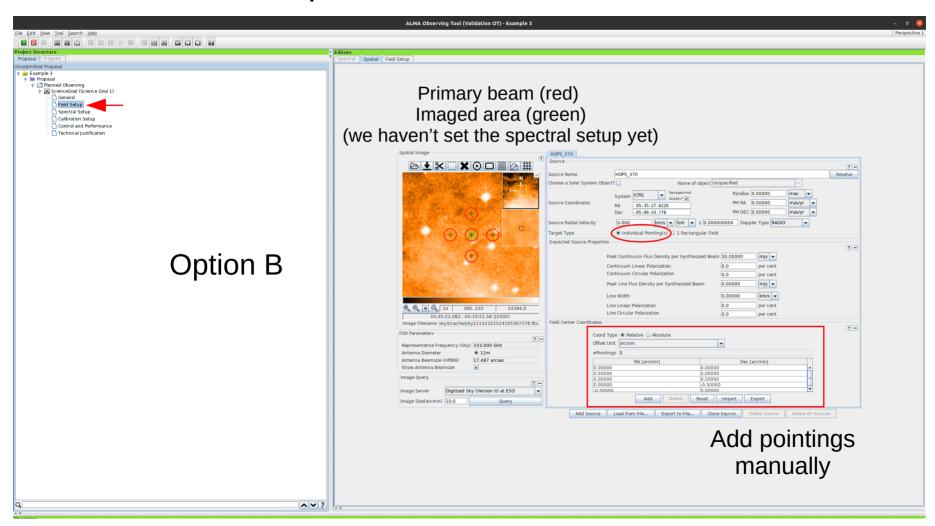
- 1) Open template created before
- 2) Add Science Goal
- 3) Fill in General field
- 4) Save it as Example 3

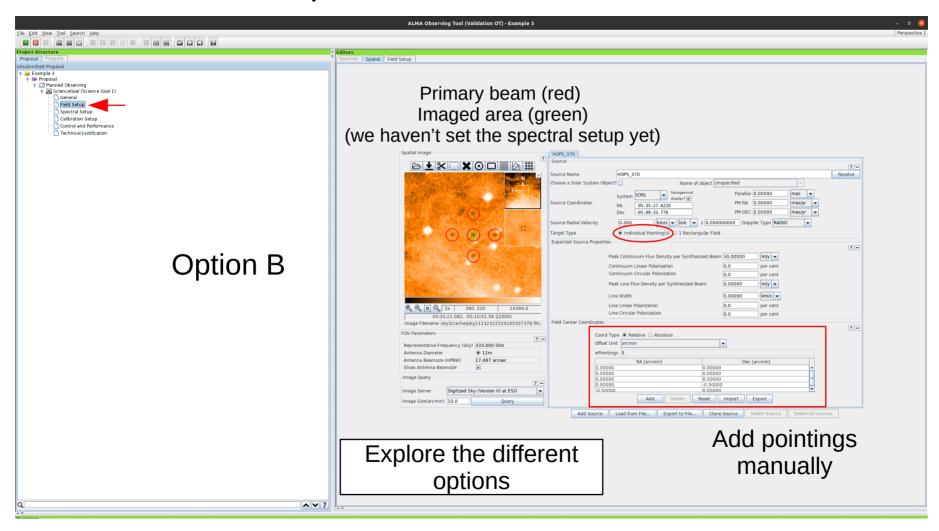


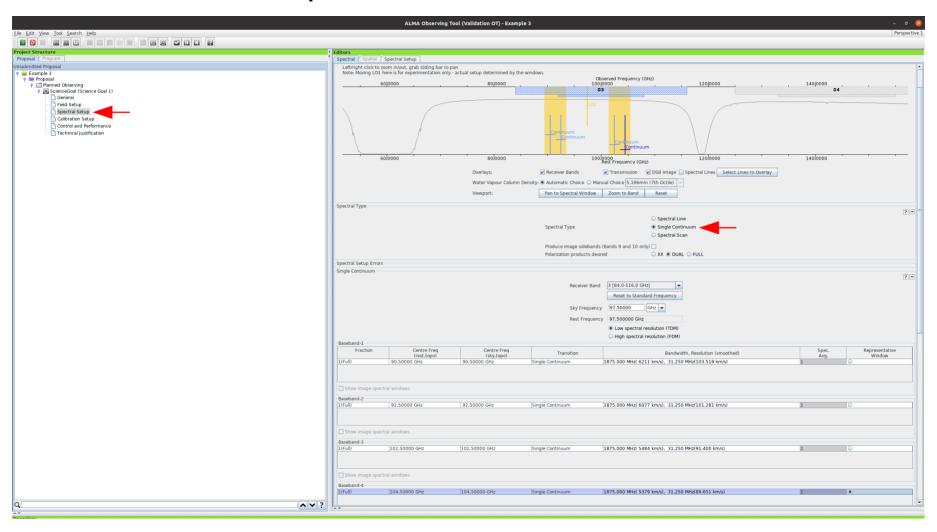


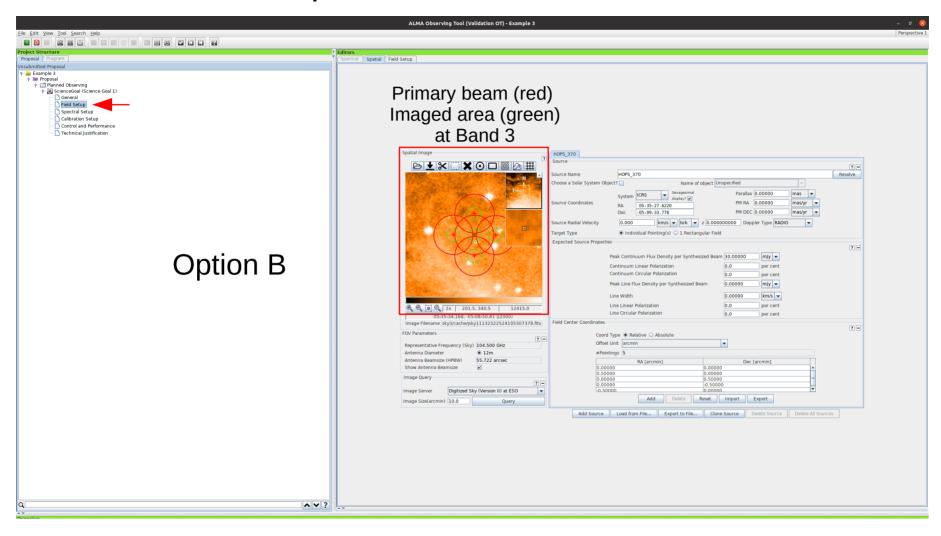


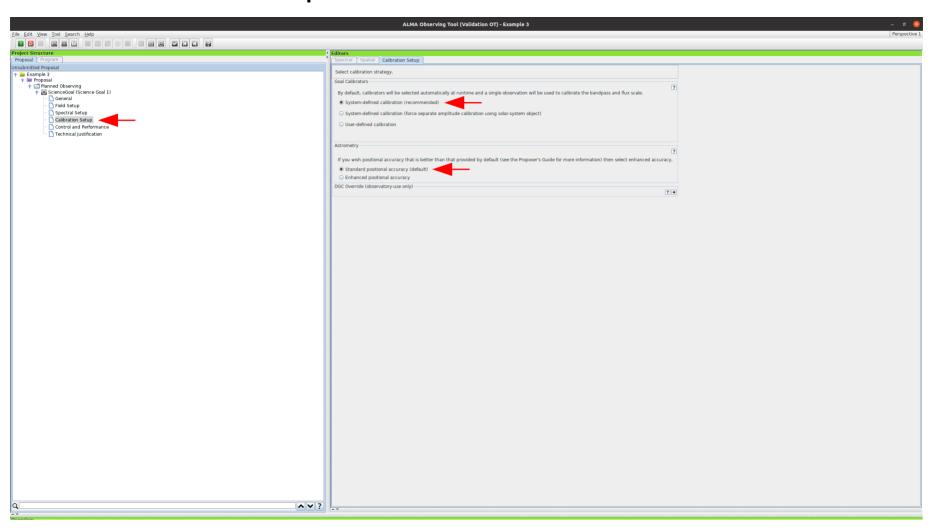


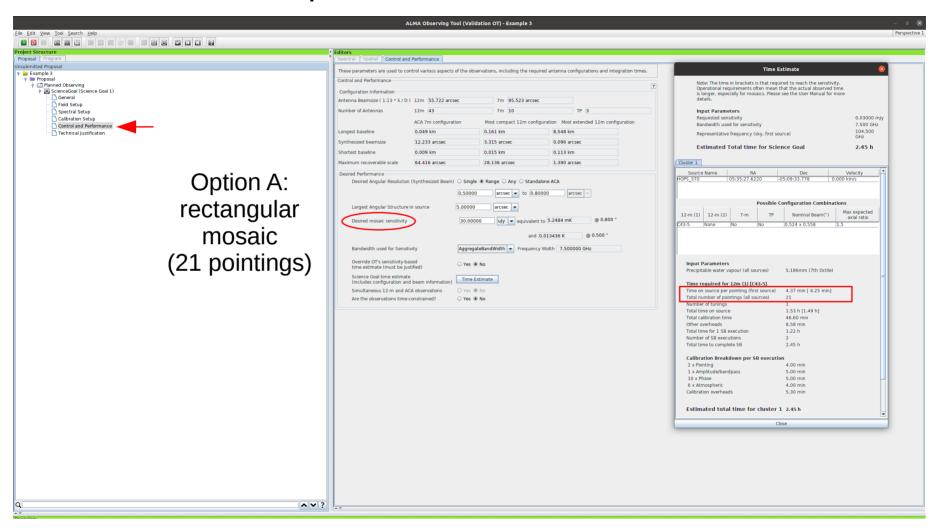


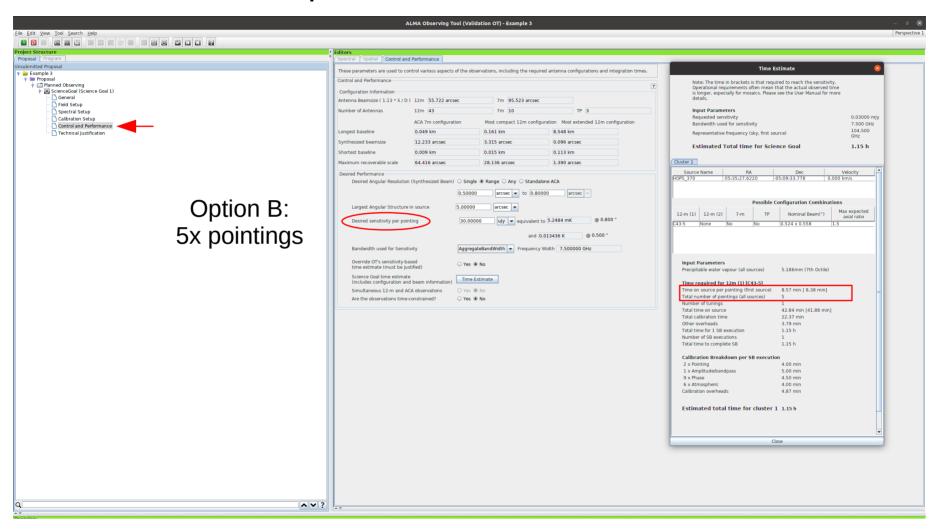


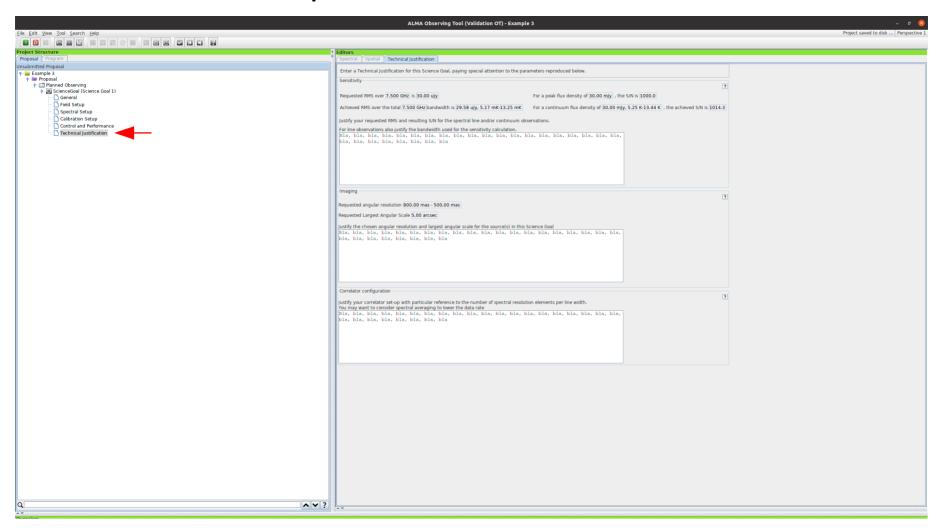


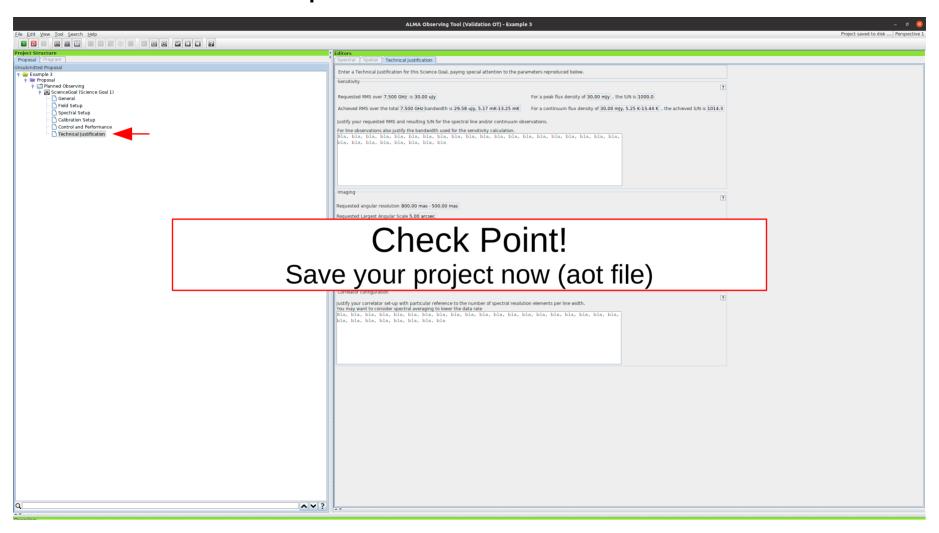












## Congratulations!

Example 1: Continuum observations of several targets (e.g. survey of protostars)

Example 2: Line observations of single target (e.g. study kinematics of a protostellar disk)

Example 3: Mosaic observations (e.g. mapping a starforming filament)

### Final Recommendations

(thanks George, Adam & Anita!)

- Make sure all Co-ls have registered for an account with ALMA (or ESO) so that they can be listed on the proposal.
- Check the source coordinates, velocities and/or redshifts, and spectral settings before proposal submission. These can be updated later, but if more changes need to be made, more errors can be introduced.
- Use at least four spectral windows. Any spectral window not covering a line of scientific interest can be used for serendipitous continuum and spectral line detection.
- Use 1920 channels per baseband. The extra channels provide extra spectral resolution if needed, and if the higher resolution is not needed, the channels can be averaged together after observing to improve sensitivity.
- Do not use 3840 channels per spectral window. The effective spectral resolution will still be equivalent to 1920 channels.
- Do not place important spectral lines near the edges of spectral windows where the sensitivity of the detectors decreases.

- Do not try to gain sensitivity by overlapping the spectral windows. The instrument doesn't work that way.
- Do not change anything under Calibration Setup unless you know what you are doing.
- Do not specify a single angular resolution unless you absolutely need to. A
  program that specifies a range is more likely to be observed.
- Use "Any" for the desired angular resolution if you only need to detect the source.
- Do not forget to account for extended source emission in terms of uv coverage.
- Do not forget to account for extent of the source emission when estimating the peak surface brightness.